

IASI-NG : OVERVIEW

L1 Processing Performances

QUENTIN CEBE . IASI-NG SYSTEM
PERFORMANCES MANAGER

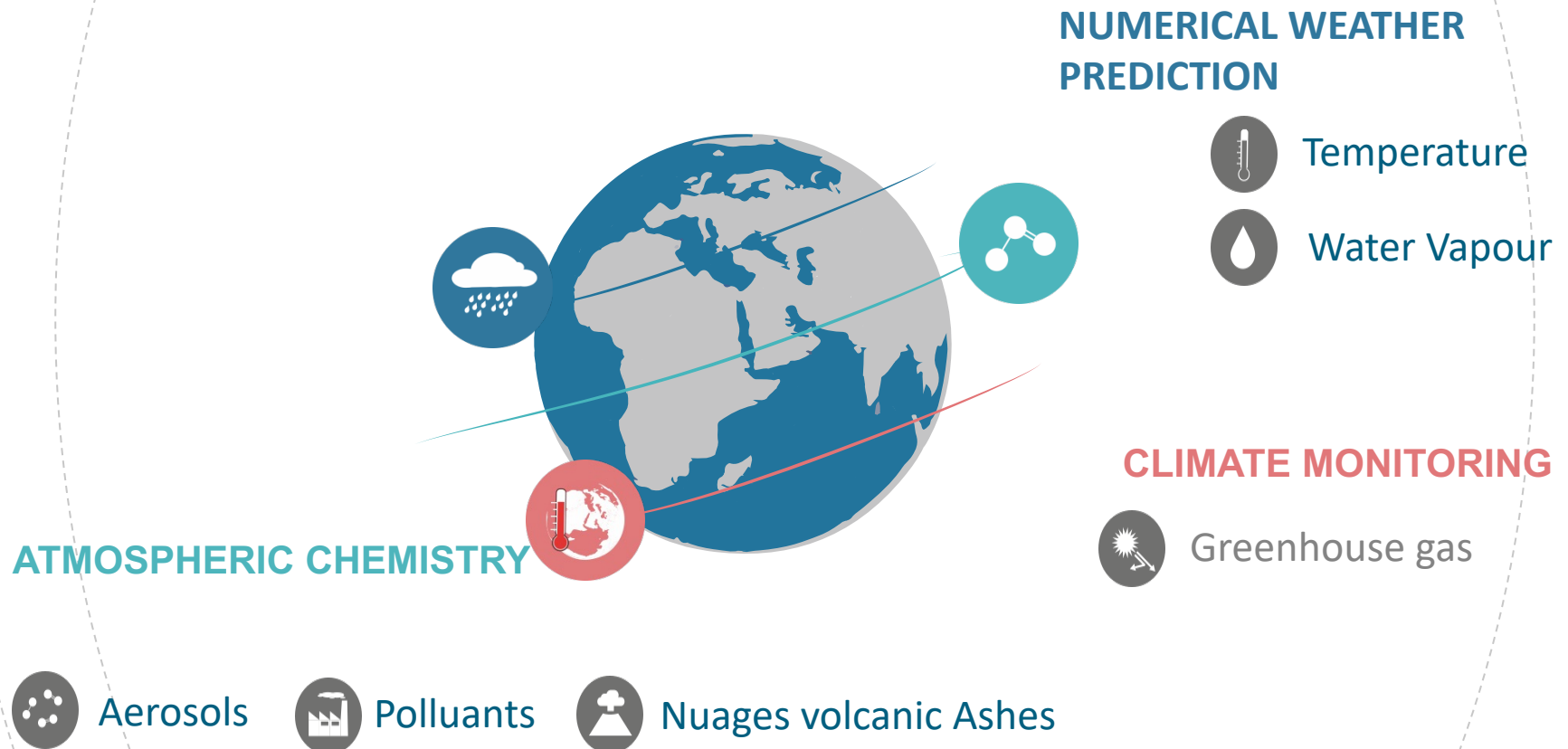
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ITSC 2025

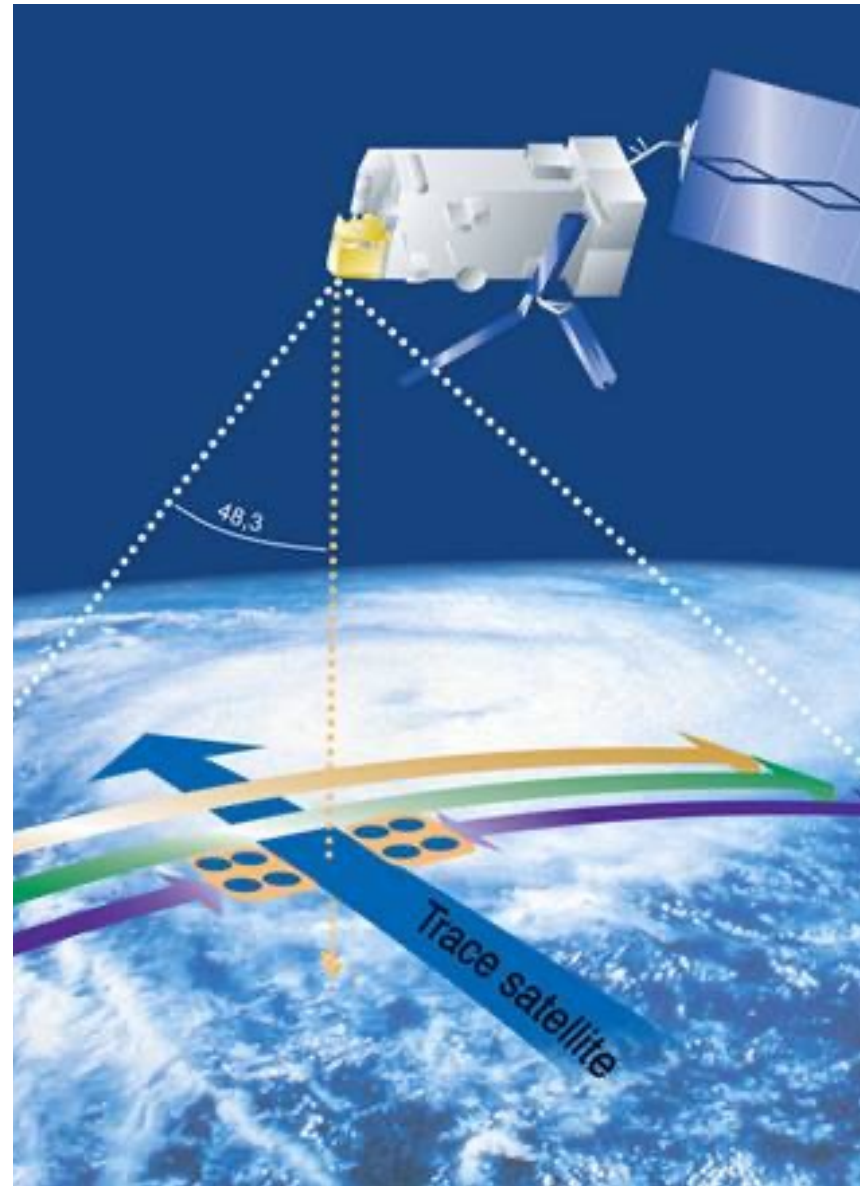
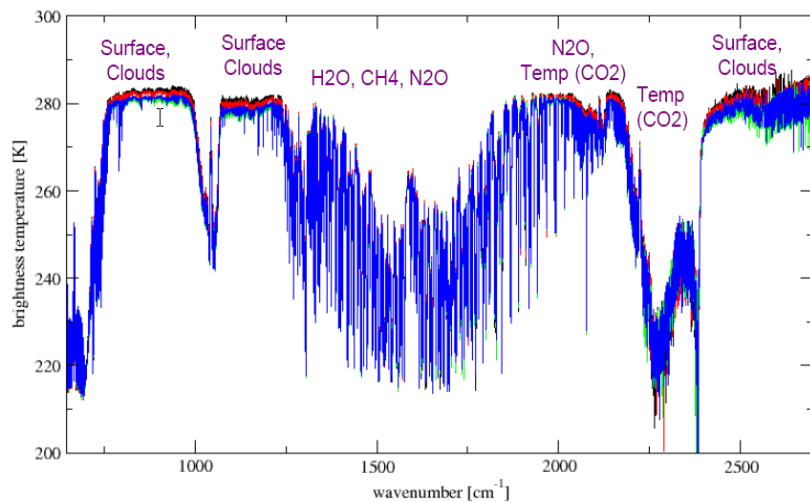
May 08 > 14 2025
Goa, INDIA

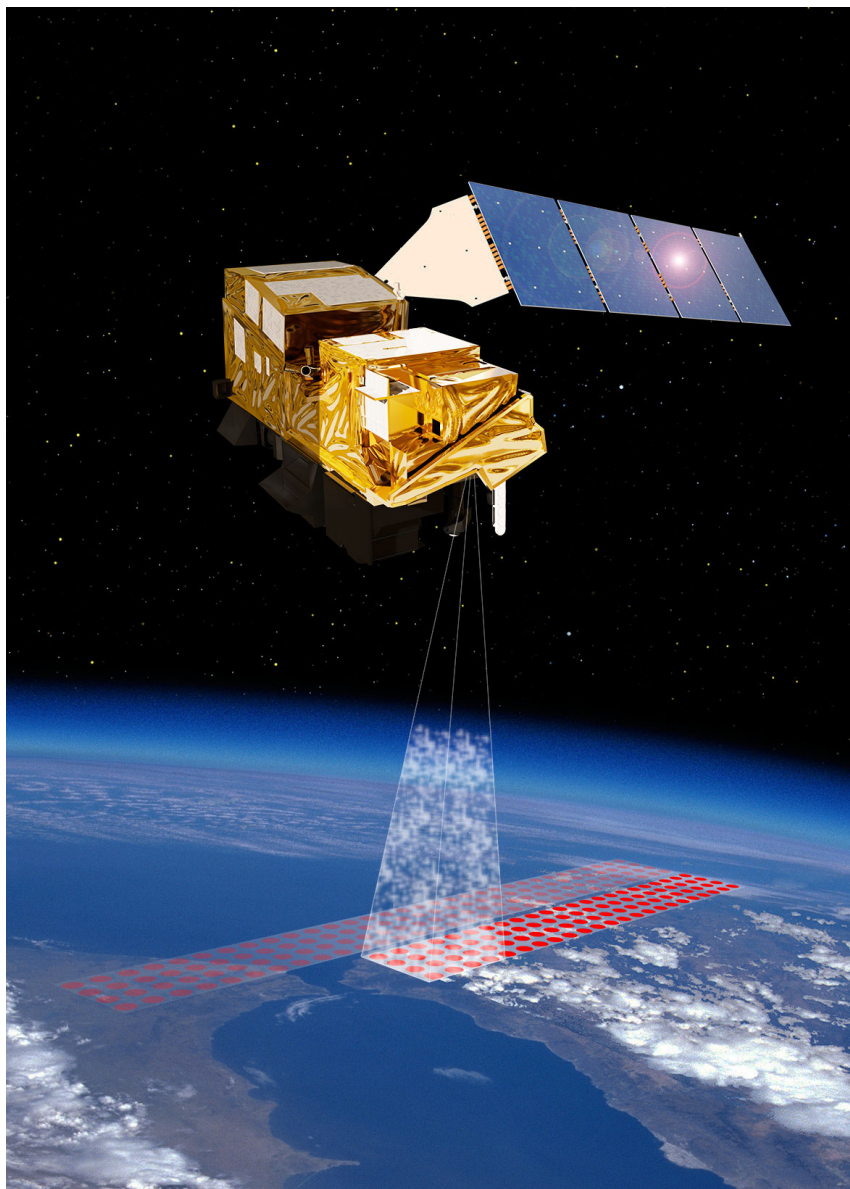
IASI & IASI-NG MISSIONS



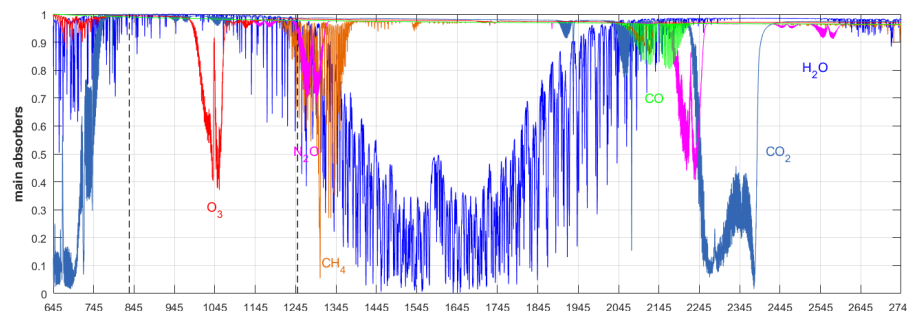
IASI

- 3 instruments aboard 3 satellites
 - MetOp A/B/C – 2006/2012/2018
 - 18 years of mission and counting
 - MetOp-A decommissioned in 2021
 - Continuous spectrum of 8461 channels





- 3 instruments aboard 3 satellites
 - MetOp-SG A 1/2/3 – 2025/2032/2039
 - Same number of products per day
 - Same pixel sizes
 - Same spatial sampling
 - Same wavelength coverage
 - $645 - 2760 \text{ cm}^{-1}$ / $3.6 - 15.5 \text{ }\mu\text{m}$



MAJOR REQUIREMENTS

- **Afocal and Imaging Telescopes**
- **Mertz Interferometer**
- **Focal Plane** : 4 detectors for 4 spectral bands /16 sounding pixels per detector
- **5 metrology lasers:**
 - 1 central metrology to give the Optical Path Difference constant triggering
 - 4 lateral lasers to monitor in real time the pupil effects (tilt, focus and astigmatism) for correction through on ground processing

SPECIFICATIONS

GEOMETRY	SOUNDER PIXEL SIZE	~12 km
	SPATIAL SAMPLING	~25 km
	GEOLOCATION ERROR	0.5 km
SPECTRAL	BAND	645 cm ⁻¹ to 2760 cm ⁻¹
	RESOLUTION	0,25 cm ⁻¹
	SAMPLING	0,12 cm ⁻¹
	CALIBRATION ERROR	$d\sigma/\sigma = 5.10^{-7}$
RADIOMETRY	CALIBRATION ERROR	0,25K @ 280 K
	NEDT	NedT ~0.1 K to 0.4 K within spectrum

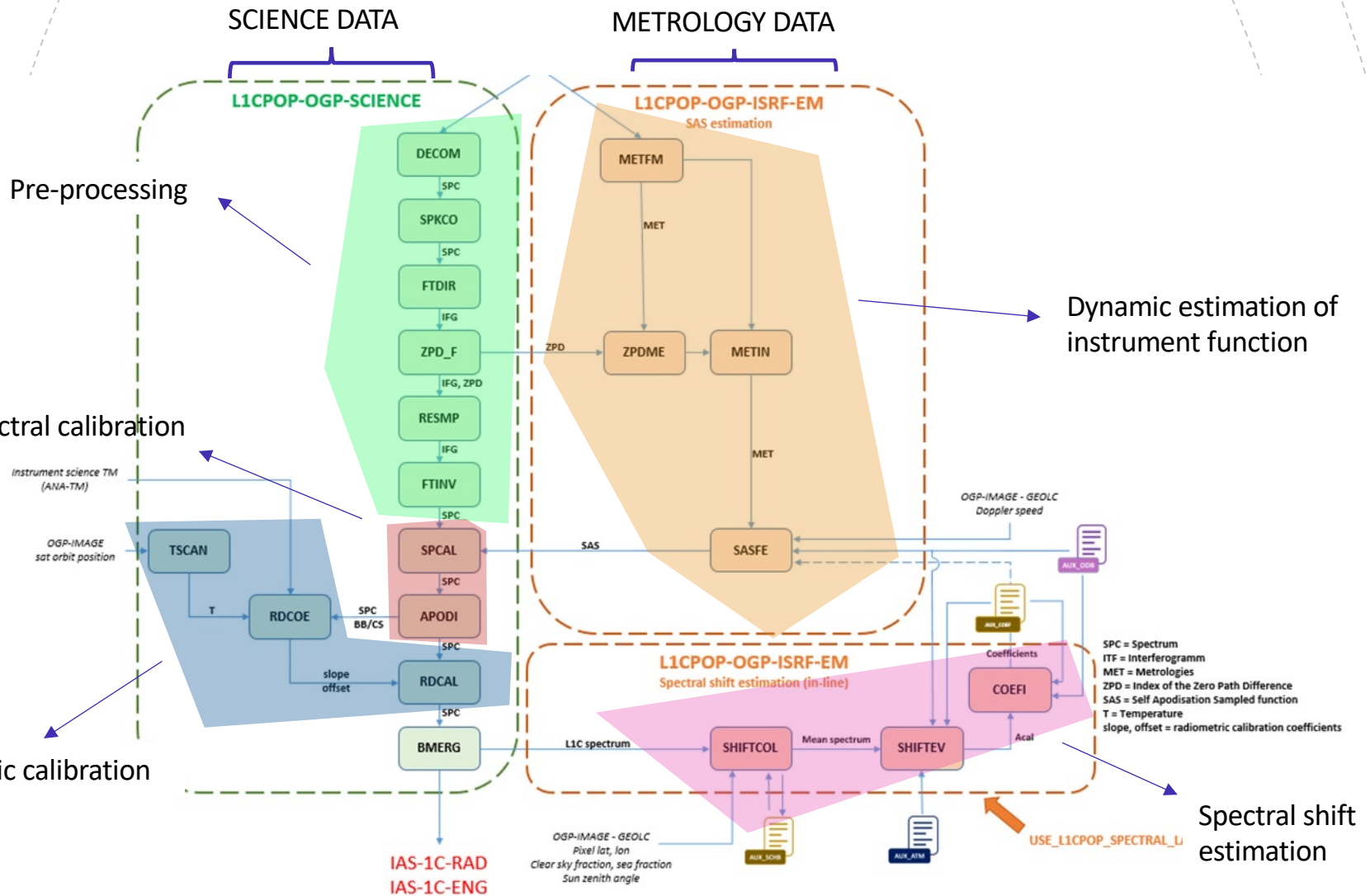
SAME AS IASI

2 TIMES BETTER THAN IASI

MAIN CHARACTERISTICS

SWATH	~ 2000 KM
FOR	+/- 3°
PUPIL DIAMETER	~ 90 MM
ATA MAGNIFICATION	2.3
MAXIMUM OPTICAL PATH DIFFERENCE	4,2 CM
ACQUISITION DURATION	~730 MS
SCAN LINE DURATION	15.6 s : 14 EARTH VIEWS + 1BB + 1CS
CO-REGISTRATION	INTEGRATED IMAGER
SPECTRAL CALIBRATION	FABRY PEROT SOURCE

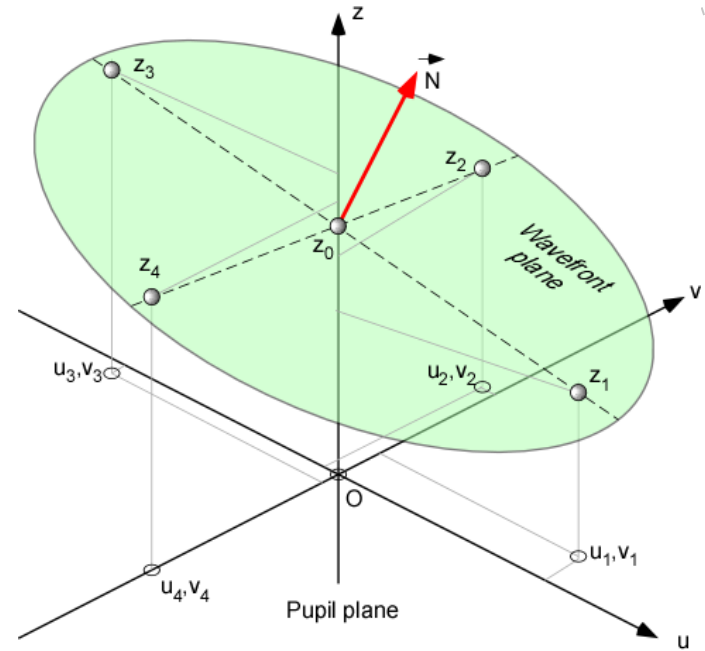
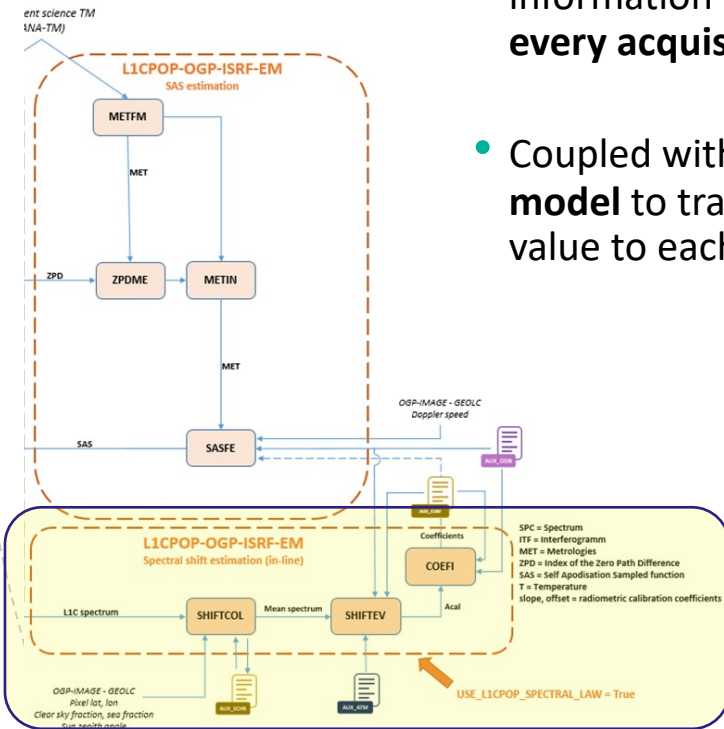
L1 SOUNDER PROCESSING



SPECTRAL CALIBRATION

ISRF estimation is a mix of measurement and modelisation

- **5 lasers** : dynamic information during **each and every acquisition**
- Coupled with an **instrument model** to translate the ISRF value to each wavenumber



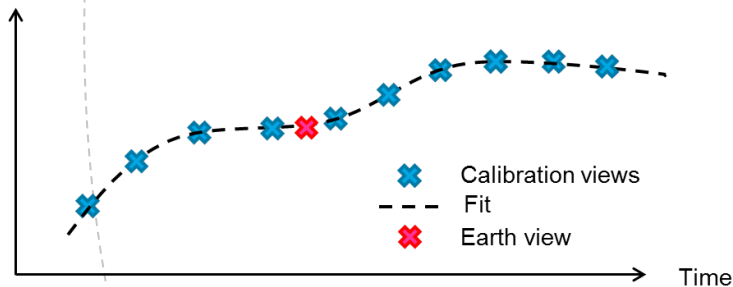
Re-estimation of the instrument coefficients using calibrated data on-flight

Asynchronous with the rest of the processing

RADIOMETRIC CALIBRATION

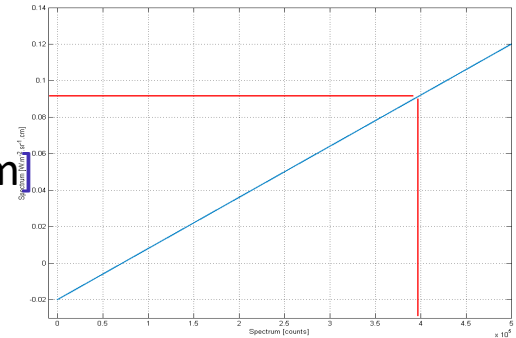
Radiometric calibration is the **same as IASI**

- 1 acquisition of cold and hot views **every 14 Earth View** acquisition
- Temporal smoothing to reduce noise



$$S_{\text{physical}}(\sigma)$$

[W.m⁻².sr⁻¹.cm]



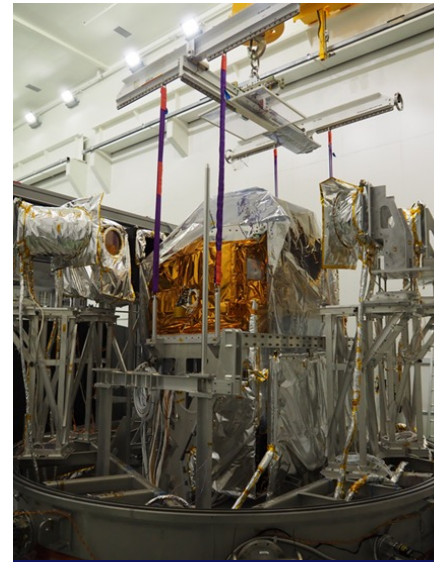
$$S_{\text{numeric}}(\sigma)$$

[counts]

- **Internal black body** is maintained at **300K**
- **2 outside views** to acquire the **cold reference**
 - Nominal view for normal acquisitions
 - One additional view for **moon avoidance**

PERFORMANCES VERIFICATION

- **Performance verification is done in two steps**
- **TVAC for assessment of spectral and radiometric fine performances**
 - 2 external warm black bodies and 1 cold black body for radiometric performances
 - 4 lasers for Instrument Spectral Response (+spectral calibration stability)
 - Gas Cell (+cold black body) for spectral calibration
 - Cold black body in front of Fabry Perot view
- **Ambient test for geometric performances (LOS, IPSF, crosstalk)**
 - Warm black body and collimator OGSE



FM2 INSTRUMENT IN TVAC CHAMBER
(COURTESY OF AIRBUS)



THERMAL TENT INSTALLATION
(COURTESY OF AIRBUS)

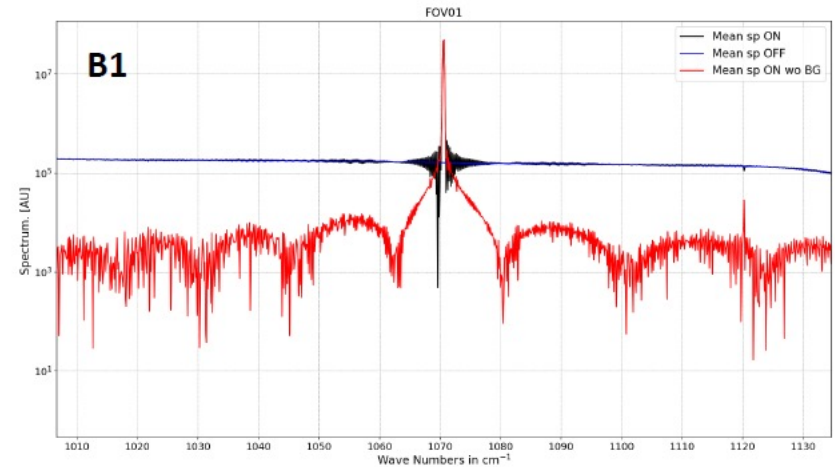


AMBIENT TEST SET-UP
(COURTESY OF AIRBUS)

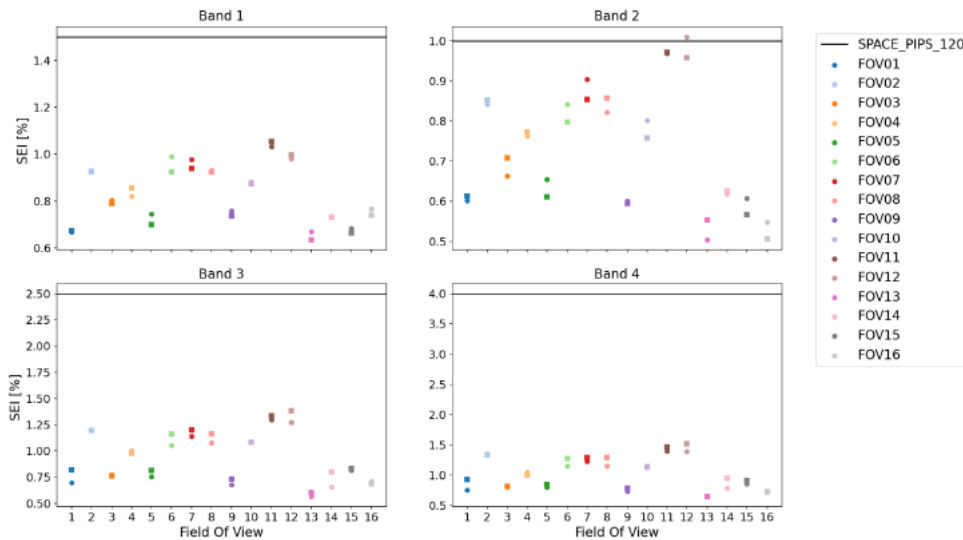
SHAPE ERROR INDEX

ISRF shape is measured against a **reference ISRF**

- 4 lasers
- Middle channel of **each band**



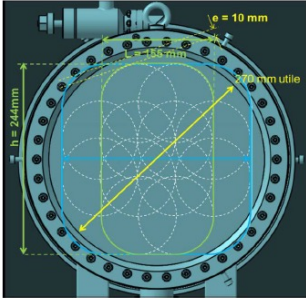
Laser test spectra with laser (black) without (blue) and background removed (red)



$$\varepsilon = \int_{-\Delta\nu_{ISRF}}^{\Delta\nu_{ISRF}} \left| \text{Re}[ISRF_2(\nu - \nu_c, \nu_0)] - \text{Re}[ISRF_1(\nu - \nu_c, \nu_0)] \right| d\nu$$

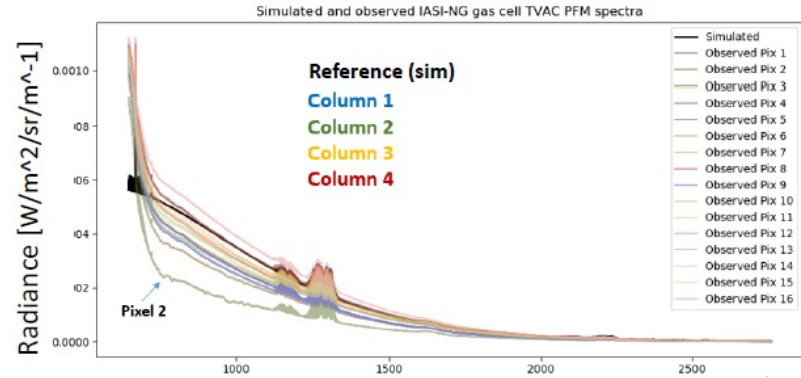
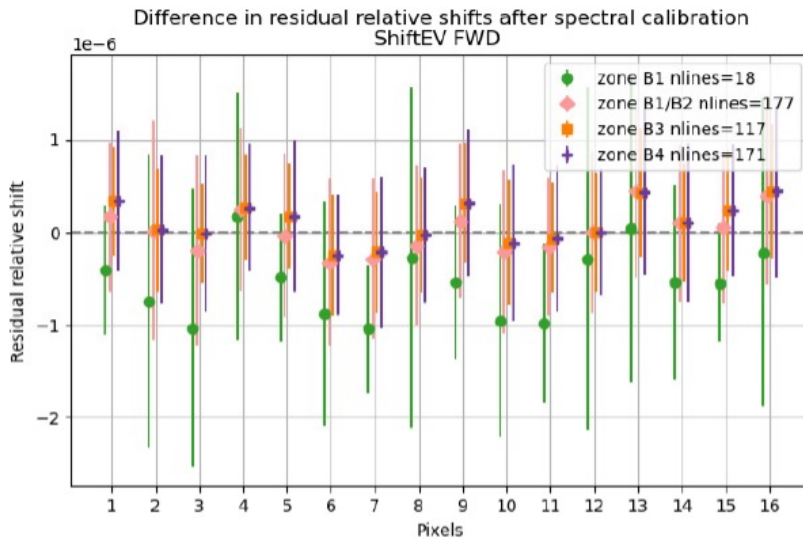
SPECTRAL SHIFT

Gas cells setup to calibrate the spectral shift on spectral lines



Pixel field of view projected onto the CAG (courtesy : ADS)

- Known concentration and pressure
- Refraction index modelled on all 4 bands at once



$$\frac{(1 + \alpha) \cdot \sqrt{n(\sigma)^2 - (\vec{r}_{\text{field}} \wedge \vec{N})^2} - (\vec{r}_{\text{field}} \cdot \vec{N})^2}{(1 + \alpha) \cdot \sqrt{n(\sigma_{\text{met}})^2 - (\vec{r}_{\text{field}}^{\text{met}} \wedge \vec{N})^2} - (\vec{r}_{\text{field}}^{\text{met}} \cdot \vec{N})^2}$$

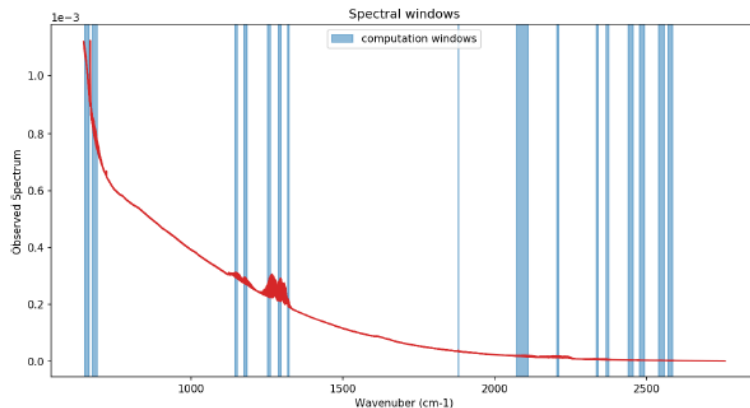
$$n(\sigma)^2 = A + \sum_{i=0}^4 \frac{B(i)}{1 - C(i)^2 \times \sigma [\mu\text{m}^{-1}]^2}$$

SPECTRAL SHIFT

The complete performance however is estimated by **analysis and simulation**

- Using contributions from **doppler residual, algorithm error, refraction index residual, reference precision, stability measurements**

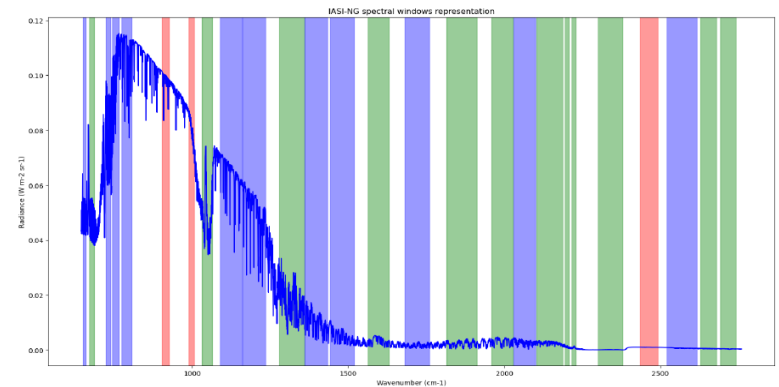
During TVAC



Gas Cells measurements

VS

In-flight



Atmospheric
simulations

Greater number of lines = Better estimation

		B1	B2	B3	B4
Specification : 1.10⁻⁶	Total	1.00 10 ⁻⁶	0.74 10 ⁻⁶	0.85 10 ⁻⁶	0.82 10 ⁻⁶

RADIOMETRIC NOISE AND BIAS

Radiometric characteristics are measured with **calibrated Black Bodies** of various temperatures

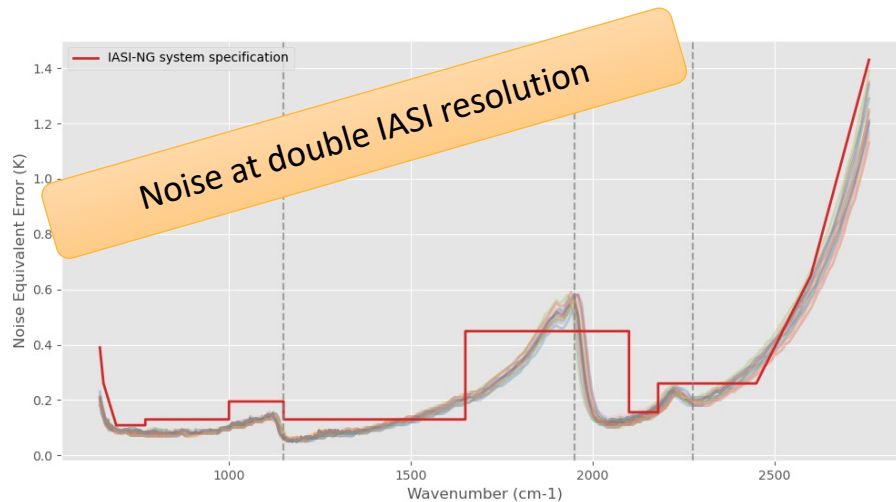
- Verification temperature range from **217K** to **310K**
- Evaluation of instrument noise and calibration bias
 - Including processing noise
 - For the complete picture, **equivalent radiometric errors** are **simulated** using spectral performances values (shape, shift, crosstalk, stability, ...)



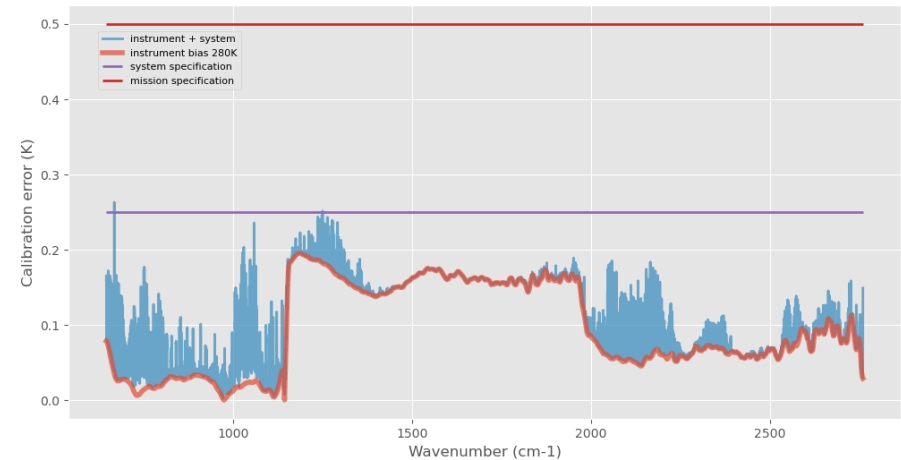
OGSE sur les vues Terre

Courtesy : ADS

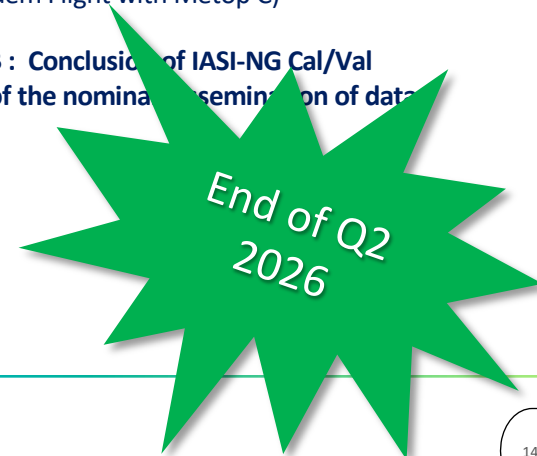
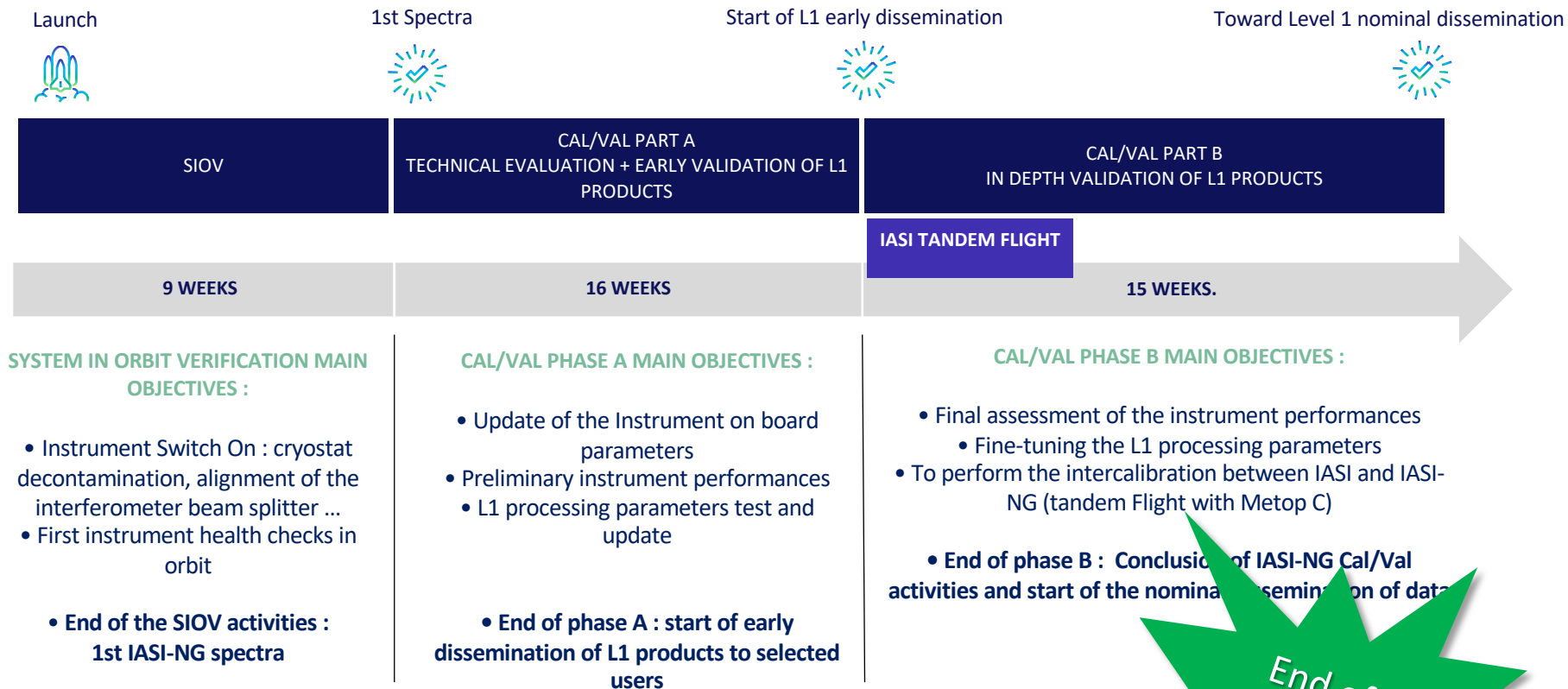
NEdT estimation with system contributors - 280K



System absolute radiometric bias at 280K - Pixel 4 (worst pixel)



SIOV/CALVAL GLOBAL OVERVIEW





THANK YOU

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