

Development and Implementation of a Physical Thermal-to-Far Infrared (TIR-FIR) Emissivity Model for Uniform Snow Surfaces within CRTM

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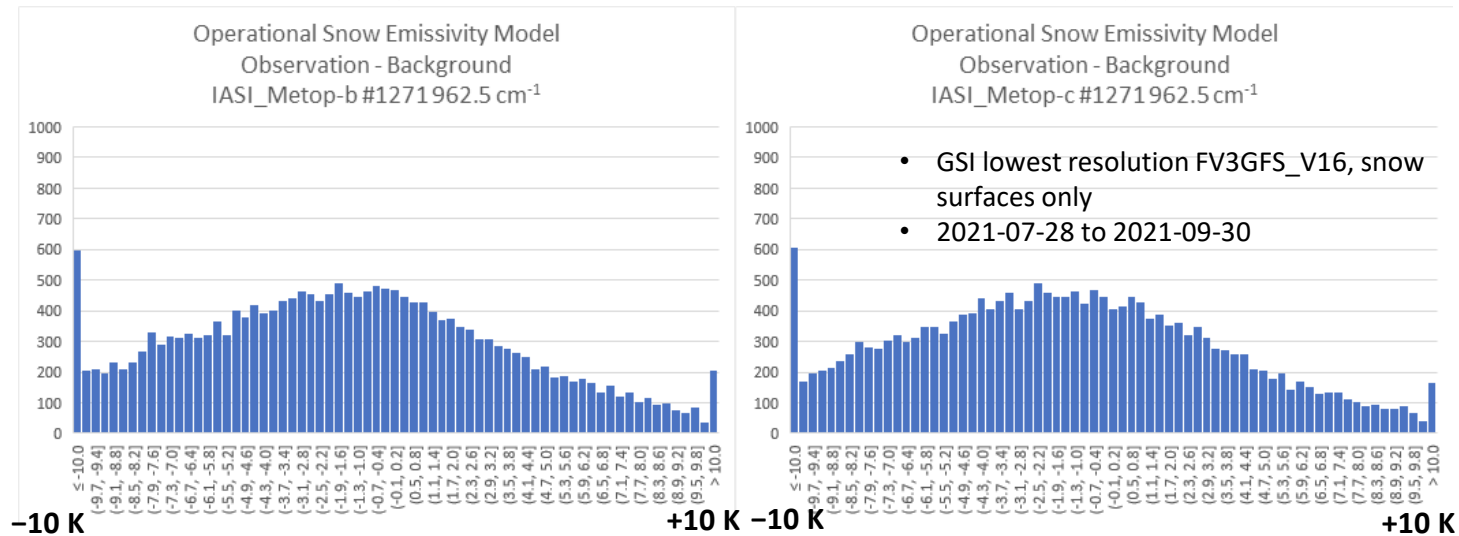
- **Background and Motivation**
 - Discrepancies in global obs – calc
 - Discrepancies in NUCAPS retrievals
- **TIR-FIR Snow/Ice Emissivity Models**
 - Physical Models
 - Empirical Models
- **Model Selection, Development, and Testing**
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 - Research code implementation
 - Problem at larger particle sizes
 - CRTM implementation and preliminary results: Improvement of global obs – calc
- **Improving the Size Dependency at Larger Snow Grain Sizes**
 - Problem at larger particle sizes
 - Observed particle size dependence
 - Hybrid Physical Model for CRTM upgrade
 - *Hori et al. (2013) Semi-Empirical Model (H13)*
 - Model results vs laboratory and field measurements
- **Summary and Future Collaborative Work**
 - Additional validation
 - CAMEL database merge
 - Implementation

Background and Motivation

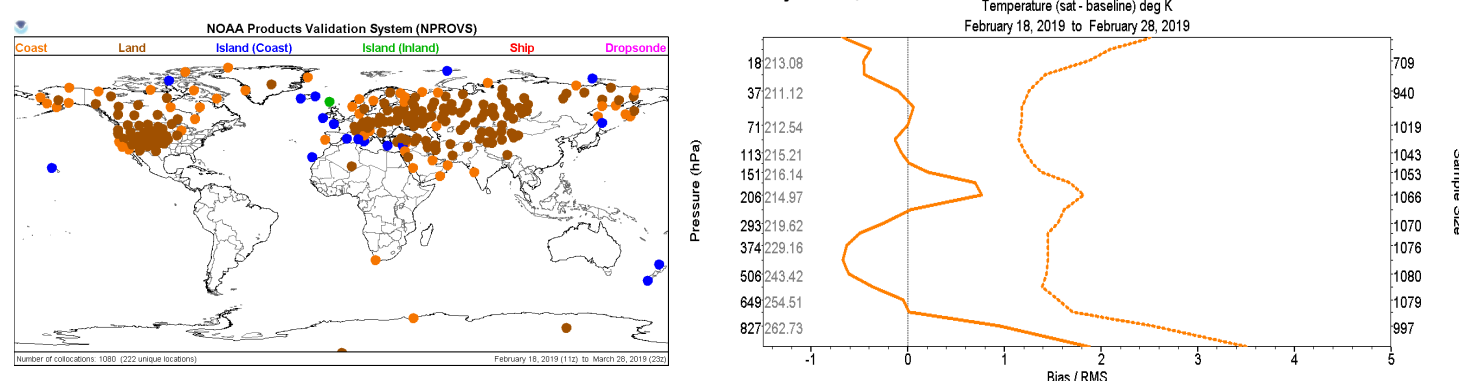


- Accurate **thermal IR (TIR)** fast-forward RT calculations (**calc**) are fundamental for radiance assimilation and operational retrieval algorithms (e.g., NUCAPS)
- **Information about the lower troposphere and surface** are derived from semi-transparent spectral window channels, which require *a priori* information about both the surface emissivity and BRDF
- **NCEP GFS assimilation studies** have revealed significant discrepancies ± 5 K RMSE in clear-sky CRTM calculations (**calc**) versus CrIS/IASI observations (**obs**) over snow/ice surfaces
- **NUCAPS temperature retrievals** exhibit similar biases associated with these factors
 - Lower-tropospheric temperatures exhibit positive bias of $\approx 1-2$ K relative to RAOBs

Global Snow obs (IASI) – calc (CRTM) GSI Control Run Metop-B Metop-C



NUCAPS Temperature Profile Statistics: Snow NOAA Products Validation System, NPROVS



Courtesy of Bomin Sun (IMSG at NOAA/STAR)

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TIR-FIR Snow Emissivity Physical Model for CRTM

TIR-FIR SNOW/ICE EMISSIVITY MODELS

Thermal IR (TIR) Snow/Ice Emissivity Models



Physical Models

- Based upon first-principles and simplifying assumptions for spectral regimes (e.g., the geometrical optics limit, etc.)
- Notable examples include
 - *Berger* (1979) model, based on geometrical optics
 - *Wiscombe & Warren* (1980) model (*WW80*), based on Mie scattering
 - *Wald* (1994) Mie correction based on “diffraction-subtraction” method

Empirical Models (Atlases/Databases)

- Based on empirical data, including both laboratory and field-measurements
- Notable examples include
 - **Combined ASTER MODIS Emissivity over Land (CAMEL)** database (*Borbas et al. 2018; Feltz-Loveless et al. 2018*), used by RTTOV
 - *Zhou et al. (2011)* hyperspectral IR derived global database



TIR-FIR Snow Emissivity Physical Model for CRTM

MODEL SELECTION, DEVELOPMENT, AND TESTING

Wiscombe & Warren (1980) Model (WW80) for Snow Albedo (1/2)



- After an extensive literature review, we chose to implement the **WW80 snow albedo model** to calculate spectral albedo, α_ν
 - Ideal in terms of both **theoretical basis** and **practical application**
 - Valid for **VIS-TIR spectrum**
 - Used by the SNICAR-ADv3 community model (*Flanner et al. 2021*)
 - Based on **Mie scattering theory** for single-scattering and **Delta-Eddington approximation** for multiple scattering in (*Joseph et al. 1976*)
 - From Kirchhoff's law, the **directional emissivity** may be computed as

$$\epsilon_\nu(\theta_0) = 1 - \alpha_\nu(\theta_0)$$

- **Delta-Eddington Approximation**

- Assumes the Eddington approximation, which assumes a truncated two-term Henyey-Greenstein (H-G) phase function
- To account for the strong forward scattering peak of snow, a **Dirac delta function term** is included
- The **phase function** is approximated as

$$P^*(\cos \Theta) \approx 2f \delta(1 - \cos \Theta) + (1 - f)(1 + 3g^* \cos \Theta)$$

where g^* is the D-E scaled asymmetry parameter

Wiscombe & Warren (1980) Model (WW80) for Snow Albedo (2/2)

- From the D-E phase function, a simplified **scaled layer RTE** may be derived as

$$\frac{dI(\theta_0, \tau_\nu^*)}{d\tau_\nu^*} \mu_0 + I(\mu_0, \tau_\nu^*) = \frac{\varpi^*}{2} \int_{-1}^1 (1 + 3g^* \mu_0 \mu) I_\nu(\mu, \tau_\nu^*) d\mu,$$

where the asterisks denote D-E scaled quantities

- The scaled RTE is then used for deriving upward and downward fluxes for deriving **albedo***, $\alpha_\nu(\theta_0)$ (i.e., **hemispherical-directional BRDF**)

*After considerable derivation; “the Devil is in the details...”

- Assuming **quasi-infinite optical depth for TIR-FIR** (e.g., *Dozier & Warren 1982*) the albedo is then shown to be

$$\alpha_\nu(\theta_0) = \varpi^* \left(\frac{1}{1 + \zeta} \right) \left(\frac{1 - \gamma \xi \cos \theta_0}{1 + \xi \cos \theta_0} \right),$$

Intermediate Parameters

$$\xi \equiv \sqrt{3(1 - \varpi^* g^*)(1 - \varpi^*)},$$

$$\gamma \equiv \frac{g^*}{1 - \varpi^* g^*},$$

$$\zeta \equiv \frac{2\xi}{3(1 - \varpi^* g^*)}.$$

DE-Transformed Mie Scattering Parameters

Single-scattering albedo:

$$\varpi^* = \frac{(1 - f) \varpi}{1 - f \varpi} = \frac{(1 - g^2) \varpi}{1 - g^2 \varpi},$$

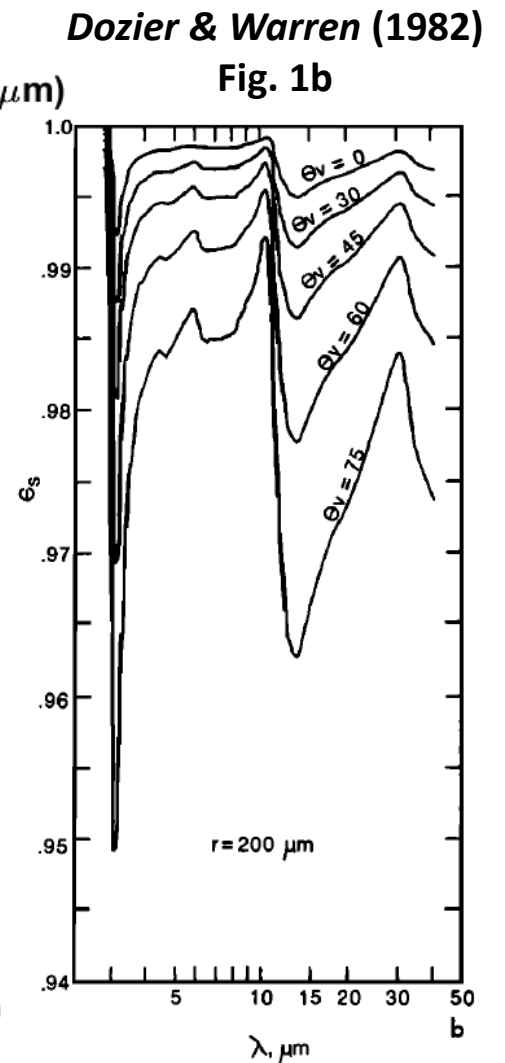
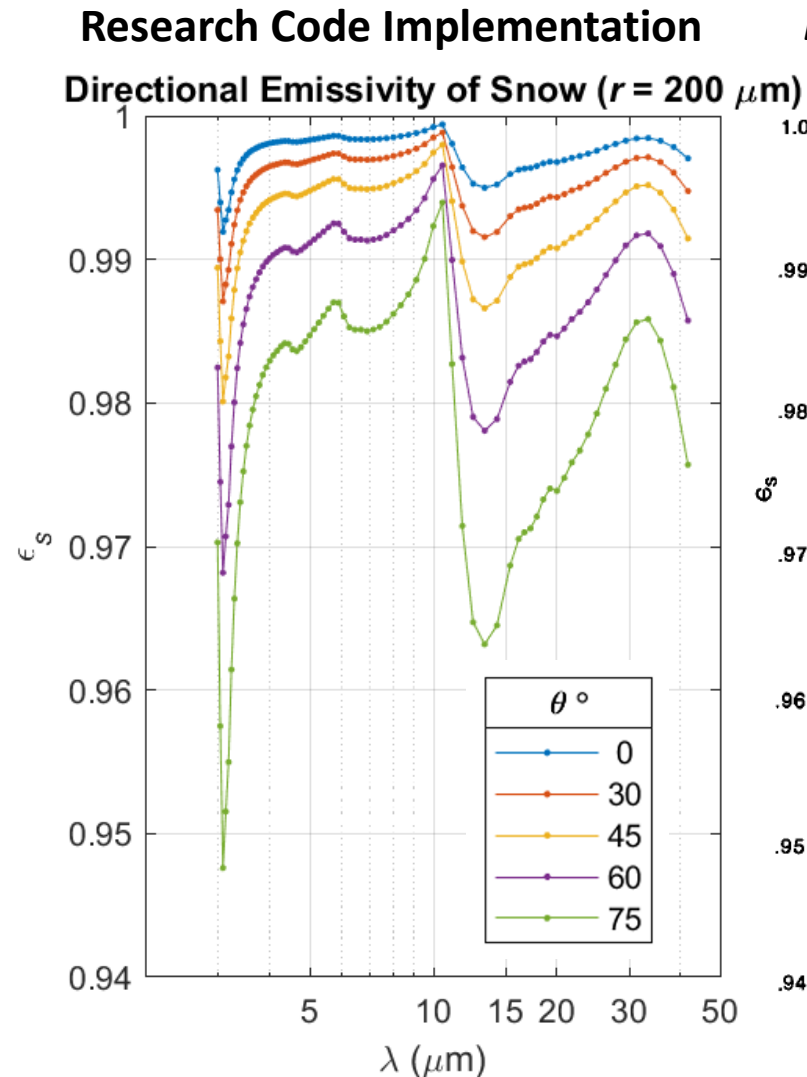
Asymmetry Parameter:

$$g^* = \frac{g}{1 + g}.$$

Research Code Implementation and Sanity Check



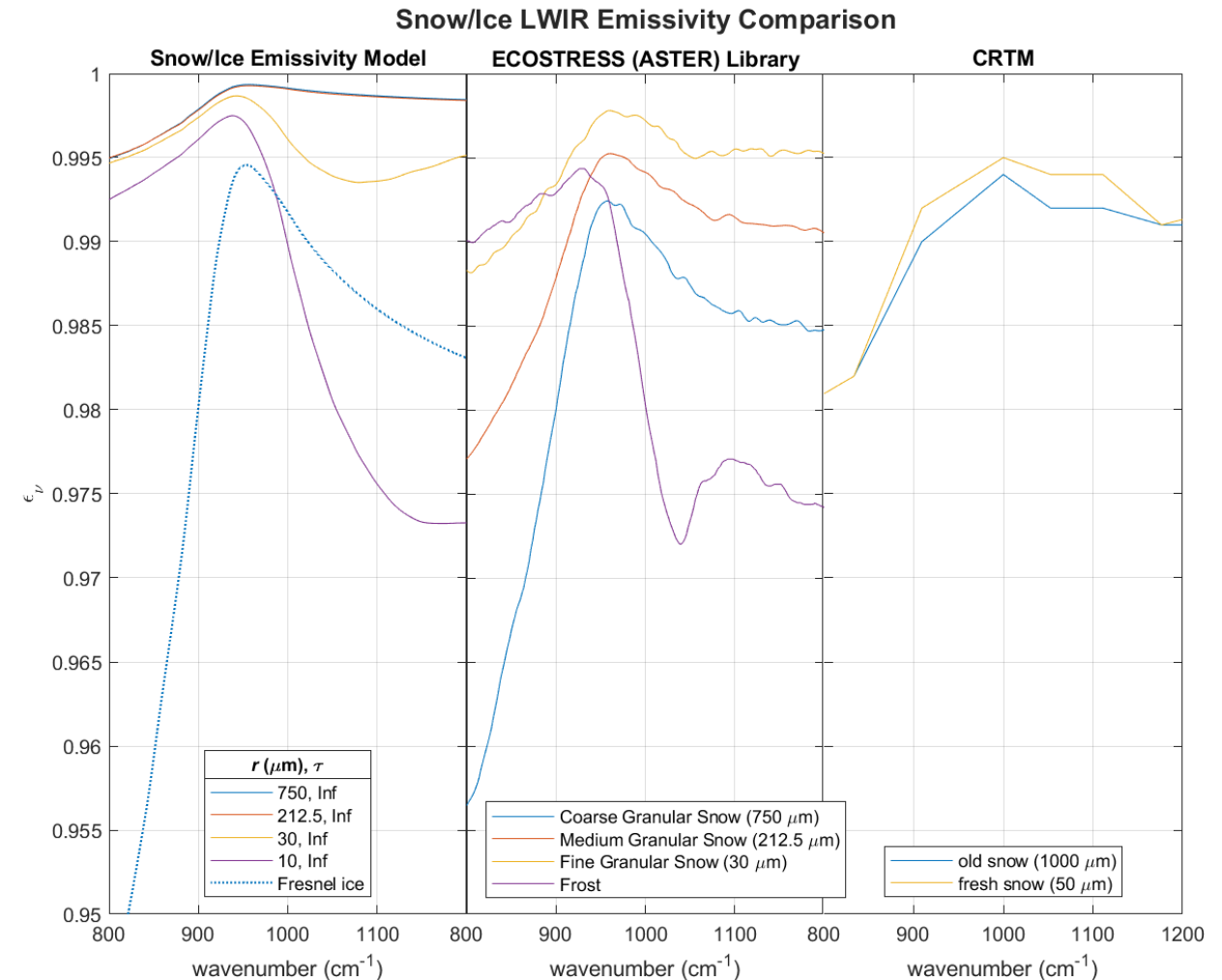
- Developed **MATLAB** research-code for performing Mie calculations
 - MATLAB is extremely well suited for performing Mie calculations, while retaining direct correspondence with theory
 - Full-scattering *WW80* model was implemented
 - Includes solar spectrum (VIS) wavelengths
- Researched the latest published **optical constants of ice in the TIR-FIR**, and implemented the **temperature-dependent dataset** by *Iwabuchi & Yang* (2011)
- Tested output of research codes against those from earlier publications and confirmed that they are functioning correctly
- **Version 1 physical model** for CRTM output as **4-D LUT $\epsilon(\nu, \theta, T_s, r)$**



Size Dependence Problem at Larger Snow Grain Sizes



- **Preliminary validation** against field measurements by *Salisbury et al.* (1994) from the ECOSTRESS (ASTER) Library (*Meerdink et al.* 2019)
- **Magnitudes are reasonably comparable**, but problems have been identified
 - Significant discrepancies seen in the **spectral dependence on particle size**
 - **Model breaks down for larger snow grain sizes** (for $r \geq 25$; e.g., *Warren 2019*)
 - There is also discrepancy and ambiguity between the sizes reported in the ECOSTRESS files versus those reported in the *Salisbury et al.* (1994) paper



CRTM Implementation and Preliminary Results (1/2)

Version 1 Physical Model



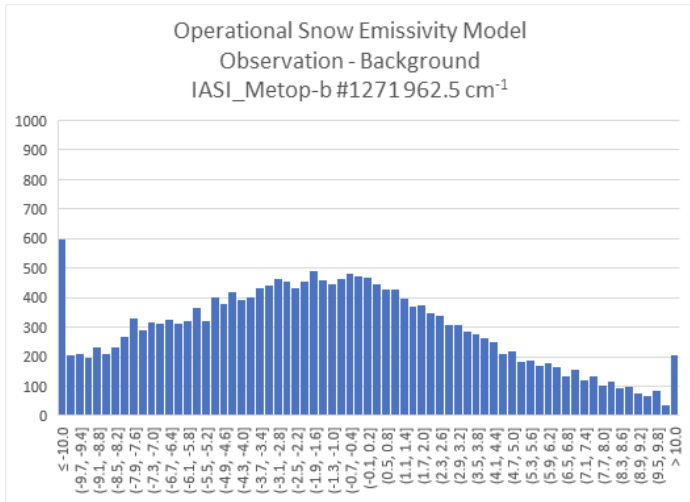
- **NCEP GSI obs – calc experiments** are being conducted testing the offline CRTM implementation (v3.0) and impact (J. Jung)
 - GSI lowest resolution FV3GFS_V16; snow surfaces only
 - **Metop-B,-C IASI**, and **NOAA-20, SNPP CrIS**, 28 July to 30 Sep 2021
 - Reset all satellite bias corrections; 5 week spin-up
- Results of **global obs – calc (O–B)** are shown in histograms (next slide)
 - Scales are –10 to +10 K
 - **Control run (Ctrl)** using the previous CRTM snow emissivity model
 - **Experimental test run (Test)** using v1 physical snow emissivity model

CRTM Implementation and Preliminary Results (2/2)

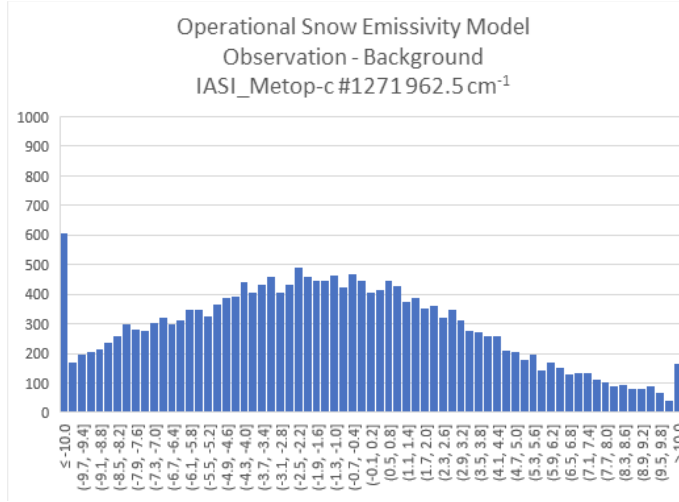
Version 1 Physical Model



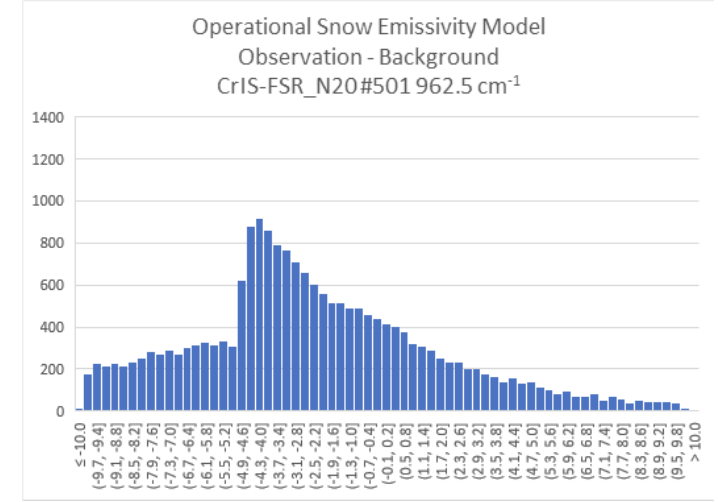
Metop-B IASI (962.5 cm⁻¹) CTRL



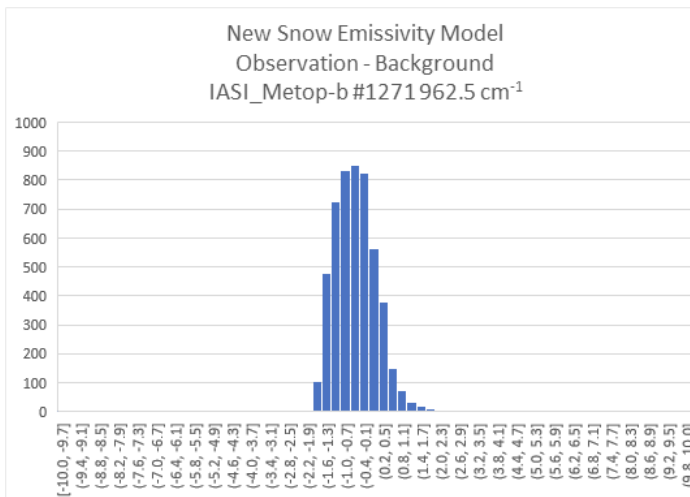
Metop-C IASI (962.5 cm⁻¹) CTRL



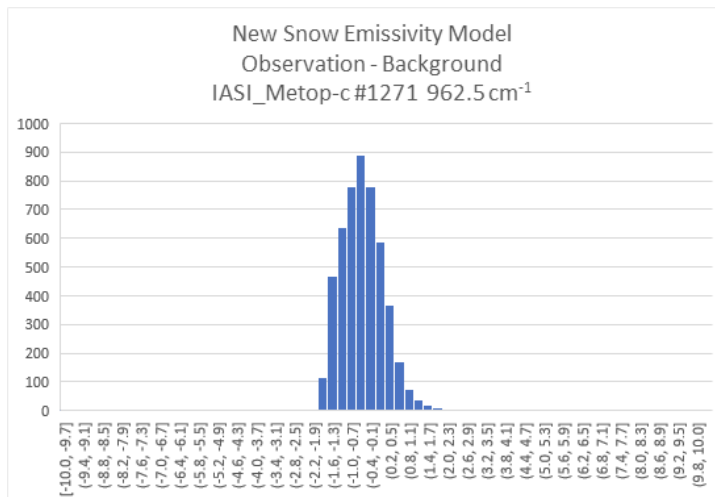
NOAA-20 CrIS (962.5 cm⁻¹) CTRL



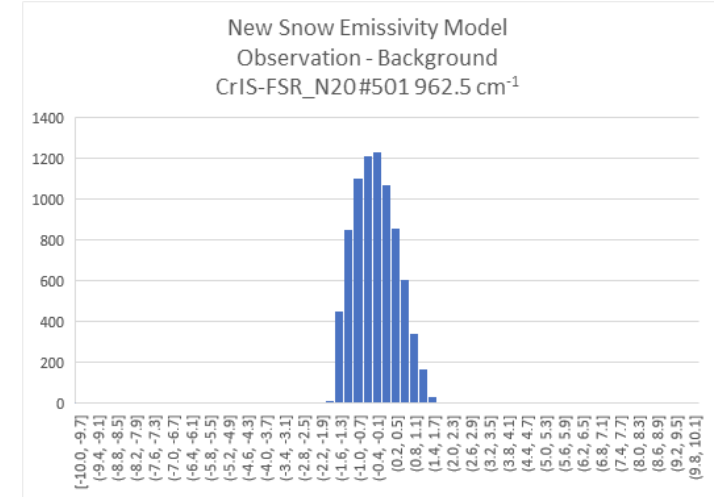
Metop-B IASI (962.5 cm⁻¹) TEST



Metop-C IASI (962.5 cm⁻¹) TEST



NOAA-20 CrIS (962.5 cm⁻¹) TEST



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TIR-FIR Snow Emissivity Physical Model for CRTM

IMPROVING THE SIZE DEPENDENCY AT LARGER SNOW GRAIN SIZES

Problem At Larger Particle Sizes

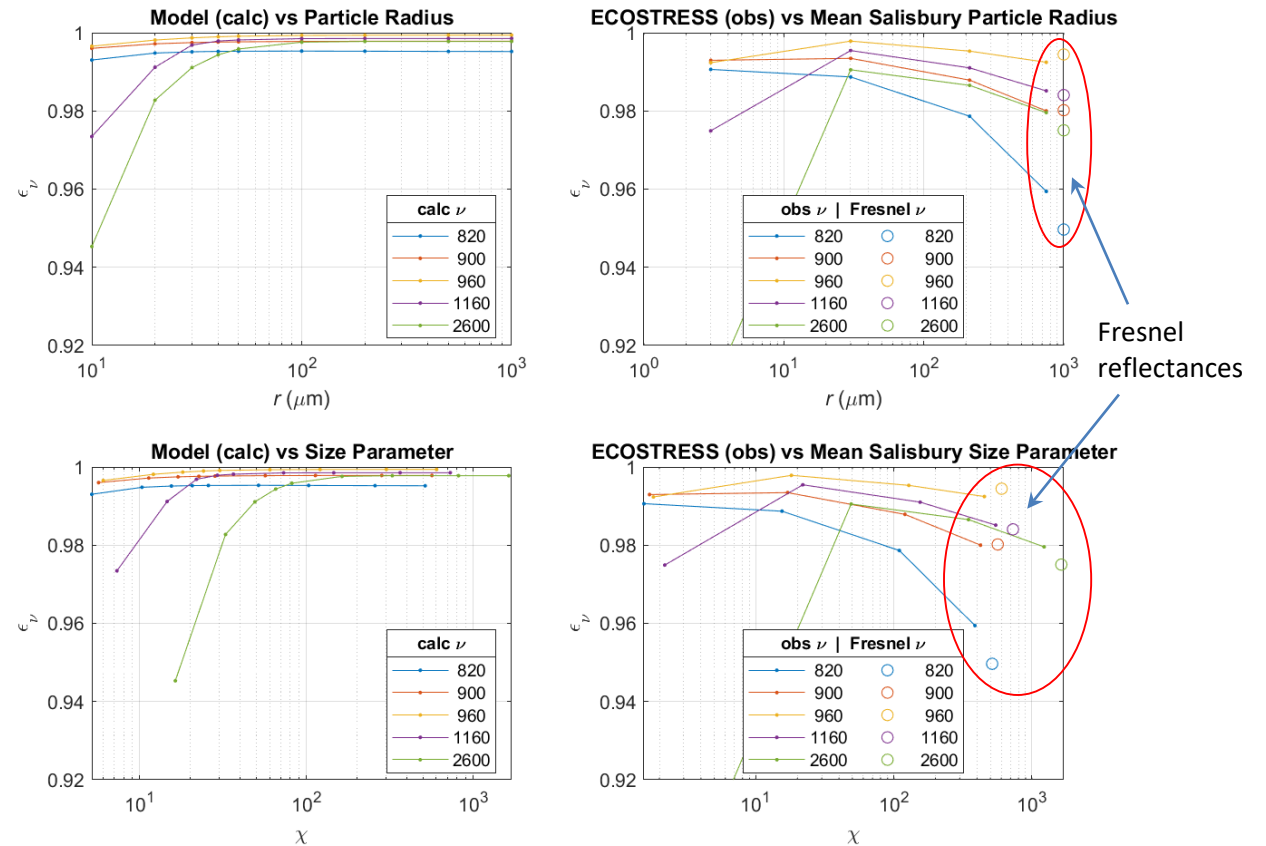


- The *WW80* model was found to agree reasonably well for smaller particle sizes ($< 30 \mu\text{m}$), but **did not capture the size dependence for larger sizes**
- It was later learned that this is a known problem with the *WW80* model (*Wald 1994, Hori et al. 2013*), but the particle size dependence was a lesser concern compared to the spectral and angular dependencies
- The improvement in global obs – calc is believed to be due to the **improved angular and spectral dependencies** afforded by the physical model
- However, **snow grain sizes are known to increase as snow ages**, so size dependency ultimately should not be ignored

Observed Particle Size Dependence

- Additional research into this problem has led to the observation that the albedo of snow surfaces **gradually become more specular** as snow ages (i.e., becomes coarser grained)
- We directly noticed this ourselves when studying the field and laboratory data, but we subsequently learned that others had also noted this previously
- As the snow ages, the particle sizes not only grow larger, but they also tend to **“weld” together** (Wald 1994) into an “ice-like” surface as opposed to a scattering layer
- Examination of the particle size dependence seemed to reveal a **linear transition** from the layer scattering regime (WW80) to a Fresnel surface reflectance regime

Snow Model v1 (calc, $\theta = 10^\circ$, $T_s = 260$ K) and ECOSTRESS (obs) vs Particle Size



Hybrid Physical Model for CRTM Upgrade

- These previous results suggested that **snow/ice surface “effective emissivity”** could be modeled such that

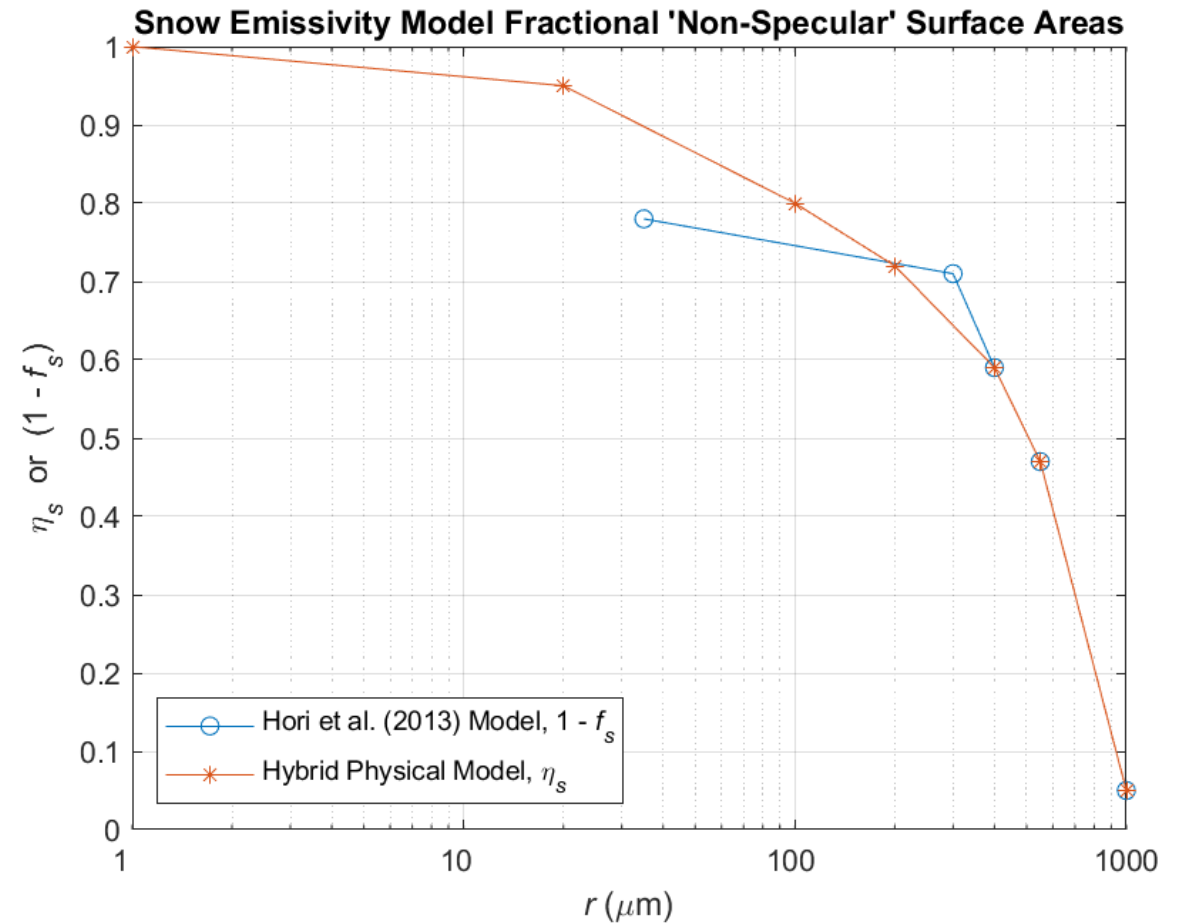
$$\epsilon_{\nu}^*(\theta_0, r) = \eta_s \epsilon_{\nu}(\theta_0, r, N_{\nu}) + (1 - \eta_s) [1 - \rho_{\nu}(\theta_0, N_{\nu})],$$

where η_s is the fractional surface area that behaves as a multiple-scattering layer, and $1 - \eta_s$ is the area of specular (Fresnel) reflectance

- In a similar vein, *Hori et al. (2013)* had proposed a **“semi-empirical” model** as a linear combination of an isotropic blackbody component and an “apparent” emissivity of the specular fraction, $f_s \sim 1 - \eta_s$
 - Empirically determined the specular fractions, f_s , for various snow samples from field observations of *Hori et al. (2006)*

$$\epsilon_{\nu}^*(\theta_0, r) = (1 - f_s) + f_s \cdot \{f_s [1 - \rho_{\nu}(\theta_0, N_{\nu})] + (1 - f_s)[1 - \rho_{\nu}(45^{\circ}, N_{\nu})]\},$$

- The ***Hori et al. (2013) (H13) model*** compares favorably for larger particle sizes, but **not** for small particle sizes
- We thus chose to adopt the “effective emissivity” approach proposed above, using specular fractions similar to *H13*, but optimized to accommodate better the smaller particle sizes treated by the *WW80* model



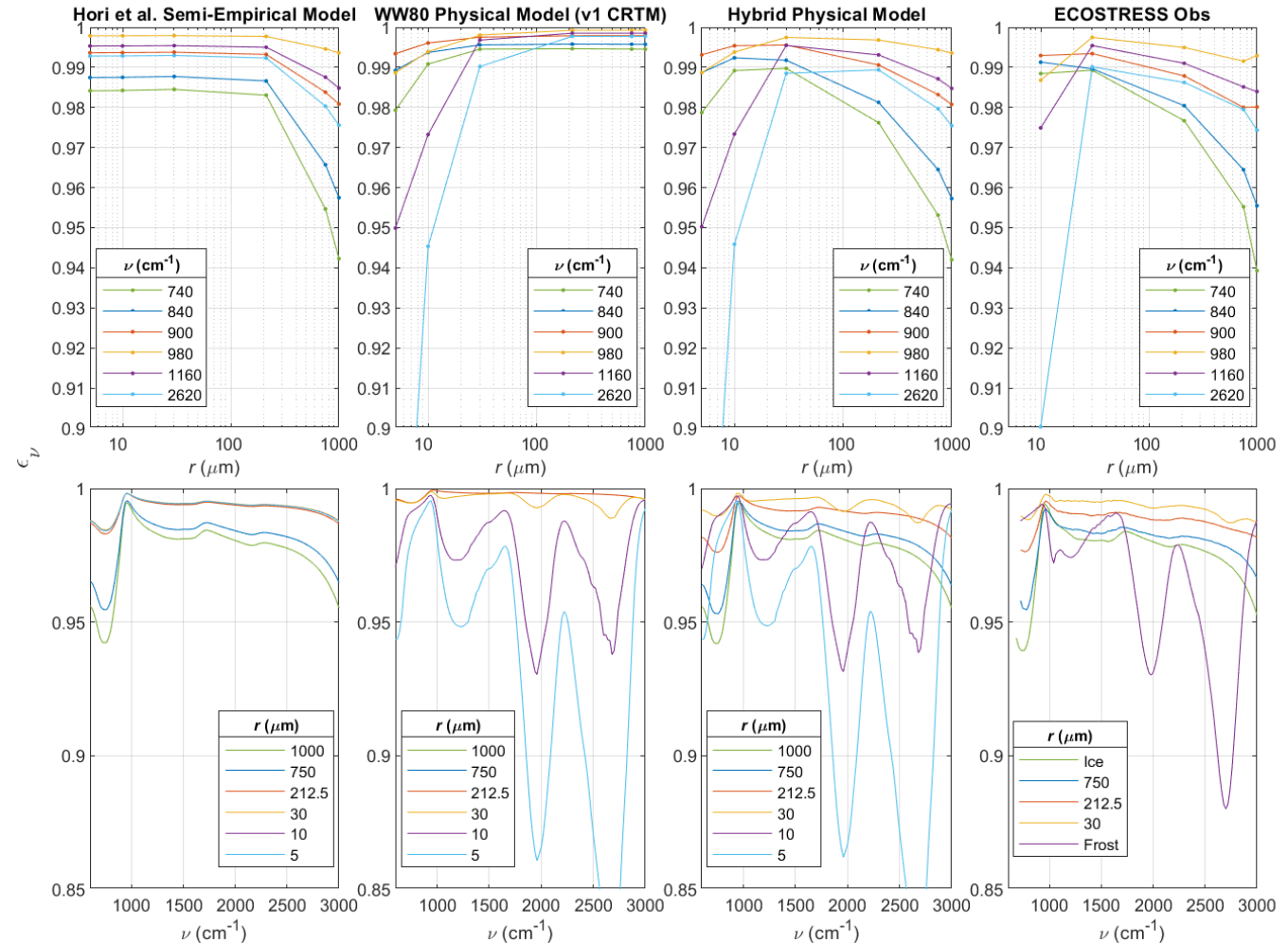
Model Results vs Laboratory Measurements

from ECOSTRESS Library



- **ECOSTRESS library** (Meerdink et al. 2019) includes measurements of snow albedo spectra taken at $\theta = 10^\circ$ (Salisbury et al. 1994)
 - Frost ($r \approx 3.25 \mu\text{m}$)
 - Granular snow
 - Fine ($r \approx 30 \mu\text{m}$)
 - Medium ($r \approx 212.5 \mu\text{m}$)
 - Coarse ($r \approx 750 \mu\text{m}$)
 - Ice ($r \geq 1000 \mu\text{m}$, flat)
- WW80 and H13 models perform well for small and large particle sizes, respectively
- The hybrid physical model captures both regimes, including the prominent low-emissivity spectral features known as **reststrahlen bands**

Snow Model Emissivity ($\theta = 10^\circ$) vs ECOSTRESS (Salisbury et al. 1994)



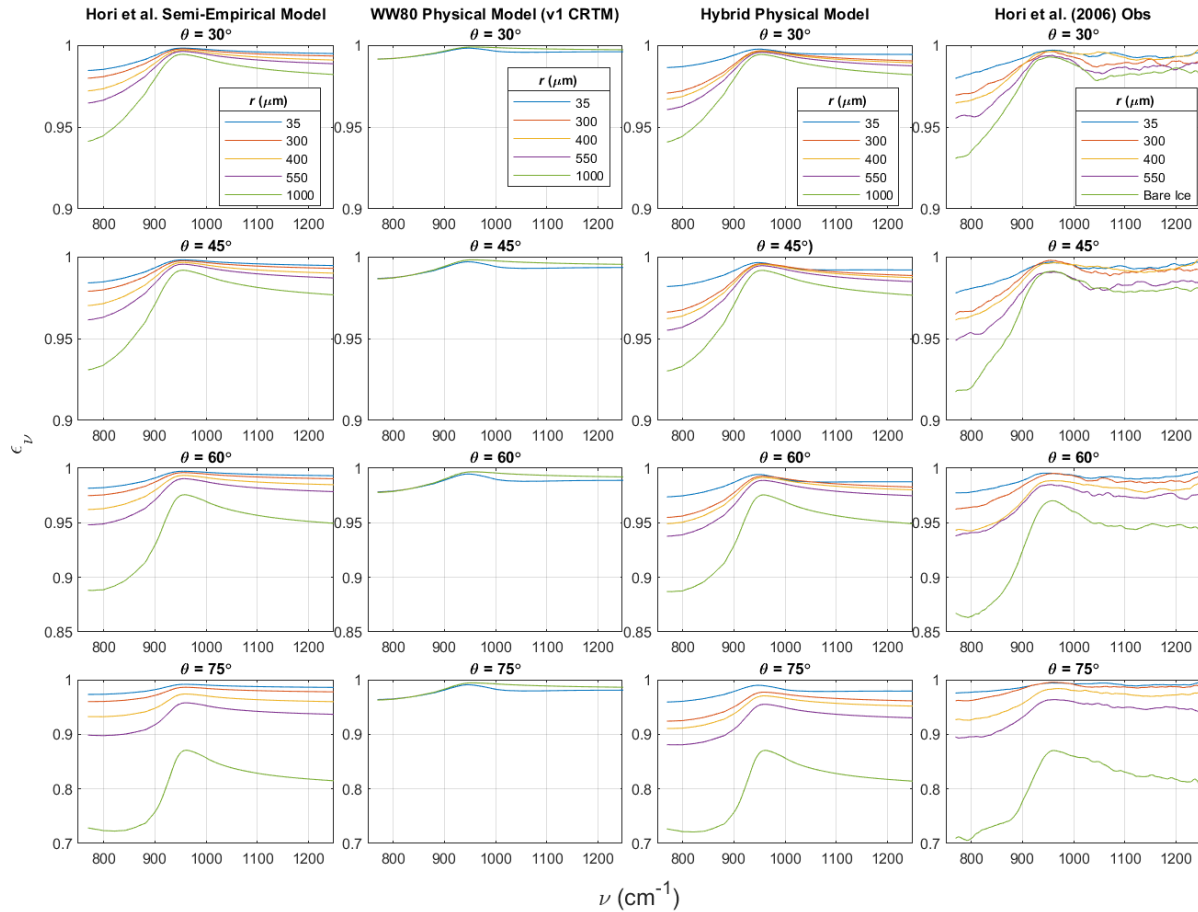
Model Results vs Multi-Angular Field Measurements

from *Hori et al. (2006)*



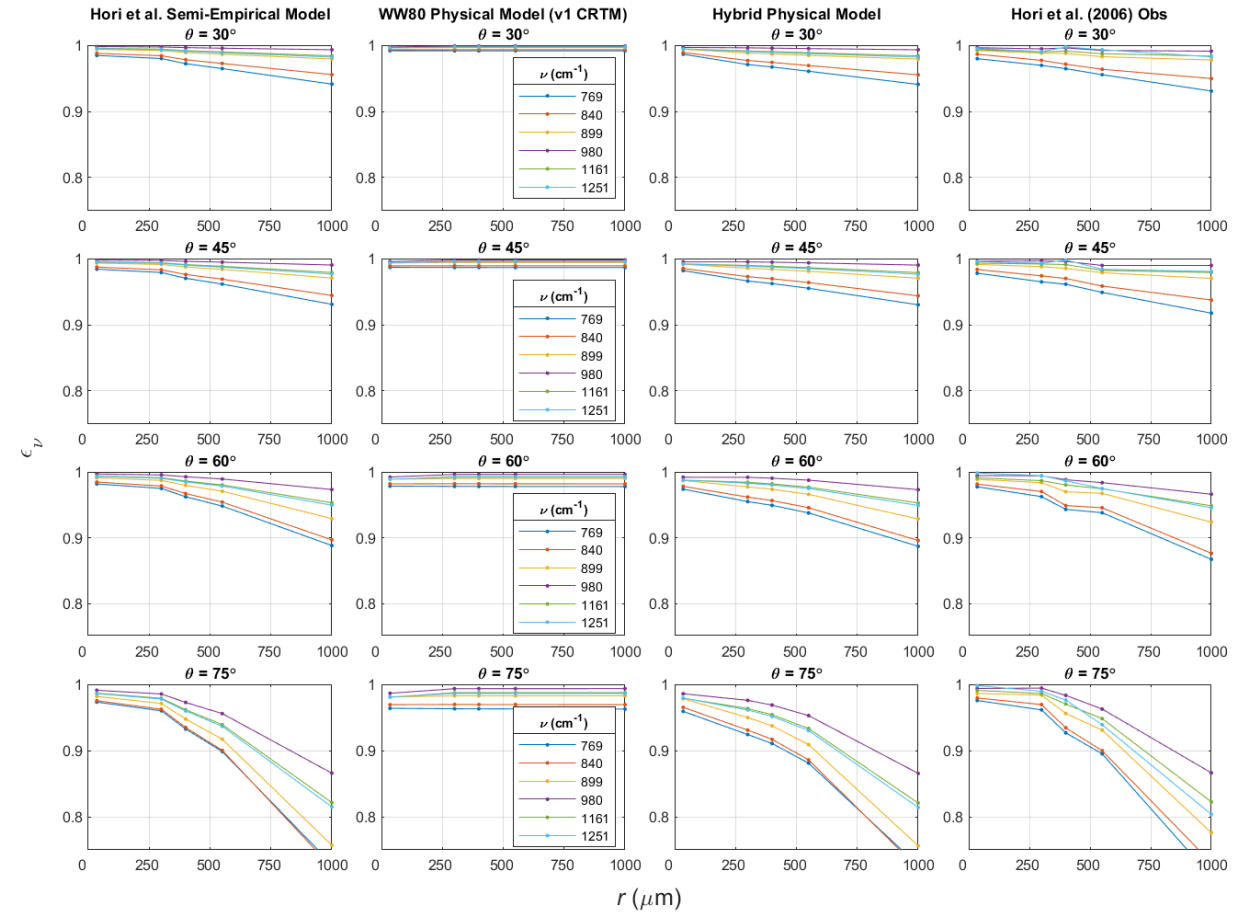
Spectral Dependencies

Snow Model Emissivity Spectral & Angular Dependence vs Hori et al. (2006) Observations



Grain Size Dependencies

Snow Model Emissivity Grain Size & Angular Dependence vs Hori et al. (2006) Observations



Summary and Future Collaborative Work



- A **physical TIR-FIR snow emissivity model** based on Mie scattering and D-E approximation (*WW80*) was developed (as research code) and implemented within CRTM (v1 physical model)
 - The albedo model itself is valid to VIS wavelengths
- Preliminary analyses of global obs – calc are encouraging, **nearly an order of magnitude improvement** over snow surfaces
 - The v1 physical model improvement in global obs – calc is believed to be due to the improved angular and spectral dependence afforded by the physical model
- A **hybrid physical model** (similar to *H13*) was developed to extend the v1 (*WW80*) physical model to larger particle sizes
- The **optical constants** used in the hybrid physical model include *T*-dependence and **span the TIR-FIR spectral range**
 - FIR is of interest for polar remote sensing (Loveless et al.)
- **Additional validation** of hybrid physical model against MAERI/ARI field/laboratory data
 - Polar MOSAIC campaign
 - SSEC “rooftop” laboratory (Loveless, Knuteson, Taylor, Revercomb, et al.)
- **Merge with CAMEL** to account for non-uniform surfaces (Borbas et al.)
- **Implementation within CRTM/GSI** and global obs – calc experiments (Jung, Dang, Johnson, et al.)
- **Implementation within NUCAPS** (Pryor, Divakarla, Zhu, Zhou, DeSouza-Machado, et al.)

TIR-FIR Snow Emissivity Physical Model for CRTM

THANK YOU! QUESTIONS?



TIR-FIR Snow Emissivity Physical Model for CRTM

BACKUP SLIDES