

# Data quality assessment of HIRAS-2 on FY-3E



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## Introduction

The Chinese satellite Feng-Yun 3E (FY-3E) was launched in July 2021 and carries the hyperspectral infrared instrument HIRAS-2. Unlike its predecessor, HIRAS, this instrument measures a continuous spectrum from 650-2550 $\text{cm}^{-1}$  with spectral sampling of 0.625 $\text{cm}^{-1}$ .

A unique feature of FY-3E is its orbital plane. The equator crossing time is 05:40, so complements the orbits of existing polar orbiting satellites.

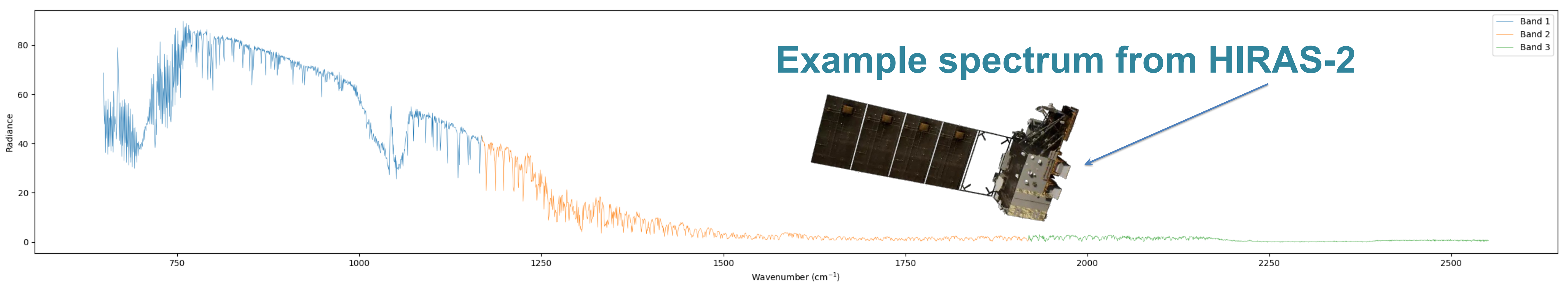
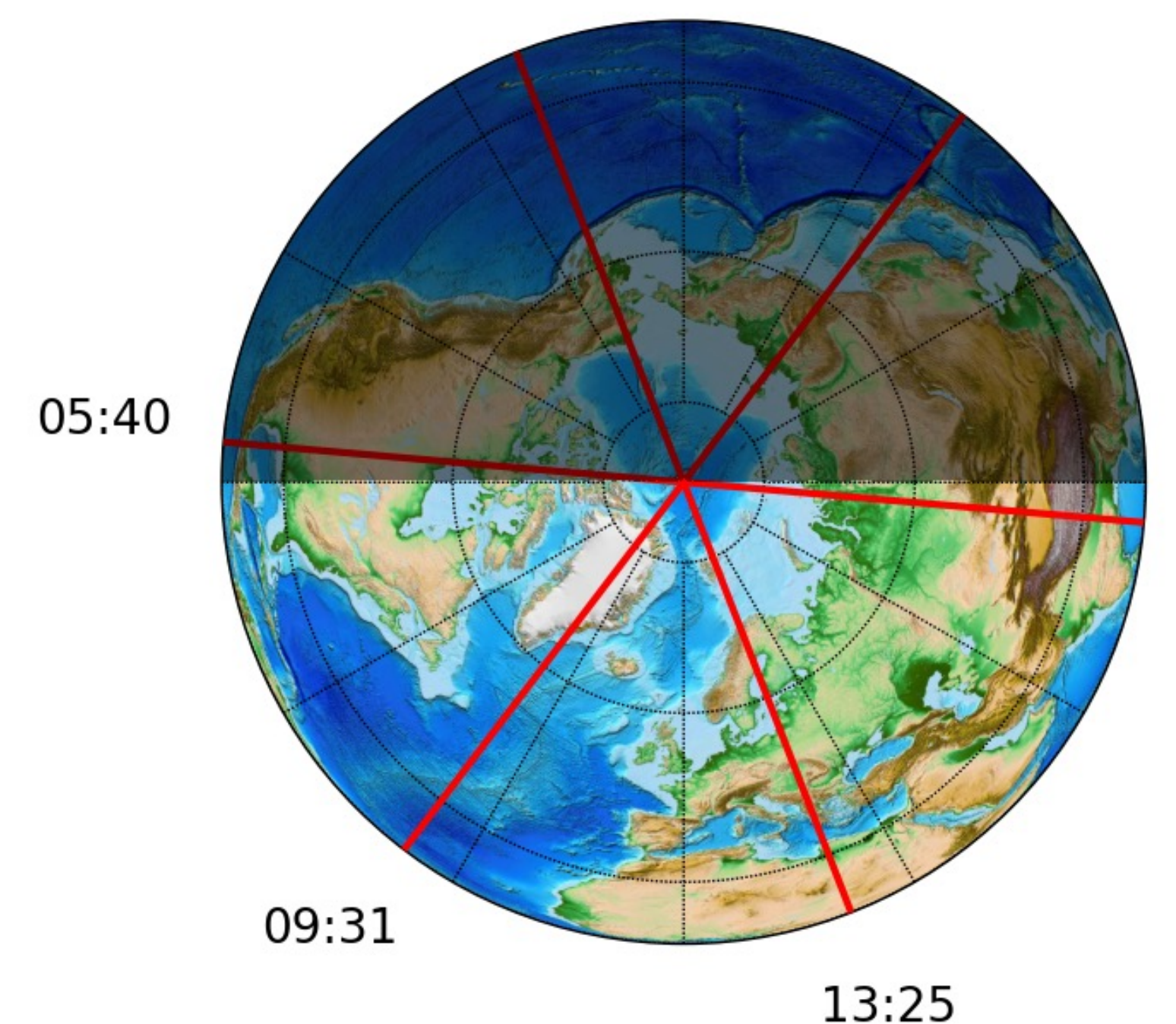
## Orbital plane

In general:

- The European polar orbiters cross the equator at around 9:30.
- The American polar orbiters cross the equator at around 1:30.

With these observations alone, only a small portion of the local times are measured, and hence the full diurnal cycle cannot be constrained by these observations alone in an assimilation system.

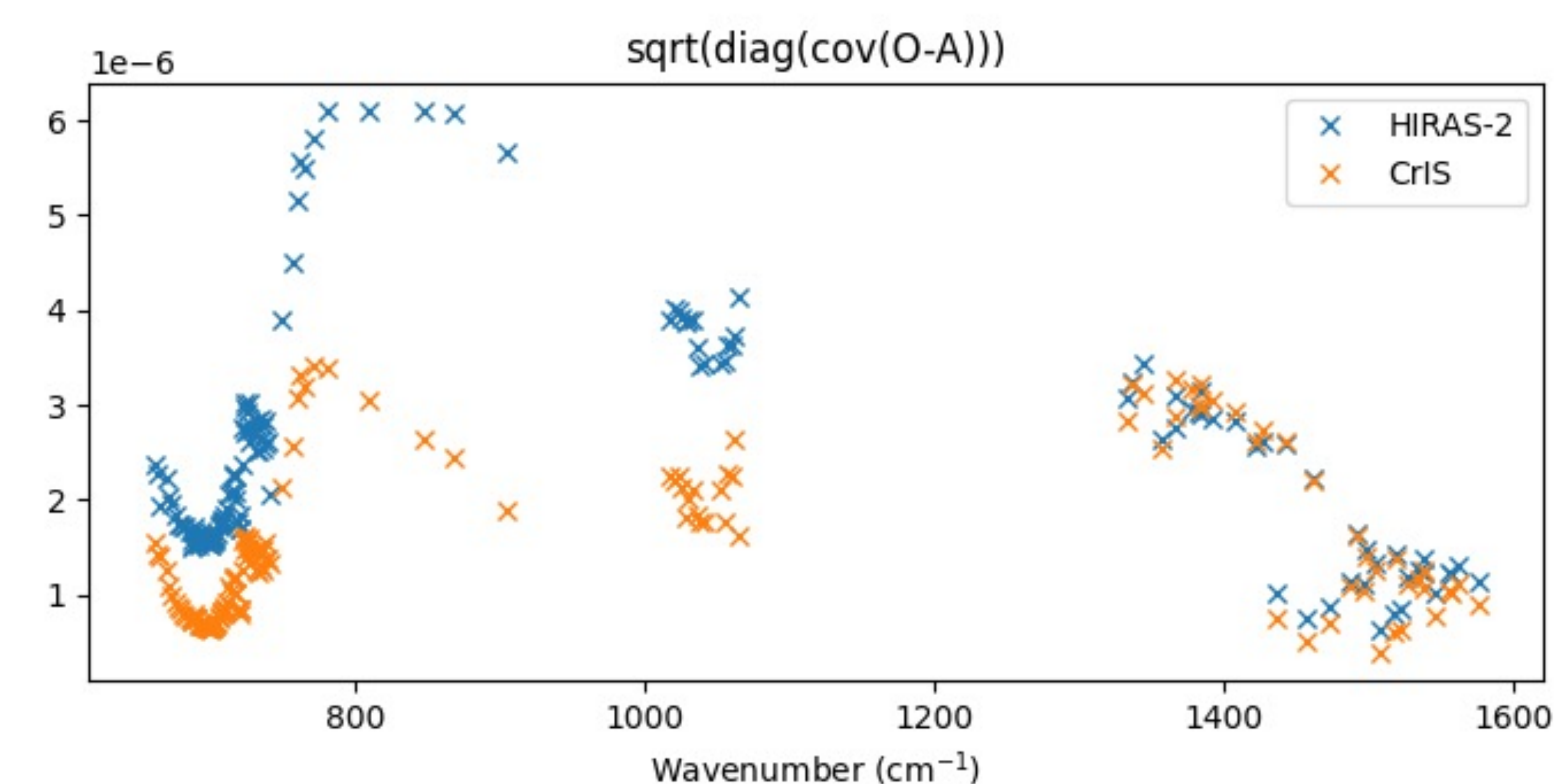
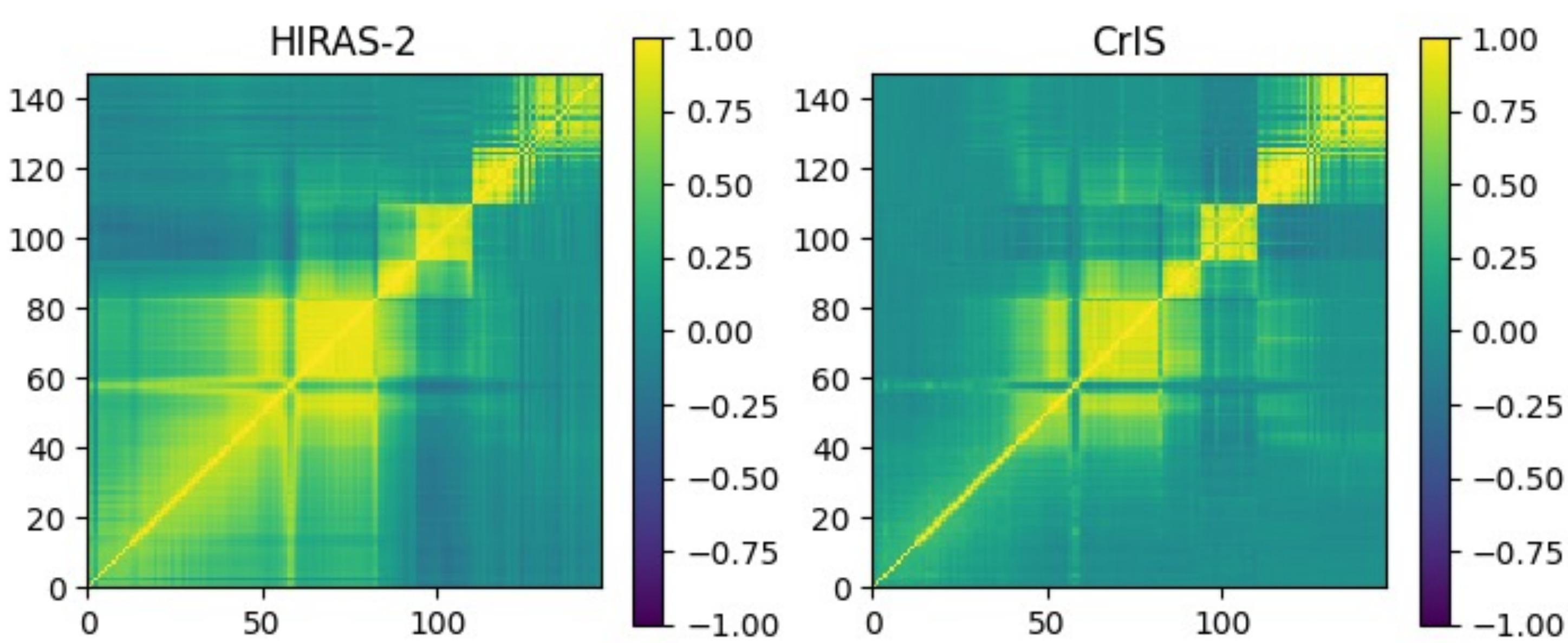
So, FY-3E's orbital plane has potential to be beneficial at improving the diurnal sampling.



## Observation error estimates

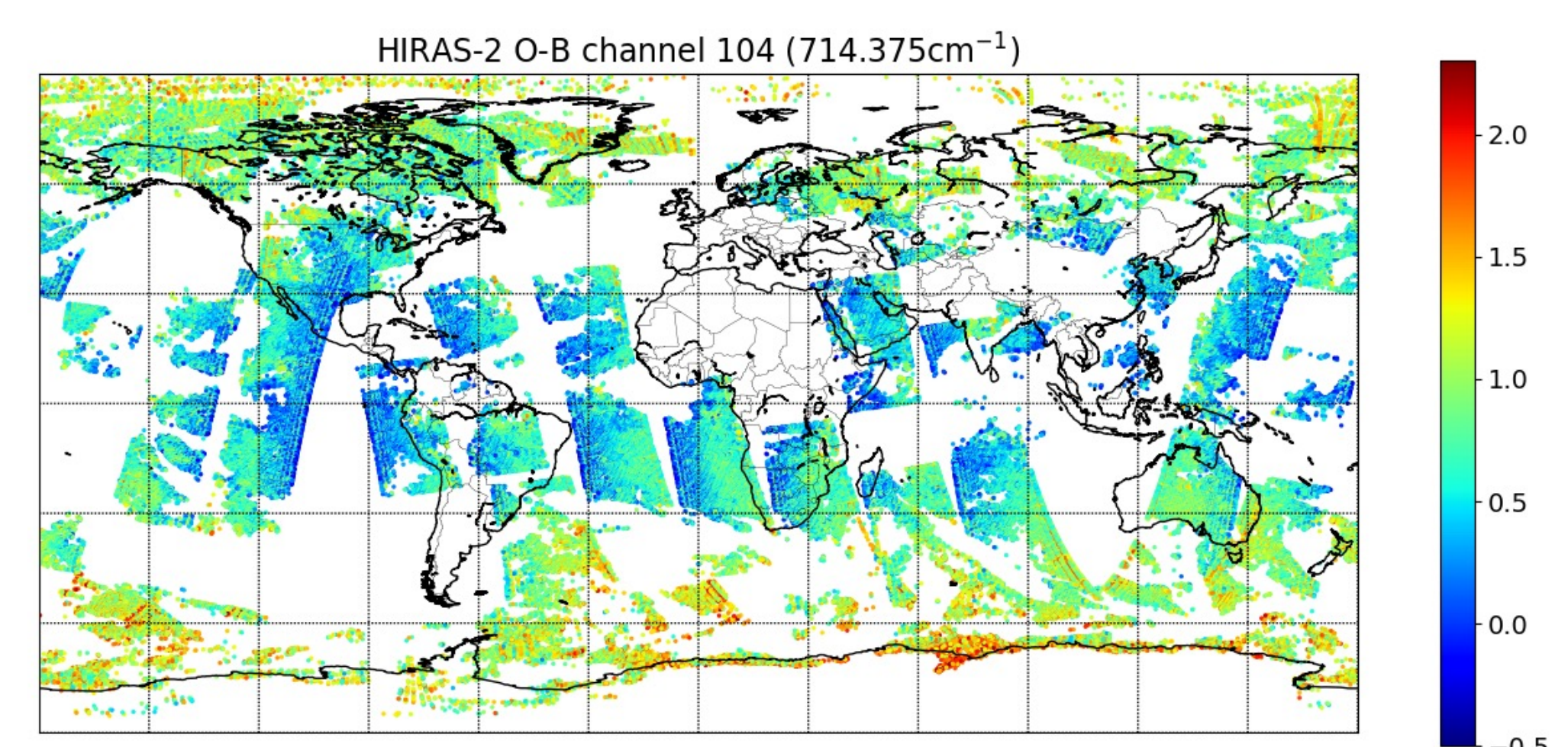
To assimilate radiance data from FY-3E, observation uncertainties need to be estimated. This is done by computing the covariance of O-A (i.e. observations minus simulated equivalents from analyses). This approach makes several approximations, but the simplicity and numerical stability make this an attractive alternative to the Desroziers approach.

Comparing HIRAS-2 with CrIS, we see that the standard deviation of the O-As are larger for HIRAS-2, as are the inter-channel correlations between the used channels.



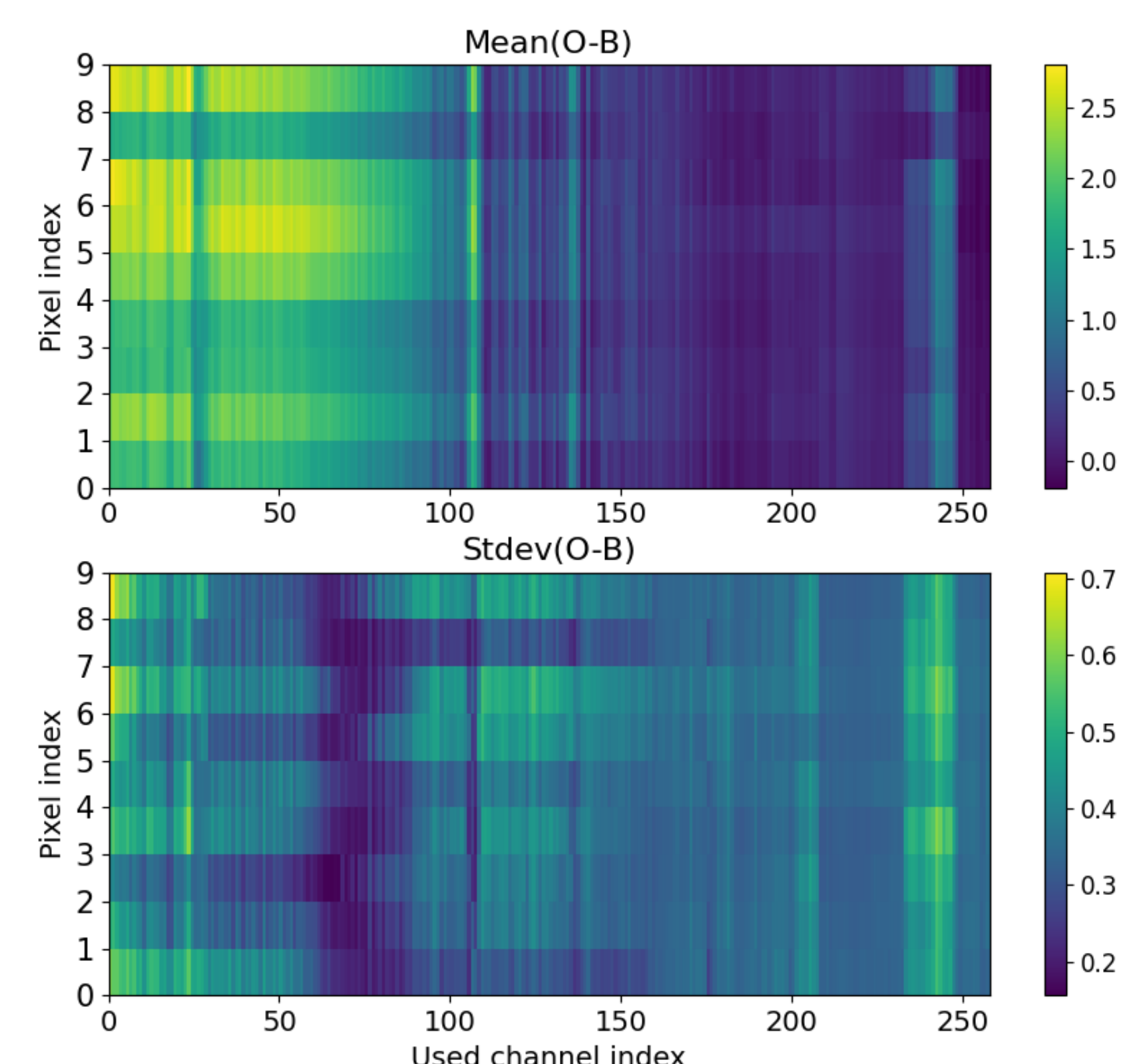
## Cross-scan bias pattern

There is a strong cross-scan bias pattern in some of the LWIR channels. This seems to be worst in the tropics. Identifying suitable bias predictors is not straightforward.



## Variation across pixels after QC

The mean and standard deviation of O-B show significant variation across the pixels. For NWP, we will consider using only one pixel per FoR. Although pixel 2 shows larger noise than others, the sample of observations passing QC and cloud detection are higher than some others, so this has been selected for assimilation experiments.



## Future work

HIRAS-2 will be passively monitored in the ECMWF system soon, along with FY-4B GIIRS. This requires in-house BUFR encoding. Care will be needed due to the large increase in the total number of observations in the ECMWF system. Assimilation experiments are in progress.

## Acknowledgements

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