

NSMC



# Performance Status of FY-3E/HIRAS-II and FY-4B/GIIRS

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Virtual Meeting

Thursday, March 24, 2022

# Outline

- FY-3E/HIRAS-II Introduction
- FY-3E/HIRAS-II On-board Performance Status
- FY-4B/GIIRS Introduction
- FY-4B/GIIRS On-board Performance Status
- Summary

# FY-3E/HIRAS-II Introduction

#### The first IR sounder on the early-morning-orbit

### HIRAS



- > FY-3E was launched successfully on July 5, 2021.
- FY-3E is the world's first early-morning-orbit meteorological satellite for civil use.
- HIRAS-II is the second hyperspectral IR sounder onboard FY-3 series satellites.



# FY-3E/HIRAS-II Introduction

### Specifications



Band	Spectral Range (cm <sup>-1</sup> )	Spectral Resolution (cm <sup>-1</sup> )	Sensitivity (NE∆T@280K)			Radiometric Accuracy		Spectral Uncertainty	
			FY-3D	FY-3E		FY-3D	FY-3E	FY-3D	FY-3E
LWIR	650~1168.125 (15.38 μm~8.56 μm)	0.625	0.4 K	$650 \sim 667  \mathrm{cm}^{-1}$	0.8 K	0.7 K	1.0 K	10 ppm	7 ppm
				$667 \sim 689 \text{ cm}^{-1}$	0.4 K		0.5 K		
				$689 \sim 1000 \mathrm{cm}^{-1}$	0.2 K		0.4 K		
				$1000 \sim 1136 \mathrm{cm}^{-1}$	0.4 K		0.5 K		
MWIR	$\frac{1168.75{\sim}1920}{(8.55\ \mu\text{m}{\sim}5.21\ \mu\text{m})}$	0.625	0.7K	$1210 \sim 1538 \text{ cm}^{-1}$	0.2 K	0.7 K	0.4 K		
				$1538 \sim 1750 \text{ cm}^{-1}$	0.3 K		0.5 K		
SWIR	1920.625~2550 (5.21 μm~3.92 μm)	0.625	1.2 K	2155~2300 cm <sup>-1</sup>	0.3 K	0.7 K	0.5 K		
				$2300 \sim 2550 \mathrm{cm}^{-1}$	0.5 K		0.6 K		

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# HIRAS-II NEdT Performance

#### NEdT@280 K inter-comparison with several IR sounders



FY-3E/HIRAS-II LWIR and MWIR have good noise performance

SWIR FOV-1 is out of family and larger than the specification

FY-3E/HIRAS-II noise performance is better than FY-3D/HIRAS except for the FOV-1 in SWIR.

# HIRAS-II NEdT Performance

### **NEdT Stability**





FY-3E/HIRAS-II mean NEdT of 9 FOVs (SW FOV-1 being rejected)

FY-3E/HIRAS-II noise performance is stable.

# **HIRAS-II** Instrument Responsivity

#### Long-Term Monitoring of Instrument Responsivity in Contaminated Channels wrt. that on 2018/04/21 1.10 1.05 1.00 0.95 0.90 0.85 0.80 8 0.70 × 0.65 0.60 a 0.55 0.50 0.45 **9**0.40 0.35 0.30 0.25 0.20 0.15 LIR InsResp @801.2cm 0.10 MIR InsResp @1259.29cn SIR InsResn @2339.2cm 0.05 0.00 020111 021101 021103 021106

#### LW InsResp 2018.12.26 2018 Apr.21 12:00 1 8 2019 Jun.20 12:00 0.7 2019 Jun.21 12:00 2019 Jun.22 12:00 1.6 2019 Jun.23 12:00 0.6 2019 Jun.24 12:00 1.4 0.5 n (1.2 (9. n.2) SRF (a. 8'0 ISRF ( 0.3 0.6 0.2 0. 0.1 0.2 2018.12.20 500 600 700 800 900 1000 1100 1200 1300 1100 Wavenumber (cm<sup>-1</sup>)

InsResp Stability



### FY-3D/HIRAS InsResp Monitoring

InsResp
$$(\sigma) = \frac{\left|C_{ICT}(\sigma) - C_{DS}(\sigma)\right|}{B_{ICT}(\sigma)}$$

 $C_{\text{ICT}}$ : HIRAS ICT FFT raw spectra  $C_{\rm DS}$ : HIRAS ICT FFT raw spectra  $B_{\rm ICT}$ : ICT radiance on sensor resolution In 2018, it had been found that the instrument responsivity (InsResp), and also the measured signal in each band of FY-3D/HIRAS would decrease with time. The signal attenuation could attribute to the silica gel gas absorption that appears in the interferometer optical path. The absorption affects almost the whole spectral range of the three bands. There are some strong absorbing channels in LW and MW bands, such as 801.2 cm<sup>-1</sup> and 1259.3 cm<sup>-1</sup>. From 2018 to 2021, HIRAS has undergone three on-orbit operations of warming maintenance in 2018-12-21, 2019-10-30, and 2021-03-21, respectively.



#### Ref:

Han Y., Chen Y. 2018, IEEE Trans. Geosci. Remote Sens. 56(2):1008-1016. Qi C., Wu C., Hu X. et al. 2020, IEEE Trans. Geosci. Remote Sens. 58(6):4335-4352.

# **HIRAS-II** Instrument Responsivity

#### InsResp Stability





FY-3E/HIRAS-II mean InsResp of 9 FOVs

Up to now, FY-3E/HIRAS-II instrument responsivity is stable.



The spectral offsets for all three bands are within  $\pm 5$  ppm after the spectral calibration parameter updated on Dec. 20, 2021.



### based on SNO IASI-C spectra



### based on SNO IASI-C spectra





MWIR spectral offsets are within ±5 ppm

#### based on SNO inter-comparison with IASI-B



HIRAS-II to IASI-B SNOs in the polar regions (from Nov.14 to Dec.01 2021)

The BT differences in LWIR and MWIR are less than 1.0 K (std. dev. < 1K).

The SWIR BT differences are larger than that in LWIR and MWIR, because of the cold scenes in Arctic regions.





1000 1400 1800

wavenumber/cm<sup>-1</sup>

2200 260

wavenumber/cm<sup>-1</sup>

Bias/K

Bias/K

Bias/K

### Based on SNO inter-comparison with IASI-B



Bias/K Bias/K -10-101800 2200 2600 600 1000 1400 600 wavenumber/cm<sup>-1</sup> fov4-south:22 Bias/K Bias/K -101000 1800 600 1400 2200 600 wavenumber/cm<sup>-1</sup> fov7-south:21 10 Bias/K Bias/K

1400 1800 2200 2600

wavenumber/cm<sup>-1</sup>

-10

600 1000

1400 1800 2200 2600

wavenumber/cm<sup>-1</sup>

fov1-south:33



-10

600

1000 1400 1800

wavenumber/cm<sup>-1</sup>

2200 2600

HIRAS-II to IASI-B SNOs in the polar regions (from Nov.14 to Dec.01 2021)

Since the SNO scenes in Antarctic are warmer than Arctic, the SWIR BT dev. has been reduced considerably.

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-1(

600 1000

### based on IASI SNO comparison & double-difference



The RTTOV spectra are calculated for clear scenes using ERA5.

LWIR radiometric accuracy is better than 0.5 K, and MWIR is about 0.5 to 1.0 K.

SWIR is larger than specification, and is need improved.







LWIR radiometric accuracy is better than 0.5 K, MWIR is about 0.5 to 1.0 K.

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#### Ref:

D. Pierre, T. Bertrand, C. Dorothee, et. al. Introduction to the ringing effect in satellite hyperspectral atmospheric spectrometry. AMT, 2021, (discussions). D. Pierre, T. Bertrand, C. Dorothee, et. al. Correction of calibration ringing in the context of the MTG-S IRS instruments. arXiv, 2022.

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# **FY-4B/GIIRS Introduction**

#### 遮光罩 南北扫描机构 吊点 东西扫描机构 坐标系 冨冷支撑 光阑 望远镜系统 主箱体 涉仪 汝热板 辐冷 激光器电控箱 **GIIRS Model**

FY-4B/GIIRS was launched successfully on June 3, 2021, which is China's second hyperspectral IR sounder on the GEO satellite.

Introduction

Unlike the demonstration of FY-4A/GIIRS, this sounder is expected to be used for NWP operational model.



- The FY-4B/GIIRS L1 science team was set up in October 2020, with two HIRAS scientists joining in. The two scientists are Chengli Qi and Lu "Richard" Lee.
- The new team redesigned the L1 calibration algorithm which is based on Revercomb et al (1988).

# FY-4B/GIIRS Introduction

### Specifications



The detector layout has been updated from 32×4 to 16×8, with spatial sampling from 16×16 km to 12×12 km at Nadir.

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# FY-4B/GIIRS Day-1 data check

6000

4000

IGM [counts]

-2000

-4000



### 

#### based on LBL simulated spectra



#### based on LBL simulated spectra



### based on SNOs spectra with IASI-C





SNOs data 2022.02.22-24

FY-4B/GIIRS LWIR spectral calibration accuracy : most FOVs within ±7 ppm (std. < 6ppm)</p>

#### based on SNOs spectra with IASI-C







Just preliminaries, fine tuning is still under going.

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### **GIIRS Brightness Temperature**



BT images in some channels

LW in 900 cm<sup>-1</sup> channel





MW in 1870 cm<sup>-1</sup> channel

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### Summary

### > FY-3E/HIRAS-II

- The spectral and radiometric calibration accuracy and stability need to be improved
- The long-term monitoring of sensor sensitivity and spectral responsivity (InsResp) are under going
- FY-4B/GIIRS
  - The radiometric accuracy will be improved by adjusting the algorithm (such as LW nonlinearity correction)
  - Spectral offset stability requires long-term monitoring (MW spectral calibration parameters need fine tuned)
  - The long-term monitoring of sensor sensitivity and spectral responsivity (InsResp) are under going

# Thank you for your time !

Fin

If you have any queries, please contact us at the address below.

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