

## Assessing the accuracy of the line-by-line 4A/OP model through comparisons with high spectral resolution IASI observations

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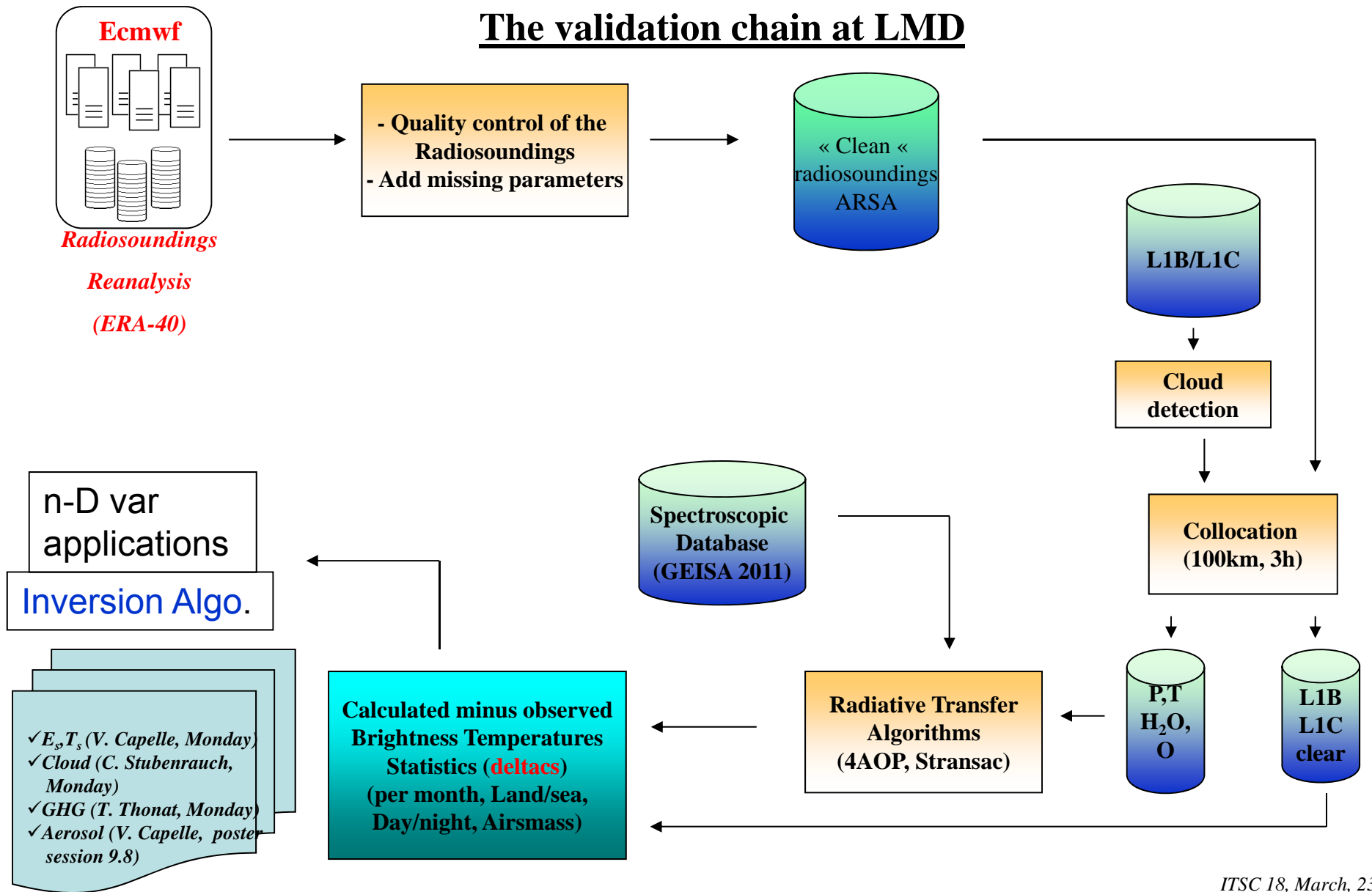
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### Plan

- The validation chain at LMD
- Results and analysis of the bias
- Conclusions and outlook

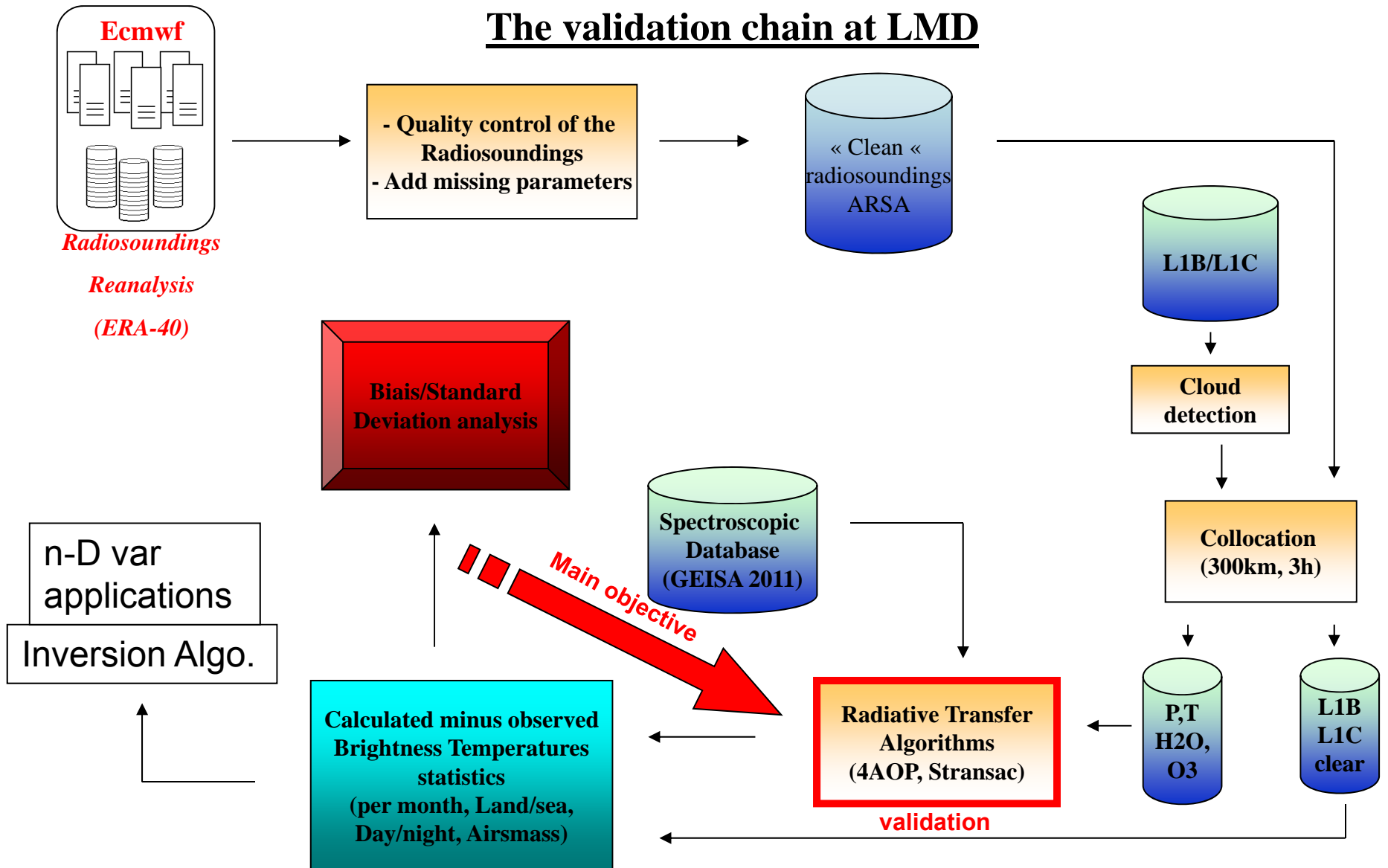


## The validation chain at LMD



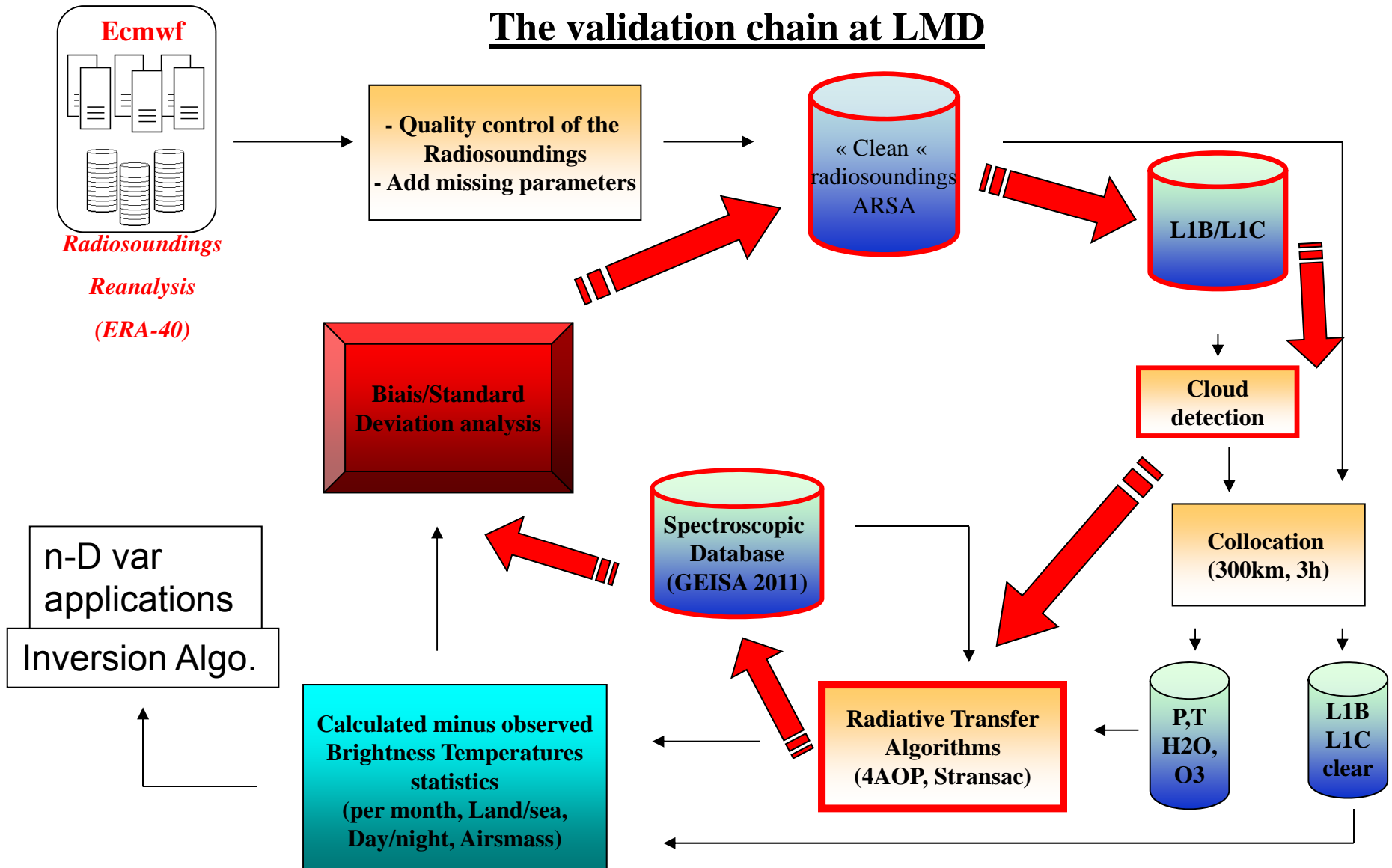


## The validation chain at LMD





## The validation chain at LMD



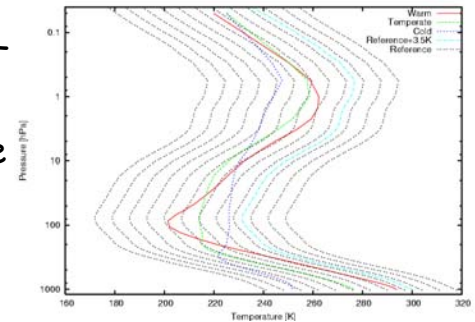
Study of the various biases: spectroscopy, RT, obs, etc.

## Radiative Transfer Model : 4AOP

4A is based on the calculation of atlases of optical thicknesses :

- ✓ for up to 53 atmospheric molecular species of the GEISA database (<http://ether.ipsl.jussieu.fr> ;
- ✓ for 12 nominal atmospheres temperature profiles representative of the T variation in the atmosphere);
- ✓ for a set of 44 pressure levels between surface and top (100km) of the atmosphere;
- ✓ for a  $5 \cdot 10^{-4} \text{ cm}^{-1}$  nominal spectral step (lower if necessary) ;

*See Chaumat et Al,  
poster session 8.3*



4A allows accurate and fast (50 times faster) computations

NOVELTIS is in charge of the industrialization and the distribution of 4A, in accordance with a convention signed between CNES, LMD/CNRS and NOVELTIS, and the LMD is in charge of the update of the physics and the validation which are needed.

### Main options :

- Radiance, transmission function and jacobians outputs
- Regular updating and improvements (line mixing CO<sub>2</sub>, continua, ...)
- all various instruments could be simulated (HIRS, AIRS, IASI, IIR, Modis, Seviri, ...)
- Limb simulations
- solar contribution
- Coupling 4AOP and Disort to take into account the diffusion
- Website <http://www.noveltis.fr/4AOP/>

### New updates :

- Scattering for cloud (cirrus...) contribution
- New atlases of absorption optical thicknesses:
  - ✓ Improvement of CO<sub>2</sub> line-mixing
  - ✓ New GEISA 2009 spectroscopy
  - ✓ Pressure shift for H<sub>2</sub>O, CO<sub>2</sub> and N<sub>2</sub>O
  - ✓ Update reference gas mixing ratio profile
  - ✓ Improved TIPS' formulation

### Coming ...

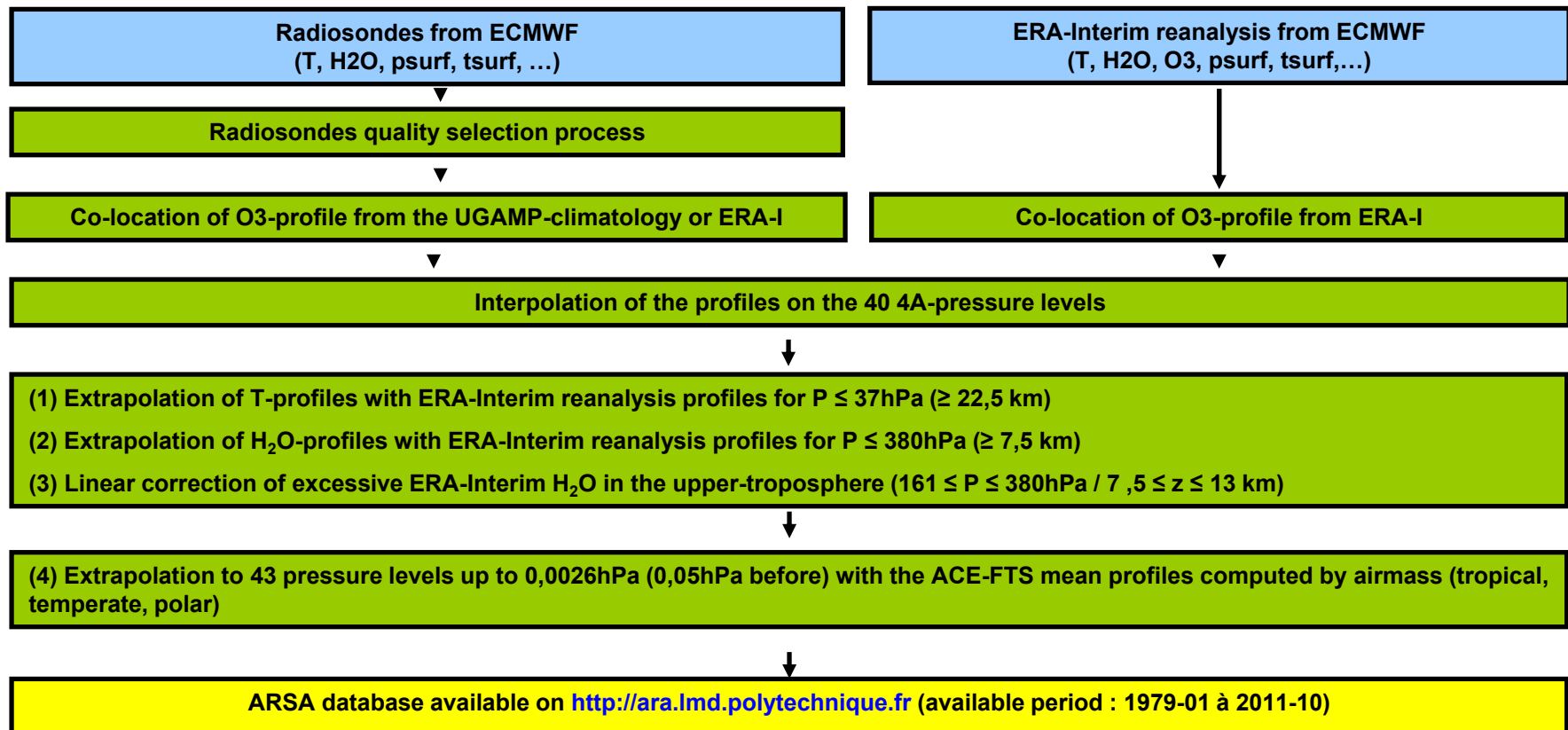
- Extension to the *SWIR* : *Merlin (CH<sub>4</sub>) and Microcarb(CO<sub>2</sub>)*

*ITSC 18, March, 23th*

## ARSA (Analyzed RadioSoundings Archive) *Scott et al, in preparation*

➤ Database gathering **radiosondes** (lat, lon, t, altitude, etc) and the corresponding thermodynamic parameters : (T, H<sub>2</sub>O) profiles, psurf, etc + O<sub>3</sub> profile **ERA\_INTERIM**

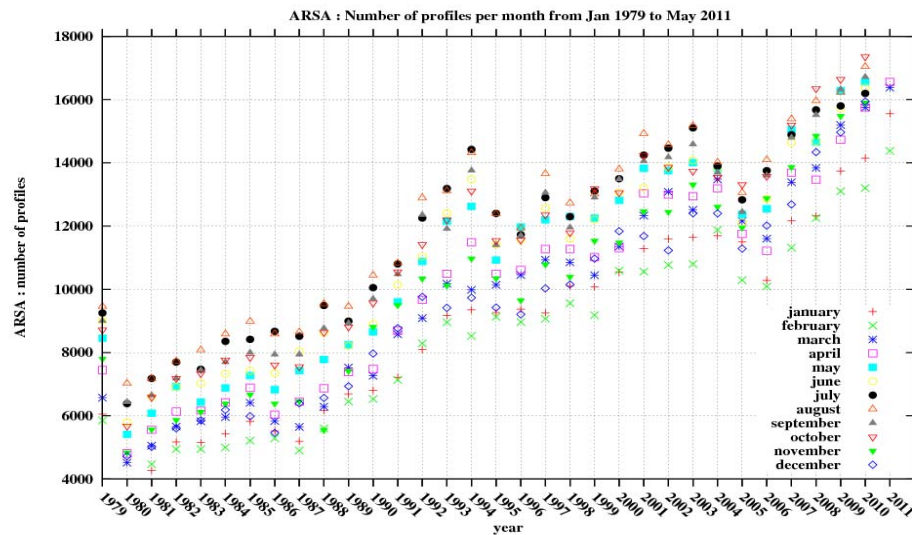
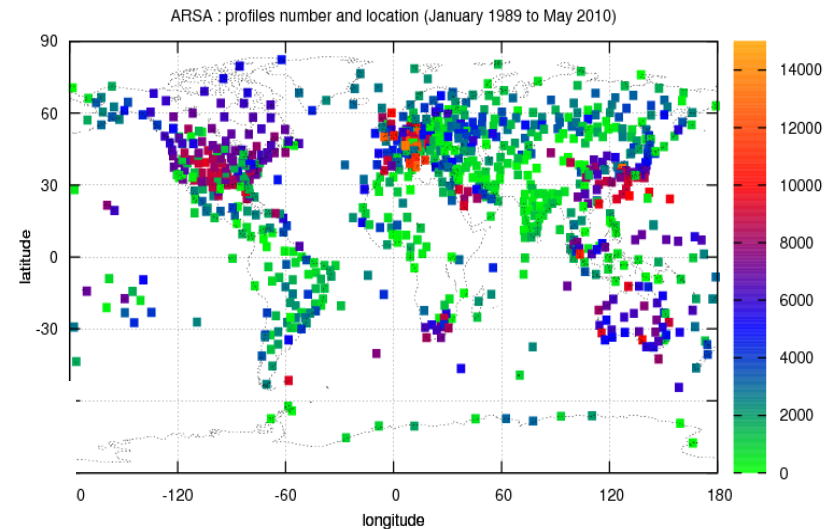
➤ **Process chain :**





# ARSA (2/2)

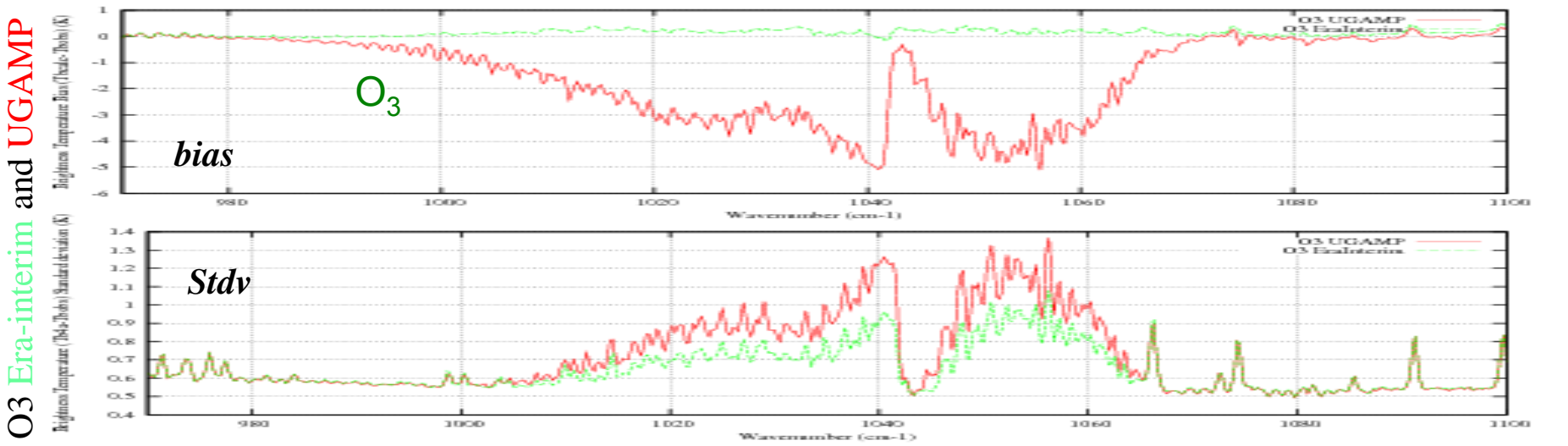
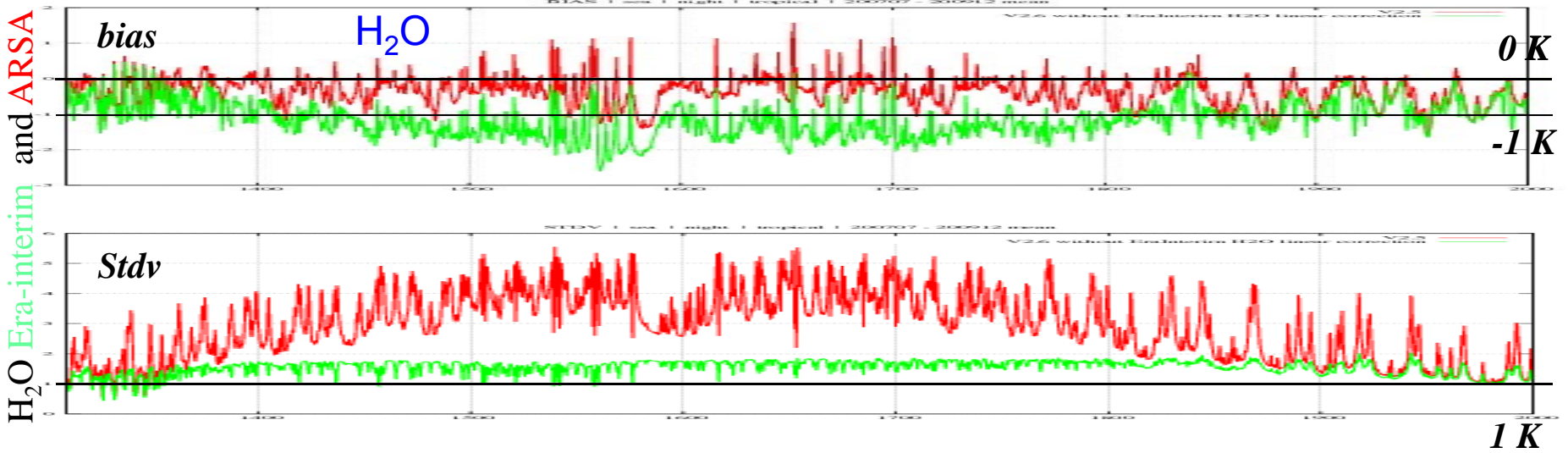
From January 1979 onwards (Jan 2012) ~  
 TOTAL > 4,000,000  
 ~ 1,100,000 tropics  
 ~ 2,600,000 midlatitudes  
 ~ 330,000 polar



➔ More than 5000 collocations IASI/ARSA

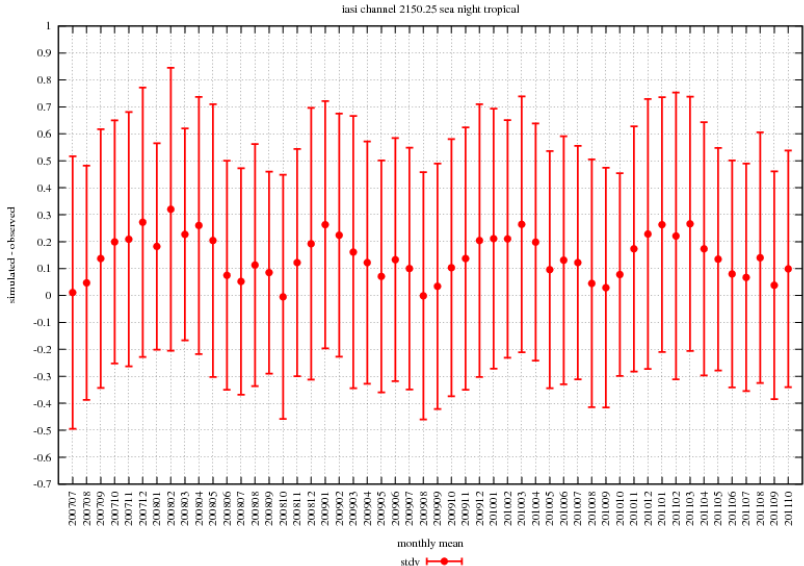
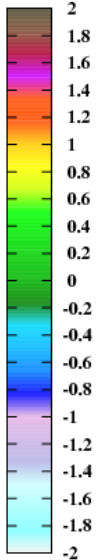
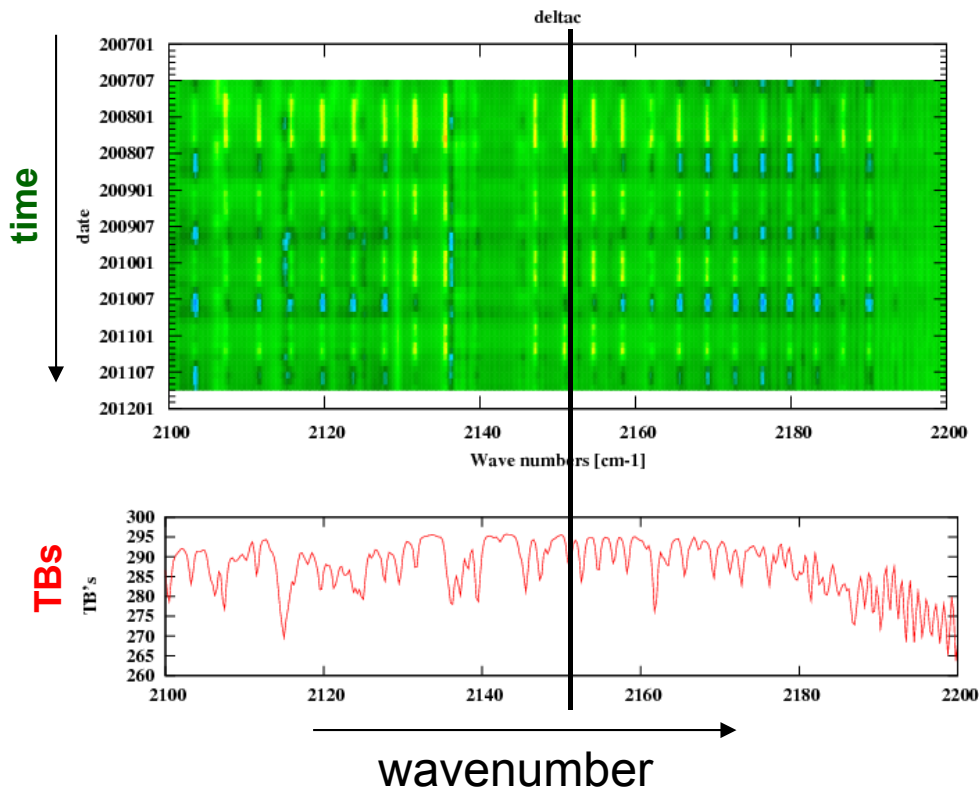
- Sea
- Night
- Tropical
- Clear

**Bias due to the thermodynamical profiles (H<sub>2</sub>O and O<sub>3</sub>)**





**« Natural » bias due to the seasonal variation of CO**



**simulated-observed** as a function of the time  
 for the 2150.25 cm-1 channel sensitive to CO  
 → Constant CO profile used in 4AOP →  
 seasonal variation of CO observed

See Thonat on monday

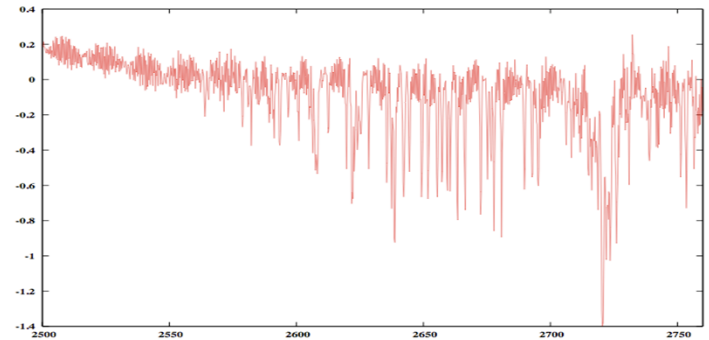
Jul 07 time Oct 11  
 ITSC 18, March, 28th ITSC 18, March, 23th

Results : In the 2500-2760 cm<sup>-1</sup> region negative bias →  
thermodynamical profiles (HDO/H<sub>2</sub>O mixing ratio)

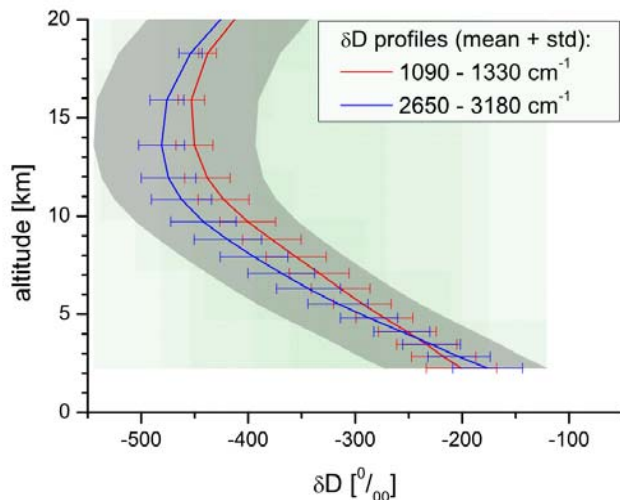
Bias between simulated and observed brightness temperatures may be as high as 1.5 K especially in the 2720. – 2730 cm<sup>-1</sup> spectral region. Sign is negative, indicating too high an absorption in this region.

Main absorber HDO or H<sub>2</sub>O? From Geisa-09 → HDO.

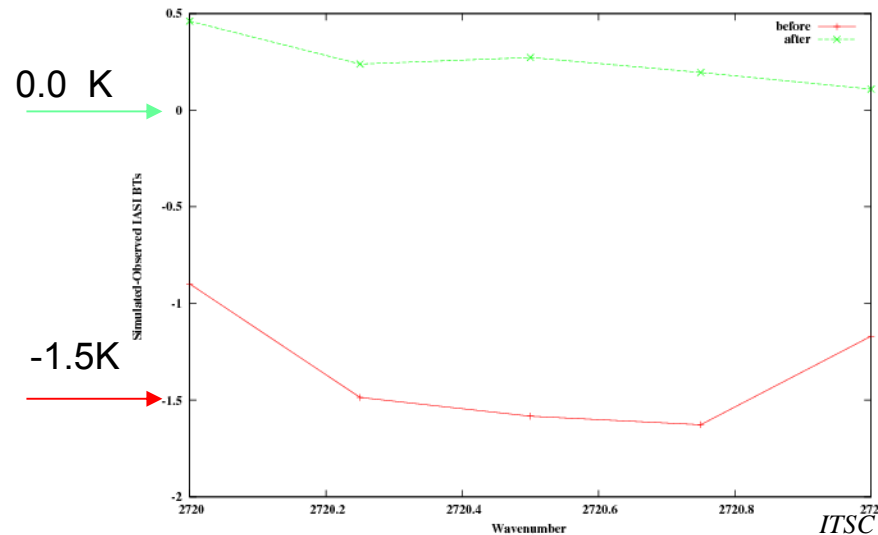
Several works indicate a **vertical variation of the δD value**  
 $\delta D = 1000 \times ([HD^{(16)}O]/[H_2^{(16)}O] / SMOW - 1)$ , with Standard Mean Ocean Water SMOW =  $3.1152 \times 10^{-4}$



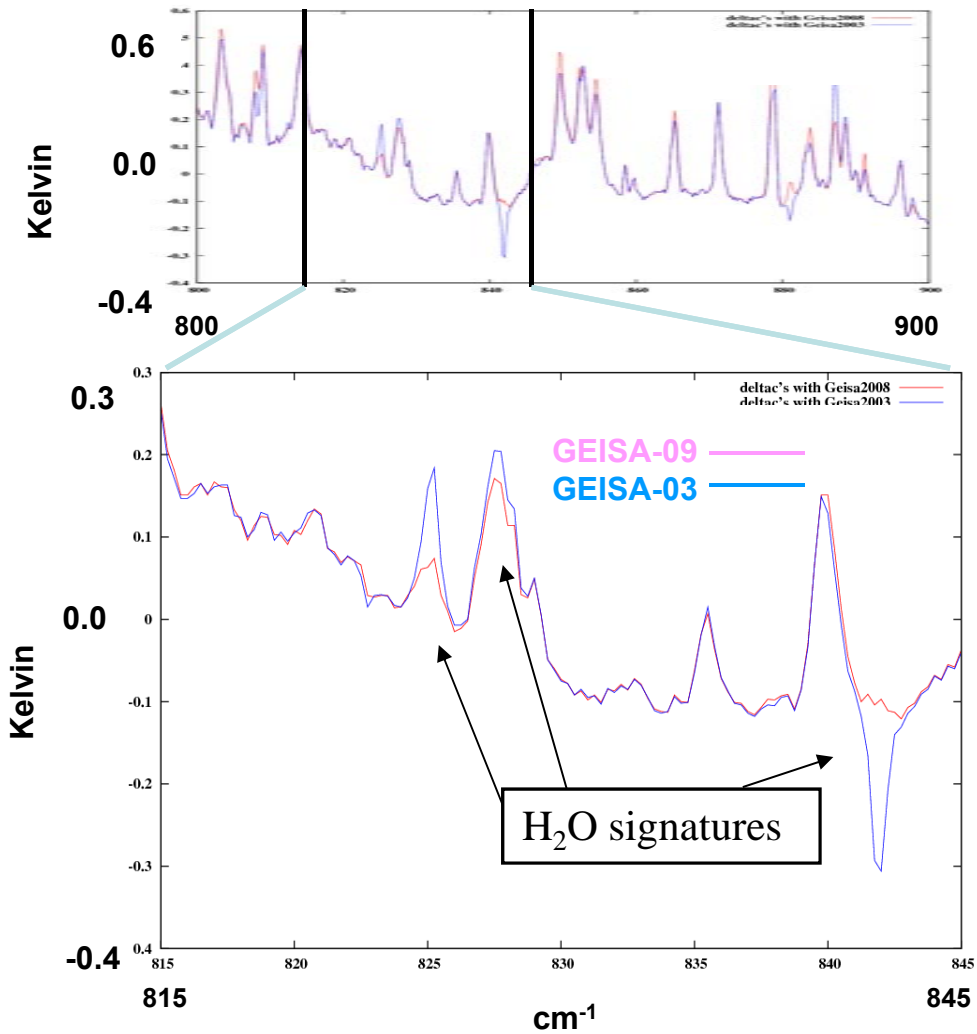
Vertical variation of the δD value :  
Impact on Simulated vs Observed  
Red = Before / Green = After



from Schneider et al, 2010.



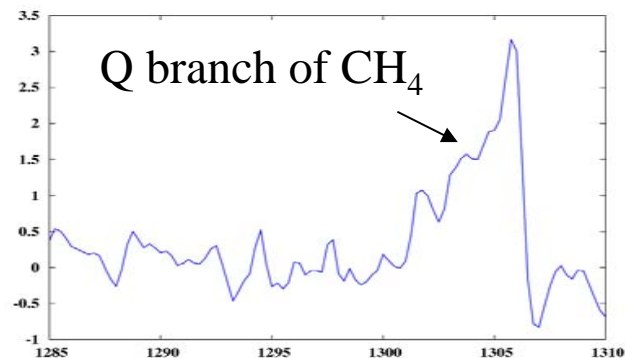
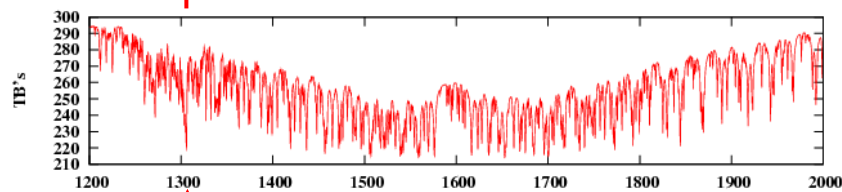
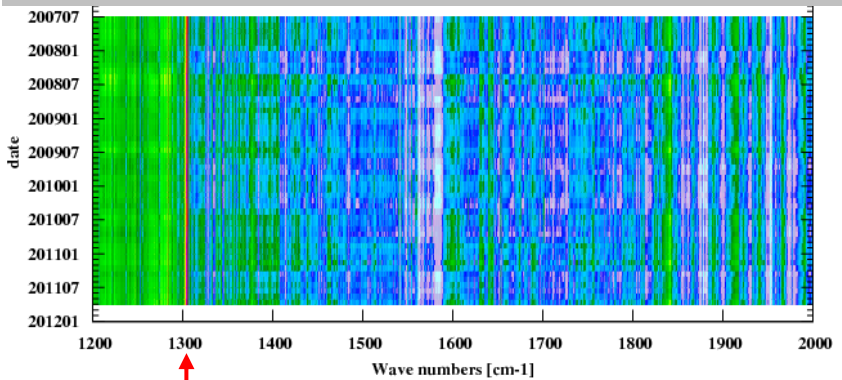
**Question: is the bias due to spectroscopic parameters ?**



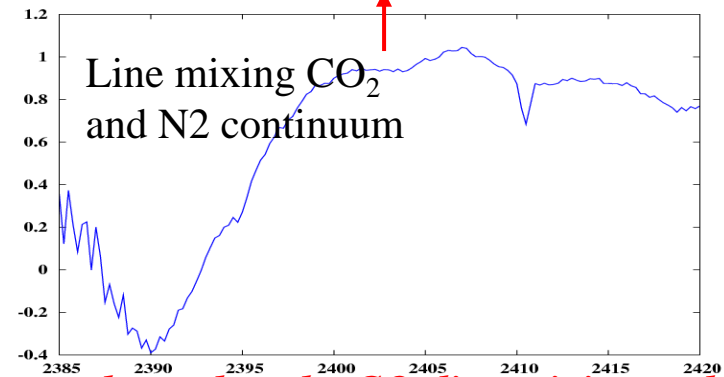
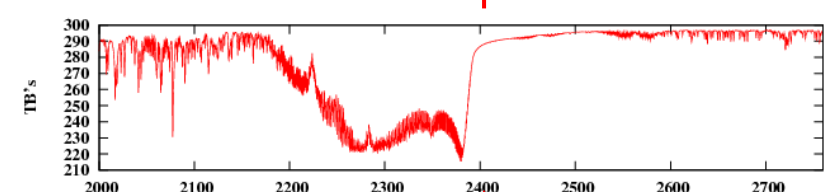
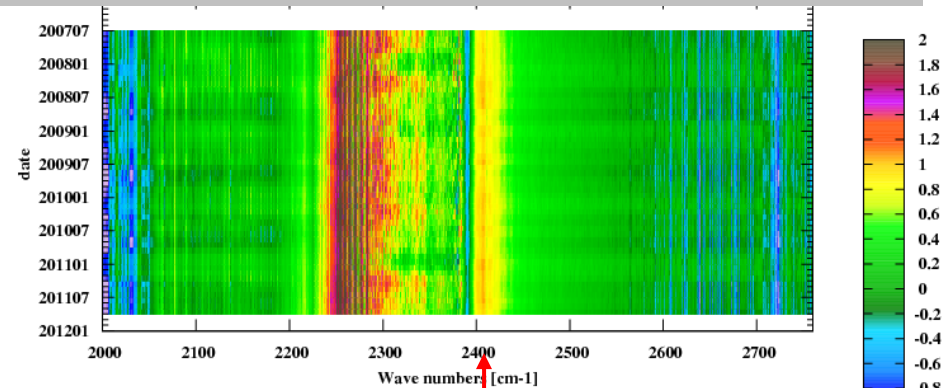
**YES → Corrected with the updated GEISA-09 version**

## Bias due to the radiative transfer algorithm (line modelling)

### CH<sub>4</sub> (7μm) and CO<sub>2</sub>/N<sub>2</sub> (4 μm)

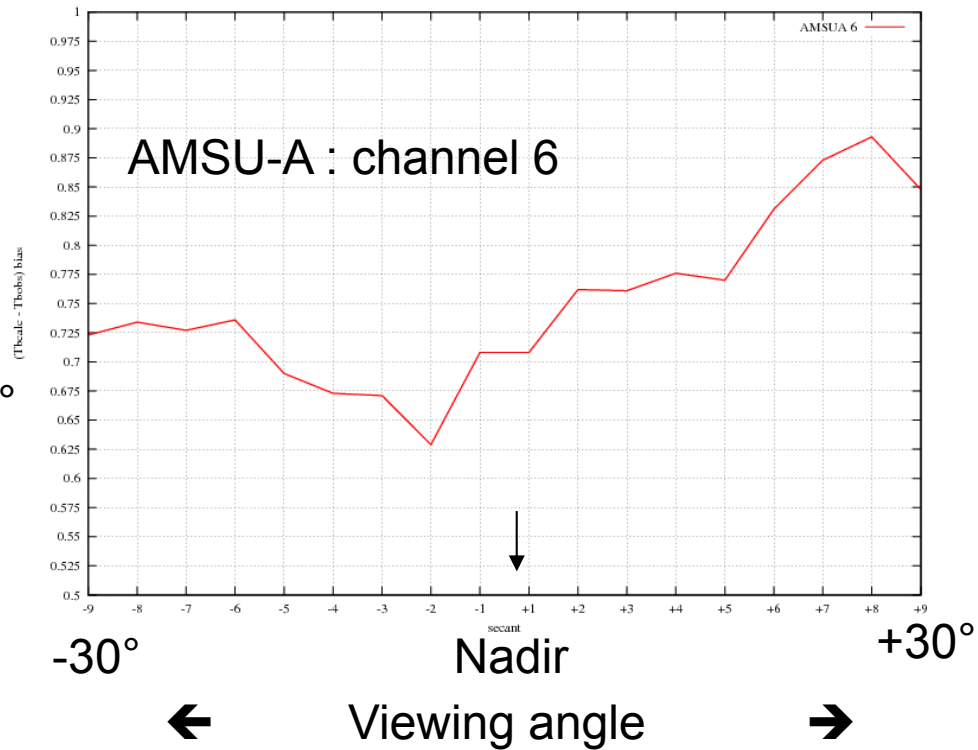
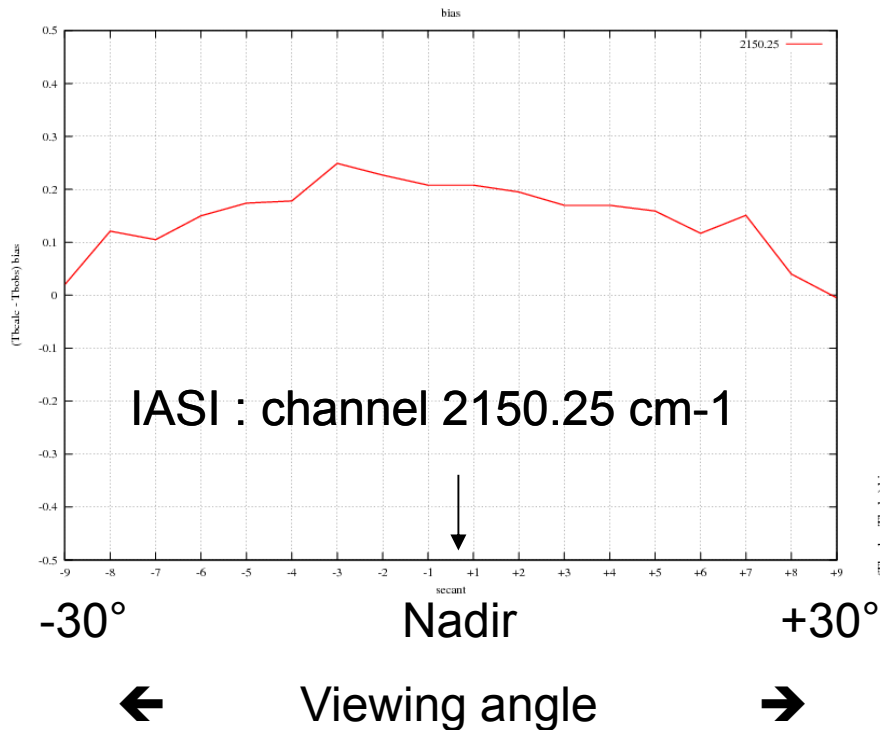


➔ *Current work : introduction of the line mixing (Hartmann and al)*



➔ *Future work : update the CO<sub>2</sub> line mixing and eventually the N2 continuum (coll. Hartmann)*

## Bias due to the instrument (viewing angle effects)



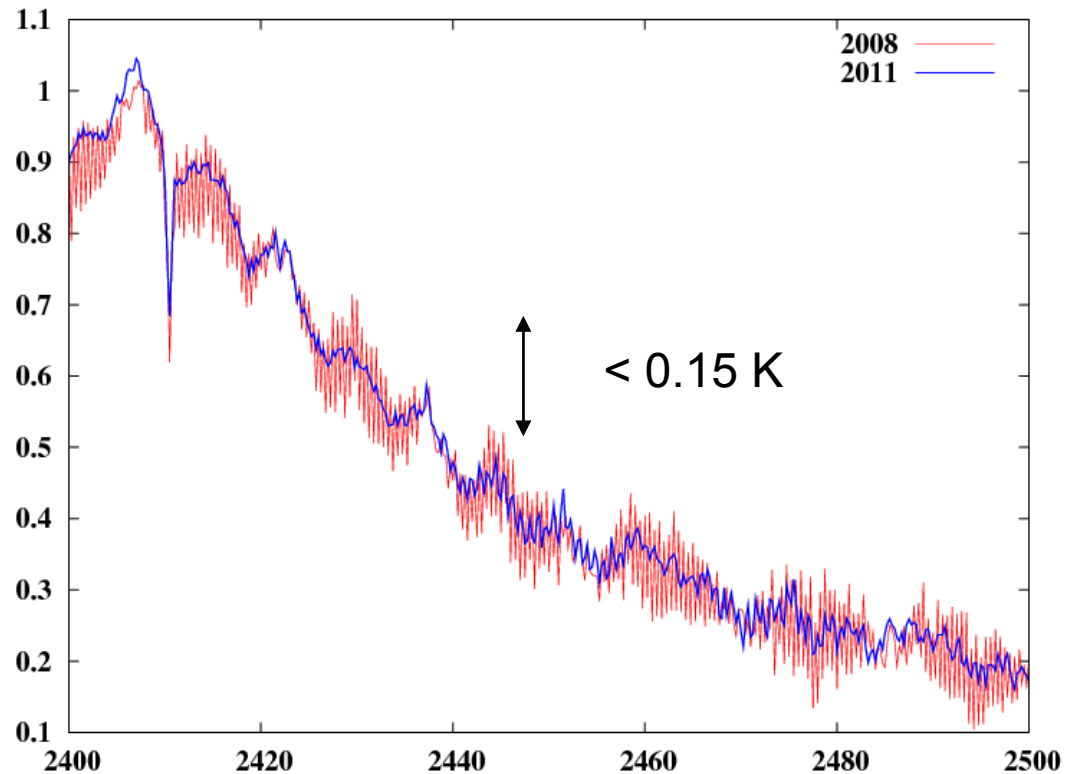
→ GSICS activities

## Bias due to the data processing

→ *Even if distance of the collocation (3h, 100 km) is important, validation with radiosoundings allow identifying small features*

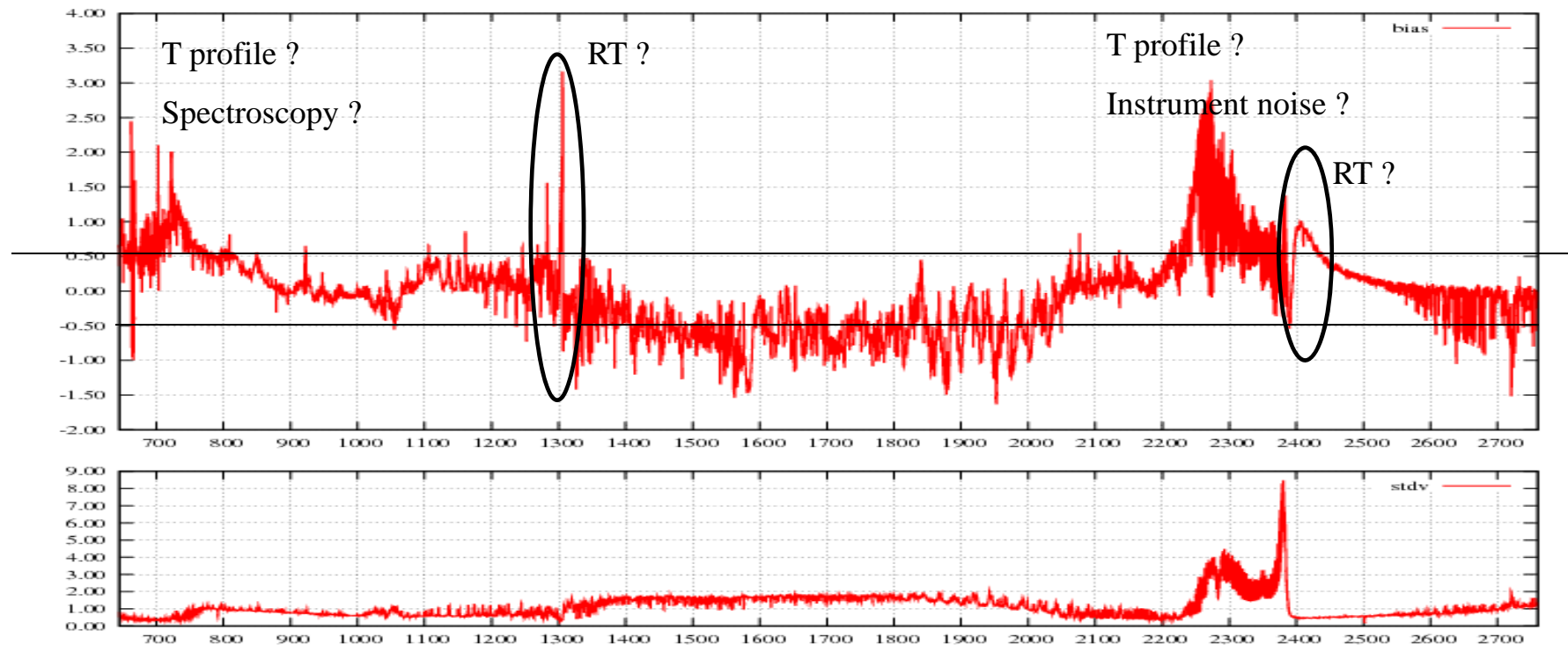
*Numerical error during the data processing due to the passage of the level1a to the level1b (Gibbs effects)*

*Corrected by the TEC (Toulouse) in 2010*



## Conclusions and future work

Biases and standard deviations are systematically investigated in order to unambiguously identify the origin of the discrepancies : natural (e.g. seasonal variations, unexpected emissions of “pollutants”, ...), spectroscopic, instrumental, thermodynamical, ...



➔ The new generation of instruments (IASI-NG, Crevoisier et al, Wednesday), it would be possible to assess the accuracy of some of the spectroscopic parameters