

## THE STATUS OF FY-3C IN NWP AND THE PREPARATION OF FY-3D AND FY-4A FOR NWP luqf@cma.gov.cn



#### **Outline**

- The evolution of FY-3 for NWP
- The status of FY3 in NWP
- The preparation of FY-3D and FY-4A for NWP

## **1. The evolution of FY-3 for NWP**



Forecast sensitivity to observations (FSO) Is an adjoint based technique for assessing the influence of observing systems on forecast accuracy

(from C. Cardinali, ECMWF)

#### The FY-3A/B/C/D/E Instrument Suites for NWP

Infrared Atmospheric Sounder (IRAS) 20 channels (~HIRS/3) HIRAS(1370channels)

Microwave Temperature Sounder (MWTS) 4 channel (~MSU) 13 channels 17 channels

Microwave Humidity Sounder (MWHS) 5 channel (~MHS) 15channels with channels at 118 GHz





Microwave Radiation Imager 10 channels (~AMSR-E)

GNSS Radio-Occultation Sounder (GNOS) (~GPS)





# The comparable data quality of FY-3C sounding instruments to its counterparts



#### The plan and status of FY-3C in NWP model from three centers..

ECMWF				UKMO					CMA NWPC			
	2014	2015	2016		2014	2015	2016			2014	2015	2016
FY-3B MWHS	Op DA	Op DA	Op DA	FY-3B MWHS	Evaluation	Monitoring	To be Op DA		FY-3B MWHS	Evaluation	Evaluation	To be Moni
FY-3C MWTS2	Evaluation	Evaluation (Now dead)	Evaluation (Now dead)	FY-3C MWTS2	Evaluation	Evaluation (Now dead)	Evaluation (Now dead)		FY-3C MWTS2	Evaluation	Evaluation (Now dead)	Evaluation (Now dead)
FY-3C MWHS2	Evaluation	Monitoring	Op DA	FY-3C MWHS2	Evaluation	Monitoring	Op DA	1	FY-3C MWHS2	Evaluation	To be Op Da	Op DA
FY-3C MWRI	Evaluation	Evaluation	To be Moni	FY-3C MWRI	Evaluation		To be Moni		FY-3C MWRI	Evaluation		To be Moni
FY-3C	Evaluation	Evaluation	Monitoring		٨٢				FY-3C			
				F 1-SC IRAS					IRAS			
				FY-3C GNOS	Evaluation	Evaluation	Evaluation	FY-3C GNOS	FY-3C	Evaluation	To be On De	0n D4
GNOS	Sample data	Sample data	Evaluation						GNOS	Evaluation	to be Up Da	UP DA

• FY3C MWHS-2 has been operationally assimilated and monitored in the Met Office global model on 15 March 2016, and in ECMWF IFS system on 4 April 2016.

• Operational assimilation of MWHS-2 with 183 GHz channels globally and GNOS in CMA/GRAPES have been activated in April 2016.

#### Monitoring OMB against the instrumental parameter to indicate the performance on orbit Server Terminal: http://satellite.nsmc.org.cn

FY3C MWHSY GLBA SM B1T 0XX SLN 20171125 LIFE MUTTS MS

FY3C\_MWHSX\_GLBA\_SM\_B1T\_0XX\_SLN\_20171125\_LIFE\_MUTTS\_MS



# Monitoring OMB against the instrumental parameter to indicate the performance on orbit

FY3C\_MWHSX\_GLBA\_SM\_NED\_A03\_OBT\_20170930\_LIFE\_ORXXX\_MS



# Monitoring OMB against the instrumental parameter to indicate the performance on orbit

FY3C\_MWHSY\_GLBA\_SM\_OMB\_103\_AVG\_20170930\_LIFE\_BSMTX\_MS



#### **Client Software**



#### **GNOS** improvement from Mi Liao and Sean Healy



 Monitored passively at ECMWF and being prepared for operational assimilation

# FY-3C Latency (GNOS) (from Mikael Rattenborg)



# FY-3C Latency (MWHS) (from Mikael Rattenborg)





## The improvement from MWRI Calibration



<sup>له</sup> [									
Frequencies (GHz)	10, 65.,	18.7.,	23.8.1	36, 5.,	89.,				
Polarization.,	V. K.,	V. K.,	¥. K.,	¥. K.,	¥. K.,				
Bandwidth (MHz) .,	180.1	2 00.1	400.,	900.1	3000.,				
Sensitivity(k).,	0, 5,,	0, 5,,	0.8.,	0, 5,,	1.0.,				
Calibration error (k) .	1.0.	2.0.	2.0.,	2.0.	2.0.,				
Dynamic Range (k)	3 = 340.,								
Sampls/scan.,	2 40.,								
Main beam efficiency.	> 90%								
Groun Resolution < (km × km)	51 × 85.,	30× 50.,	27 × 45.,	18×30.,	9 × 15.,				
Scan mode.	Conical scanning.								
Orbit width (Km)	1 400.,								
Viewing Angle (* .)	45 ± 0.1.,								
Scan period (\$)	1.7±0.1.								

Instrument Characteristics



#### original

High values of antenna emission have been observed from TMI and SSMIS.

> adjust the emissivity of hot reflector and cold reflector

adjust the emissivity of hot reflector

# The statistics of OMB



original Ascending
original Descending
adjust the emissivity of hot reflector Ascending
adjust the emissivity of hot reflector Descending
adjust the emissivity of hot reflector and cold reflector Ascending
adjust the emissivity of hot reflector and cold reflector Descending

## 3. The preparation of FY-3D and FY-4A for NWP

# Five payloads from FY-3D are of particular interest to NWP community

- MicroWave Temperature Sounder 2 (MWTS-2)
- MicroWave Humidity Sounder 2 (MWHS-2)
- High spectral Infrared Atmospheric Sounder (HIRAS)
- Global Navigation Satellite System Occultation SoundeR (GNOS);
- Microwave Radiation Imager (MWRI)

# Two payloads from FY-4A are of particular interest to NWP community

- Geostationary Interferometric Infrared Sounder (GIIRS)
- Advanced Geosynchronous Radiation Imager (AGRI)

#### HIRAS instrument specification improvement from FY-3D to FY-3E

Band	Spectral Range	Spectral Resolution	(N	Num of			
Danu	(cm <sup>-1</sup> )	(cm <sup>-1</sup> )	FY-3D	FY-3	Channels		
LWIR		0.625	0.15(Expectation) 0.4K(Requirement)	$650 \sim 667 \mathrm{cm}^{-1}$	0.8K	778	
	650~1136 (15.38µm~8.8 µm)			667~689 cm <sup>-1</sup>	0.4K		
				689~1000 cm <sup>-1</sup>	0.2K		
				$1000 \sim 1136 \mathrm{cm}^{-1}$	0.4K		
MWIR1	1210~1750 (8.26μm~5.71 μm)	1.25	0.1(Expectation)	1210~1538 cm <sup>-1</sup>	0.2K	433	
			0.7K(Requirement)	1538~1750 cm <sup>-1</sup>	0.3K		
MWIR2	2155~2550	2.5	0.3(Expectation)	2155~2300 cm <sup>-1</sup>	0.3	150	
	(4.64μm~3.92 μm)		1.2K(Requirement)	2300~2550 cm <sup>-1</sup>	0.5	139	

# **FY-3D/HIRAS TVAC**

![](_page_21_Figure_1.jpeg)

![](_page_21_Figure_2.jpeg)

FY-3D /HIRAS instrument NEdT MW1 : All channels meet NEdT specification LW&MW2: Most of channels meet NEdT specification except few of edge channels

#### **The FY-4A Instrument Suites for NWP**

Geo. Interferometric Infrared Sounder(GIIRS)(1650channels) by the Shanghai Institute of Technical Physics of the Chinese Academy of Sciences

![](_page_22_Picture_2.jpeg)

![](_page_22_Picture_3.jpeg)

![](_page_22_Figure_4.jpeg)

- The on-orbit spectral resolution for LWIR and MWIR are 0.625cm<sup>-1</sup>, better than the specified (0.8/1.6 cm<sup>-1</sup>), similar to NPP/CrIS;
- The NEDT for all the 1650 channels except some contaminated channels, generally is less than 0.1K, consistent with the specified;
- The comparisons of LWIR and MWIR with the counterpart channels from METOP-A/IASI shows that the calibration difference is about 0.64K and 0.99K separately, spectral difference is about 8ppm.

![](_page_23_Figure_4.jpeg)

Spectrum Comparison with METOP/IASI

# What we are doing for the interferometer

#### Generally, there are 4 components of energy received by the detector

![](_page_24_Figure_2.jpeg)

Instrument emission after interfere Earth scene radiance after interfere

# **Items affecting calibration precision**

Modules that have been or will be incorporated in the

ground segment algorithms:

- Interferogram alignment
- Non-linearity correction
- Self apodization correction
- Different calibration equation
- Doppler shift correction
- Polarization correction --- not been incorporated yet

# The nonlinear simulation of interferometer

1.Polynomial:

![](_page_26_Figure_2.jpeg)

## Effect of Nonlinear Correction, simulation

### **Before Correction**

### After Correction

![](_page_27_Figure_3.jpeg)

#### Polynomial

## **Effect of Nonlinear Correction, TVAC**

![](_page_28_Figure_1.jpeg)

## **Effect of Nonlinear Correction, TVAC**

![](_page_29_Figure_1.jpeg)

### **Simulation of the GIIRS**

280

275

- 270

265

260

255

250

- 310

- 300 - 290

280

- 270

260

280

- 275 - 270 - 265 - 260

> 255 250 245

![](_page_30_Figure_1.jpeg)

date: 20160801~20160823

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### **Simulation of the GIIRS**

The comparison of CRIS and GIIRS

![](_page_31_Figure_2.jpeg)

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![](_page_32_Figure_0.jpeg)

#### Skew T-InP diagram over Fangshan, Beijing during 00:00 UTC-12:00 UTC (08:00-20:00 LST) 02 Aug 2017

#### With LMI+AGRI+GIIRS, what can we see?

![](_page_32_Figure_3.jpeg)

À PE

**Convective Available Potential Energy** (CAPE) map during 00:00 UTC-12:00 UTC (08:00-20:00 LST) 02 Aug 2017 (grey to white areas represent cloud observed by FengYun-4 satellite; the asterisk denotes the location of **Fangshan of Beijing**)

## **OSSE: Preparing FY-3D and FY-4A for NWP**

#### Improve the data precision and stability

- Monitor the OMB and instrumental parameters to indicate the data quality
- Characterize the instrumental biases
- Control the data quality

#### Support the earlier preparation of data assimilation

Generate the initial coefs of the fast radiative transfer modeling for NWP data assimilation

Simulate the sample data by RTM

Release the sample data to cooperative users (after the agreed coordination of CMA and WMO)

Prepare the assimilation in NWP model

#### **Optimize the performance**

Evaluate/improve the data quality and its impact on NWP model