



# IASI-NG Level 1 processing

## How to estimate the instrument spectral response function in real-time ?



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**<sup>1</sup> CNES, <sup>2</sup> Noveltis**



# IASI – NG level 1 processing

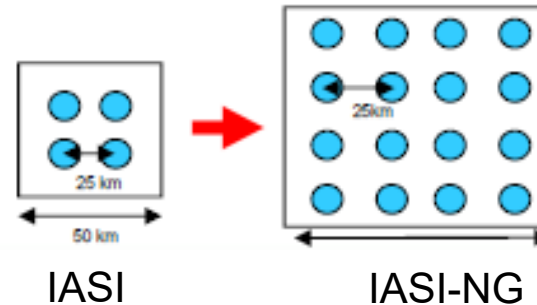


- ❖ From IASI to IASI-NG: innovations and new challenges
- ❖ Overview of the IASI-NG Level1 processing
- ❖ The ISRF – Estimation Model
- ❖ System performance budget



## IASI-NG Level 1 main characteristics

| Main figures                             | IASI                 | IASI-NG                             |
|--|----------------------|-------------------------------------|
| Radiometric Resolution (NeDT)            |                      | IASI/2                              |
| Spectral resolution                      | 0.5 cm <sup>-1</sup> | IASI/2 (0.25 cm <sup>-1</sup> @L1C) |
| Absolute Radiometric Calibration         | < 0,5K               | IASI/2 (<0,25K@280K)                |
| Spectral bands                           | 3 bands              | 4 bands                             |
| Number of sounder pixels per acquisition | 4 pixels             | 16 pixels                           |
| Ground Pixel diameter                    | 12 km                | 12 km                               |
| Ground sampling                          | 25 km                | 25 km                               |



- IASI vs IASI-NG ?

→ With great performance comes great complexity

## IASI interferometer

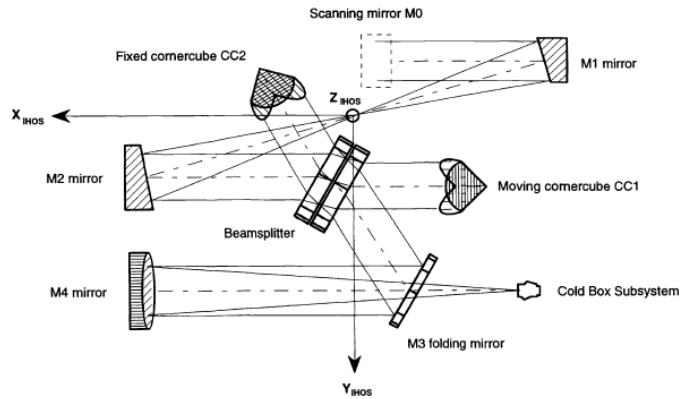
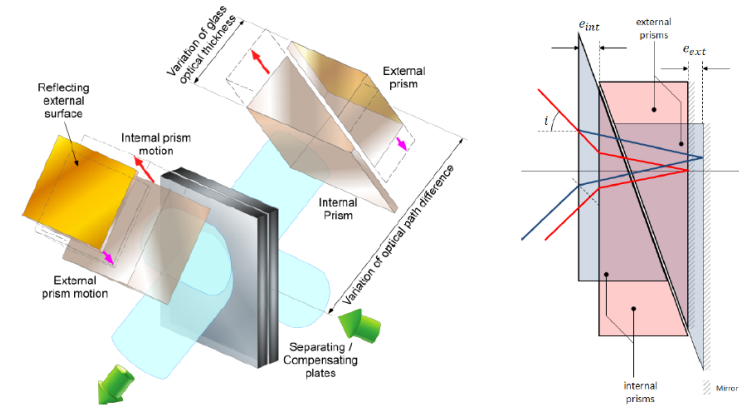


Figure 3.1-1 : IHOS optical configuration

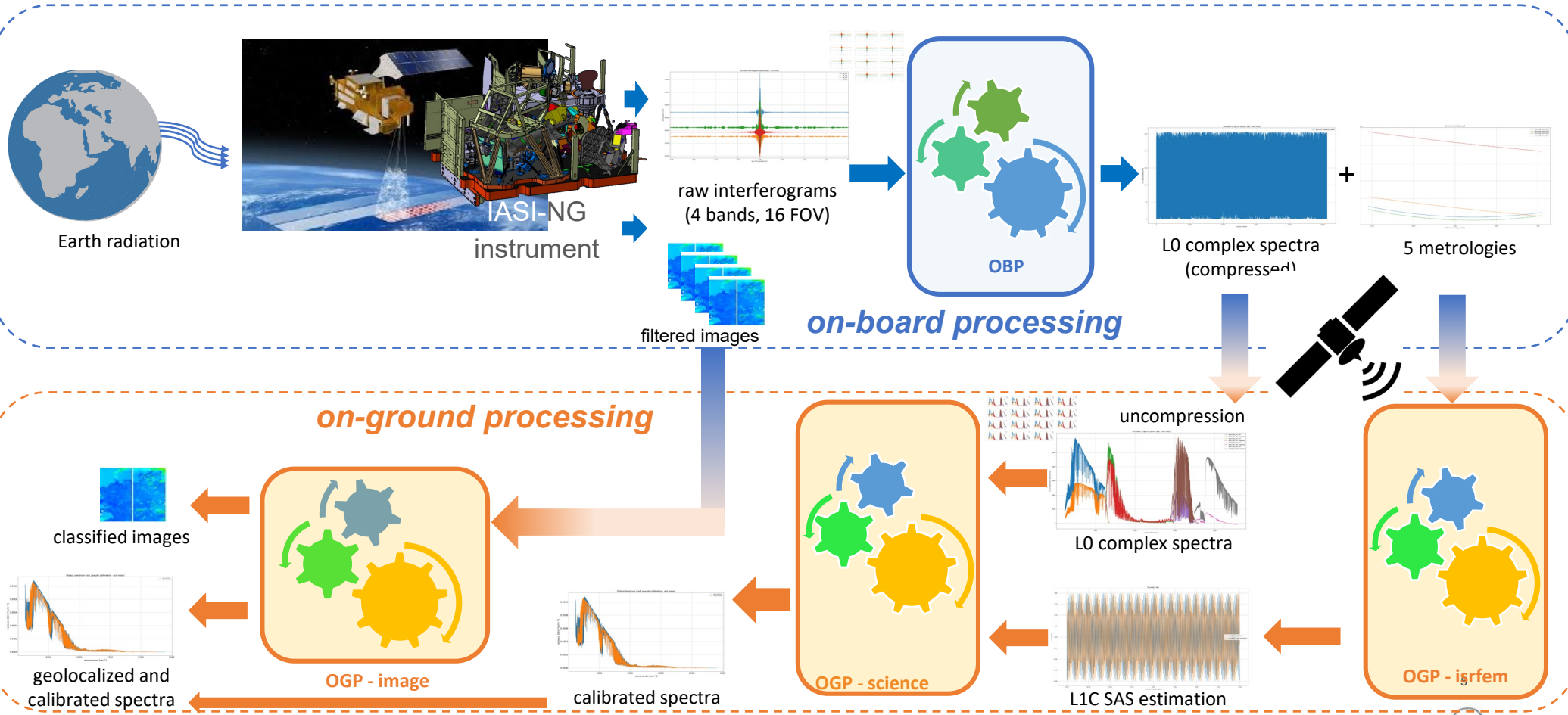
- 4 optical components
- 1 moving corner cube
- nearly achromatic / all reflective design
- 1 laser trigs the acquisition

## IASI-NG interferometer



- 6 optical components
- 4 moving prisms, simultaneously
- chromatic and refractive design
- 1 laser trigs the acquisition, 4 additionnal lasers for opd estimation

# IASI-NG Level1 chain: overview





- IASI-NG ISRF estimation principle

➤ ISRF is estimated for every set of 16x4 acquisitions

estimating the ISRF = knowing the science opd

Airbus D&S shows that the opd for every wavelength in the science band can be approximated by a linear combination of the 5 metrologies opd :

$$\text{opd}(\sigma) = (\mathbf{A}_{\text{cal}}(\sigma) - \mathbf{C}_0(\sigma, \mathbf{Z}_0)) \cdot \mathbf{Z}_0 + \sum_{i=1}^4 \mathbf{C}_i(\sigma, \mathbf{Z}_0) \cdot \mathbf{Z}_i + \text{OFFSET}(\sigma, \mathbf{Z}_0)$$

$\mathbf{Z}_{i=1..4}$  : metrologies opd

$\mathbf{Z}_0$  : reference metrology opd

$\mathbf{C}_{i=0..5}$  : opd coefficients

- IASI-NG ISRF estimation principle

## ➤ ISRF-Estimation Model parameters - Coefficients

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They are a combination of :



### Computed coefficients

- Using a numerical model of the interferometer
- Parameters of the model can be updated in-orbit using WFS

- IASI-NG ISRF estimation principle

- ISRF-Estimation Model parameters - Coefficients

$$\text{opd}(\sigma) = (\mathbf{A}_{\text{cal}}(\sigma) - \mathbf{C}_0(\sigma, \mathbf{Z}_0)) \cdot \mathbf{Z}_0 + \sum_{i=1}^4 \mathbf{C}_i(\sigma, \mathbf{Z}_0) \cdot \mathbf{Z}_i + \text{OFFSET}(\sigma, \mathbf{Z}_0)$$

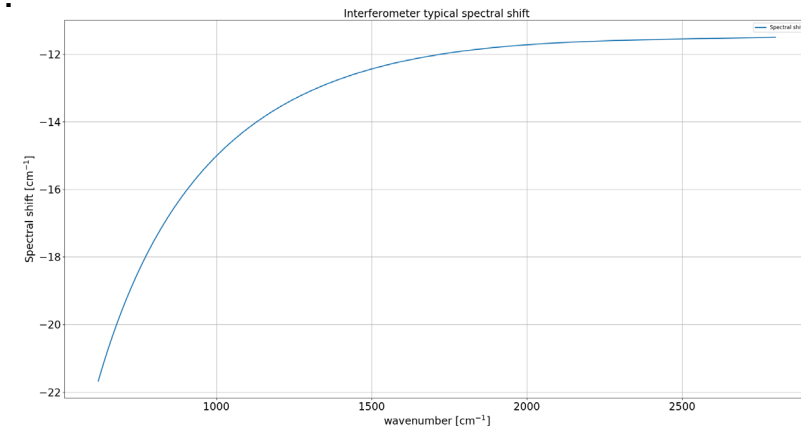
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They are a combination of :



### Measured coefficient

- On-ground spectral shift (first guess)
- In-orbit determination using dedicated acquisitions sequences (atmospheric spectra correlation + FPI)



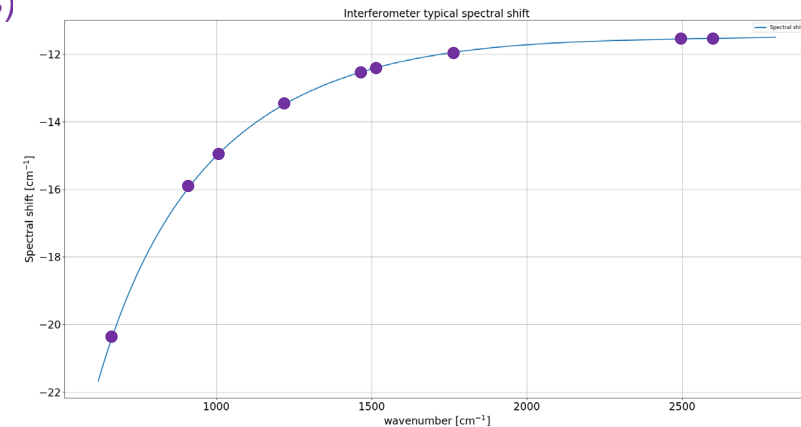
- IASI-NG ISRF estimation principle

- ISRF-Estimation Model parameters – Coefficients –  $A_{\text{cal}}$

The  $A_{\text{cal}}$  coefficient is related to the spectral shift  $\Delta\sigma$  through the simple relation :

$$A_{\text{cal}}(\sigma) = \frac{\Delta\sigma}{\sigma}$$

1-  $A_{\text{cal}}(\sigma)$  values are computed for few wavenumbers (anchor points) across the IASI-NG band using correlation between observed and modeled spectra.



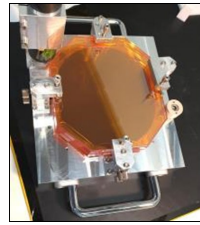
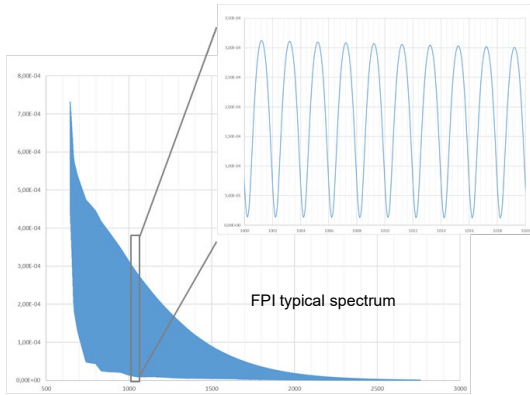


- IASI-NG ISRF estimation principle

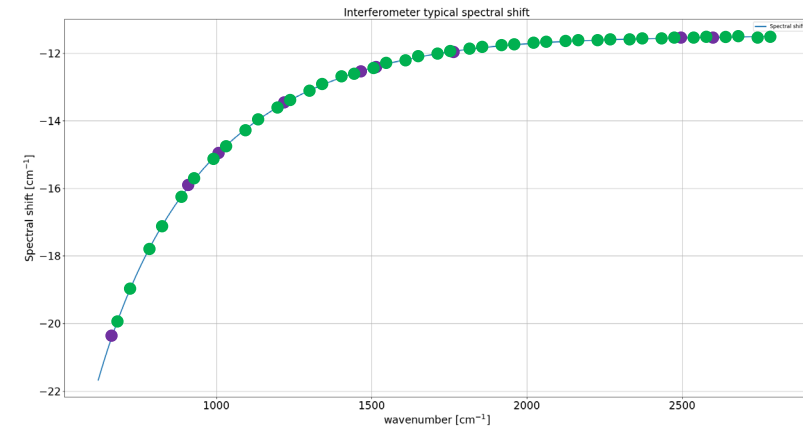
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The  $A_{cal}$  coefficient is related to the spectral shift  $\Delta\sigma$  through the simple relation :

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FPI plates (ZnSe)  
Winlight optics



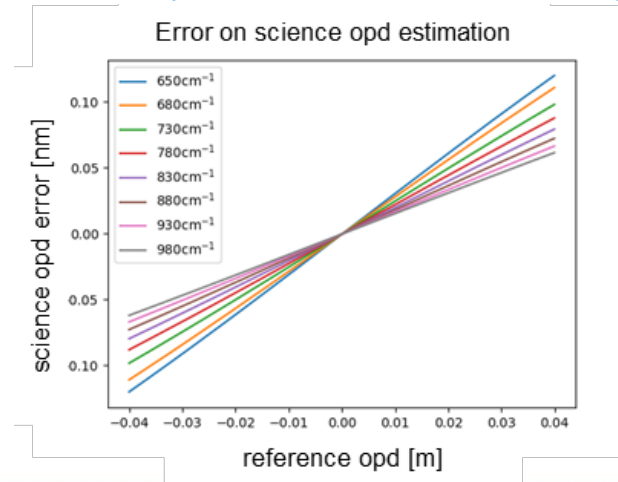
How on board calibrations will allow high level of performance during IASI-NG Mission  
E. Baldit and al – Poster session – Joint EUMETSAT/AMS/NOAA conf. 2019 (Boston)

- IASI-NG ISRF estimation performances

➤ Preliminary assessments of ISRF estimation model are very promising :

**Knowing the science opd = estimating the ISRF**

Without defects, the opd estimation is better than few tenth of nanometers  
(~ 1/1000 OPD sampling)



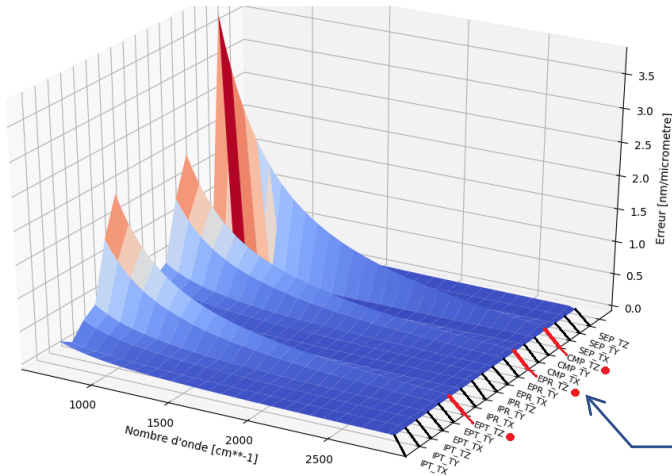
- IASI-NG ISRF estimation performances

➤ Preliminary assessments of ISRF estimation model are very promising :

Knowing the science opd = estimating the ISRF

opd variations induced by kinematical defects are also well corrected

Sensitivity of the opd error – translation defect

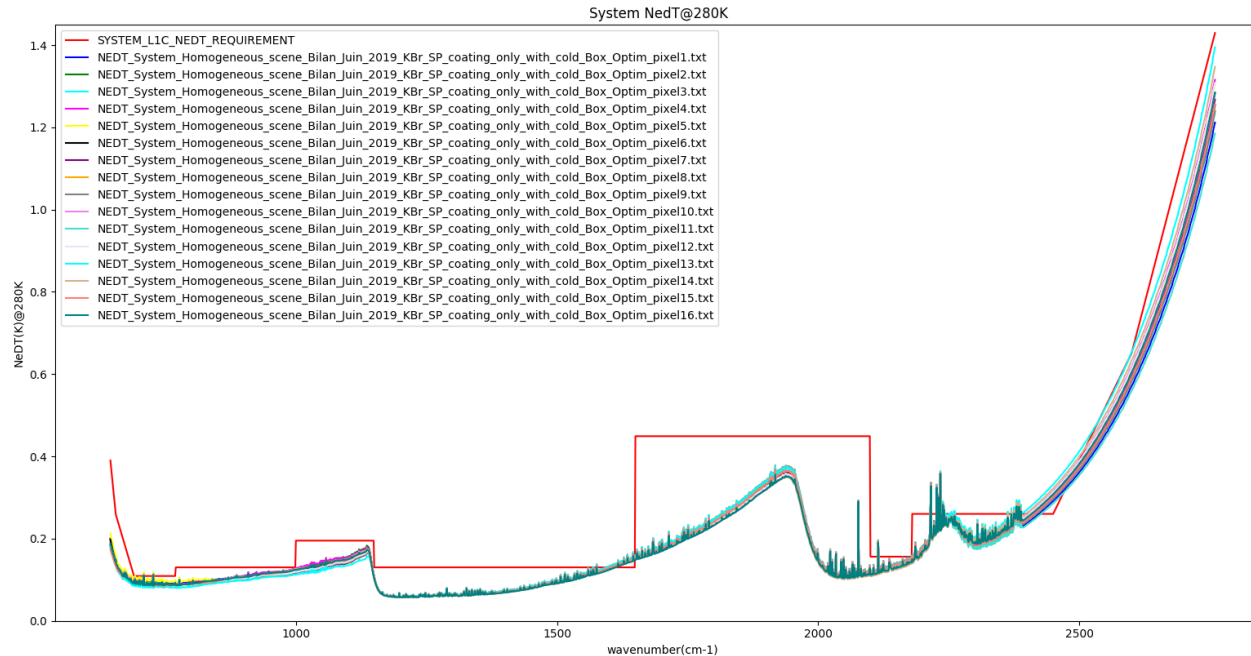


Sensitivity of the opd error below :

- **3nm/μm of translation** for every component / every wavenumbers
- Movement of 1μm of the compensating plate corresponds to a worst case

6 optical components x 3 translations / rotations

- The current budget of system performances shows very encouraging results in terms of
  - ✓ Geometric performances
  - ✓ Spectral performances
  - ✓ Radiometric performances





**Thank you ...**

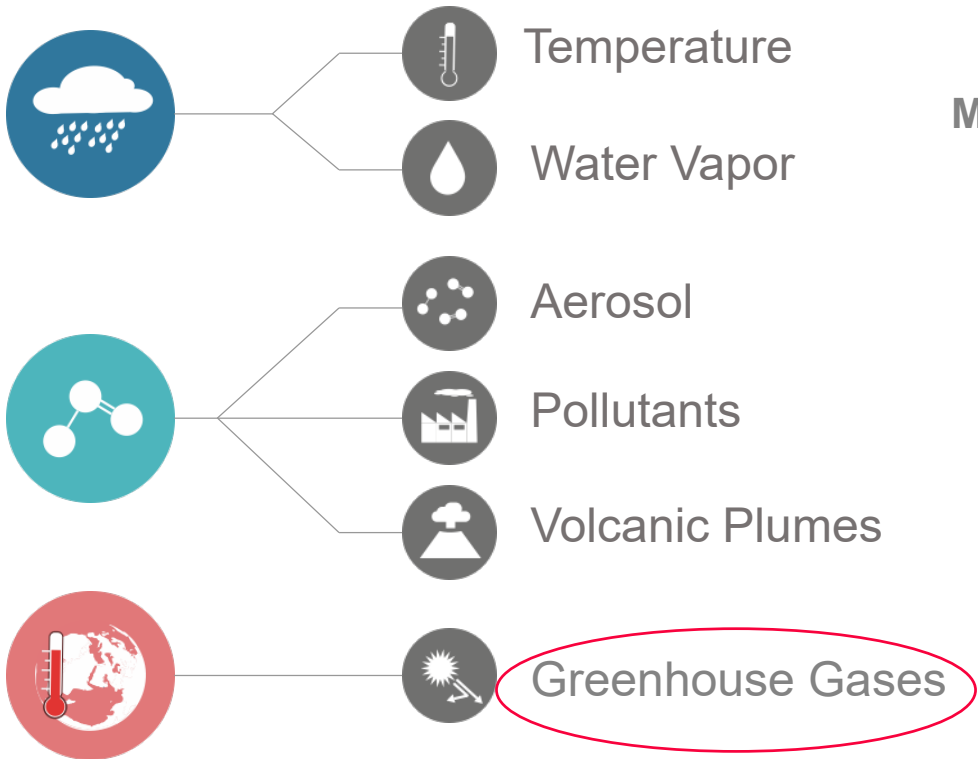
# 5<sup>th</sup> IASI Conference, 20-24 April 2020, Evian (France)



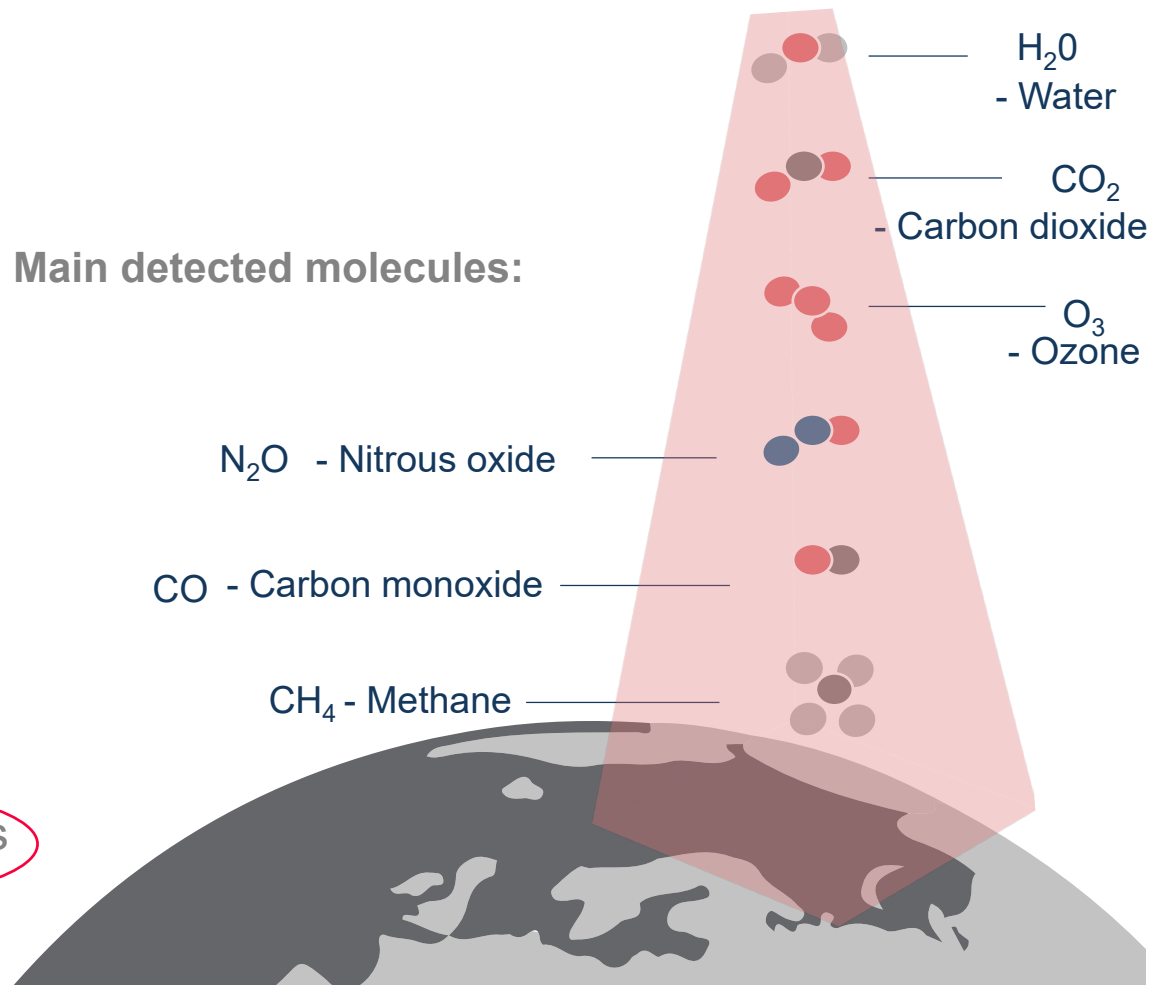
## **Back-up slides**

# From IASI to IASI-NG: innovations and new challenges

Applications are the same:

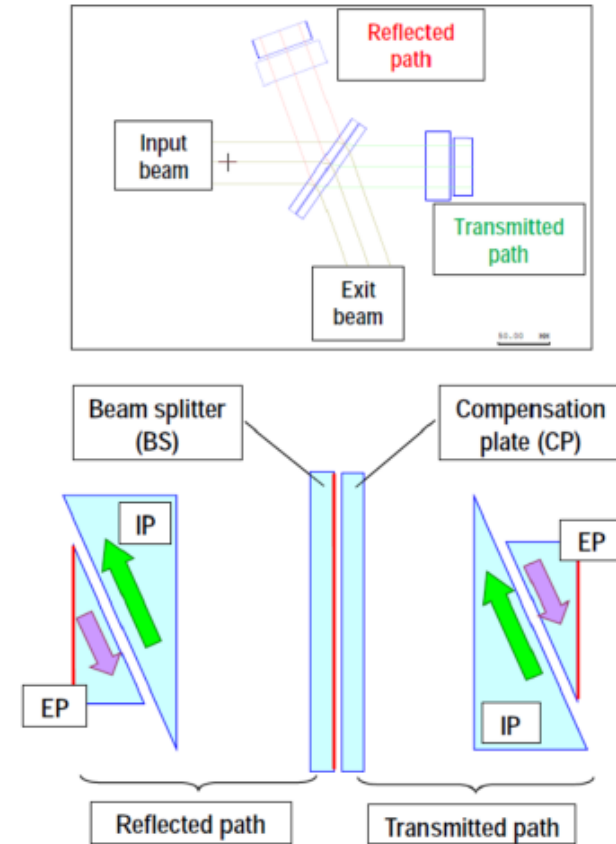


Main detected molecules:





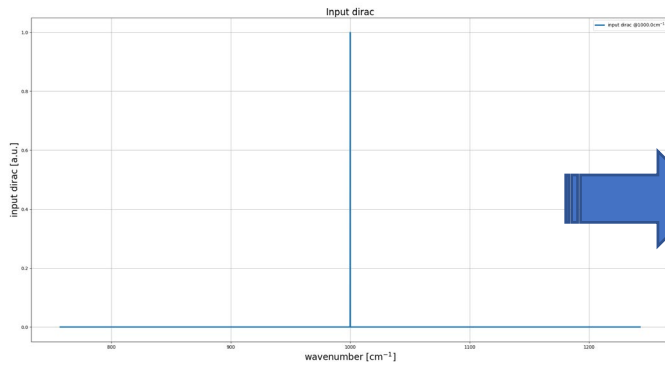
- **To deal with the stronger requirements in terms of performances, a new instrumental concept has been proposed:**
  - ✓ The Mertz interferometer allows a field compensation (self-apodisation correction)
  - ✓ Field compensation is achieved by introducing optics with correct optical index
  - ✓ A single ‘dual swing’ mechanism translates two pairs of prisms proportionally and creates simultaneously the OPD change and the self-apodisation compensation
- ✓ **The level 1 processing has been modified consequently, especially to estimate the Instrument Spectral Response Function (ISRF)**



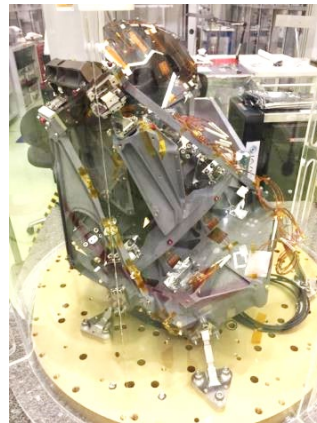
- Why estimate the instrument spectral response function (ISRF) ?

→ The impact of the instrument on measured spectra

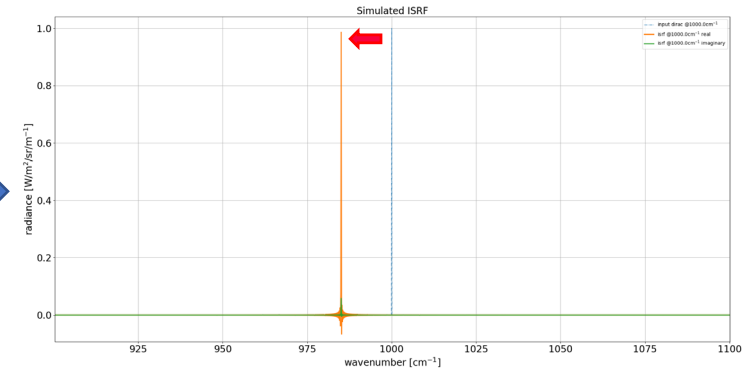
ISRF = response of the instrument to a spectral dirac



input dirac



FM1 IFM Model  
(courtesy Airbus)

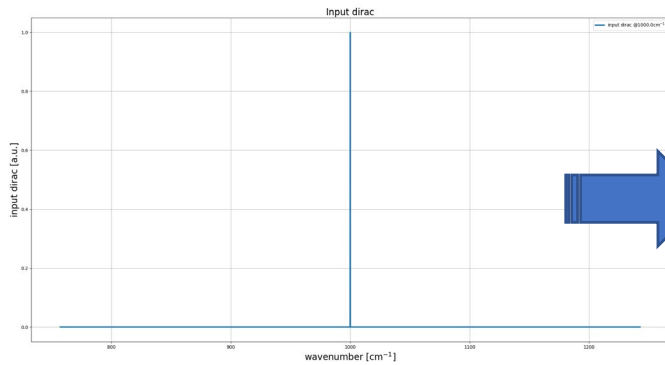


ISRF

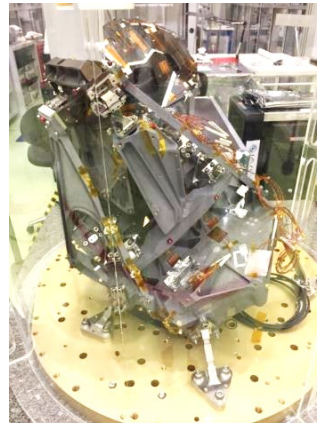
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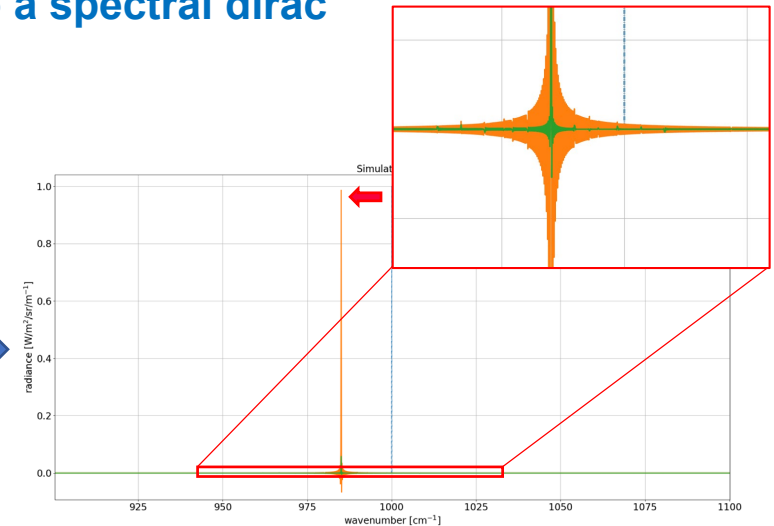
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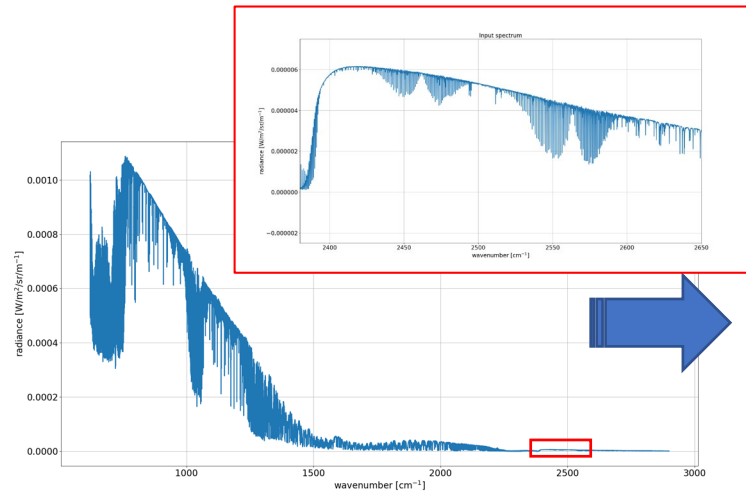


ISRF

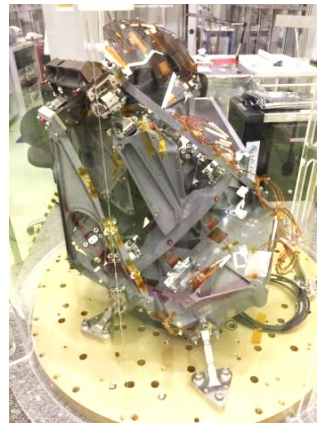
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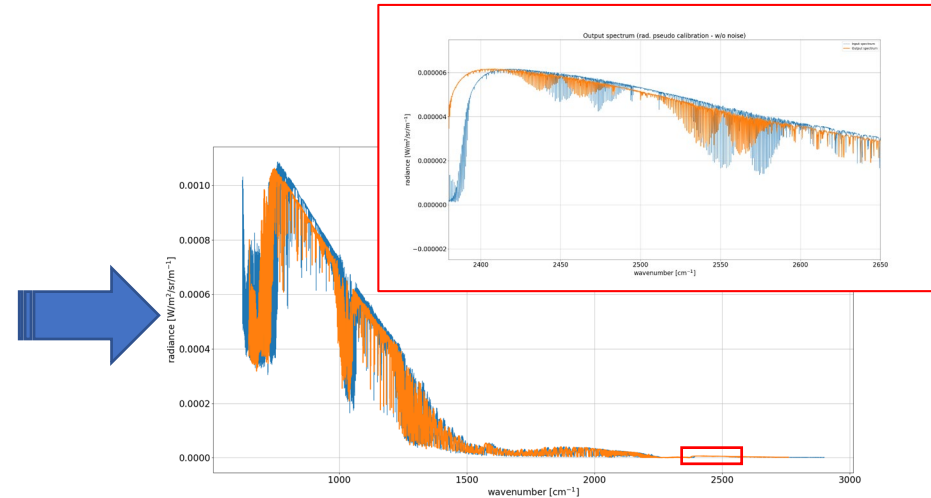
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input spectrum



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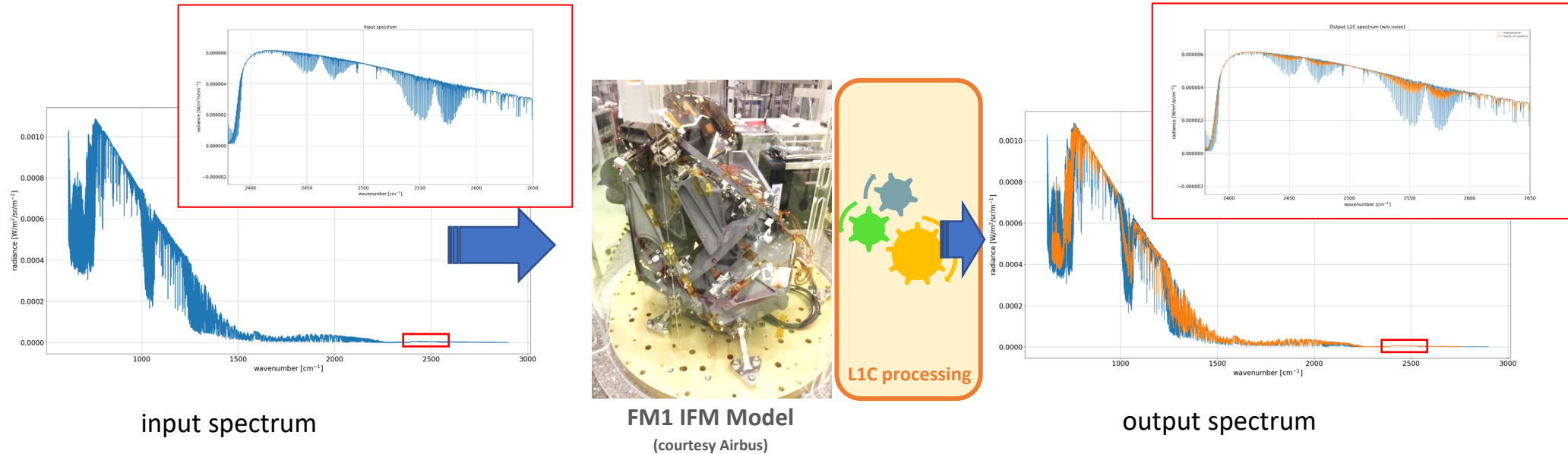


output spectrum

- Why estimate the instrument spectral response function (ISRF) ?

→ The impact of the instrument on measured spectra

ISRF = response of the instrument to a spectral dirac

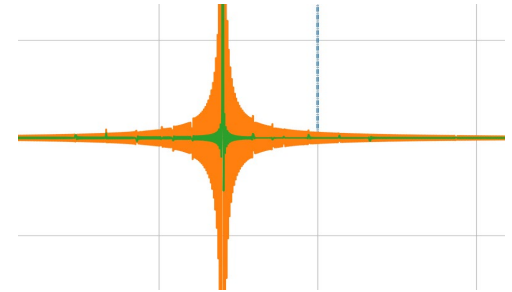


- IASI-NG ISRF defects

→ Instrument impact on ISRF can be separated into 2 main categories :

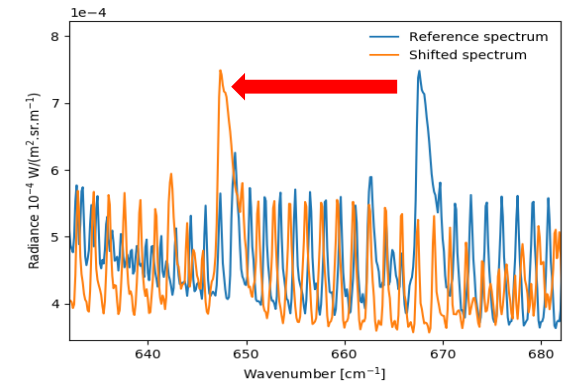
- Defects induce by the gap between the realized instrument and the ideal ones :

- optics realization (MSE,...),
- alignments,
- cinematic perturbations during the stroke, ...
- ...



- Defects induce even if the instrument were ideal :

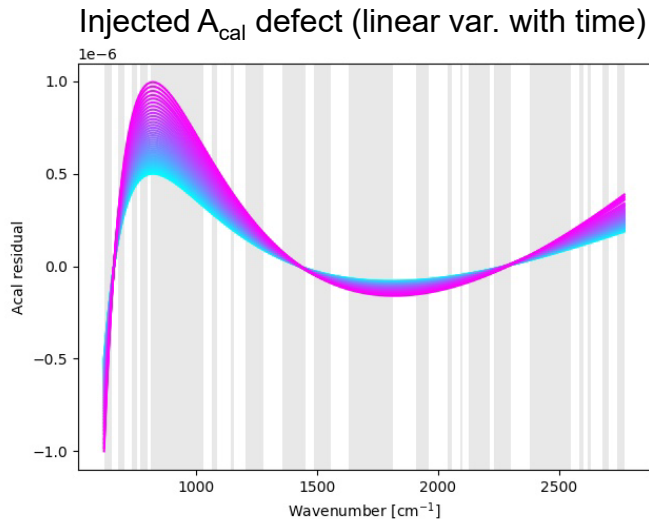
- variable spectral shift due to the opd chromaticity  
( $\sim 20\text{cm}^{-1}$ @645 and  $\sim 12\text{cm}^{-1}$ @2760 $\text{cm}^{-1}$ )



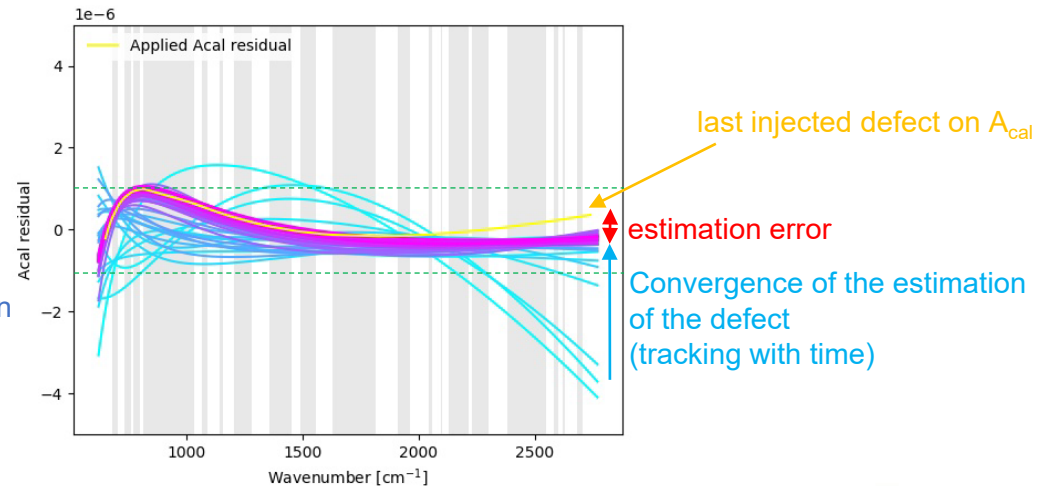
- IASI-NG ISRF estimation performances

➤ Preliminary assessments of ISRF estimation model are very promising :

spectral shift estimation is below the  $1.10^{-6}$  requirements when considering around 40 tropical (night/sea) atmospheric spectra (correlated with ref. mean tropical atmos.)



➔  $A_{cal}$  estimation by correlation



- IASI-NG ISRF estimation principle

- ISRF-Estimation Model parameters - Coefficients

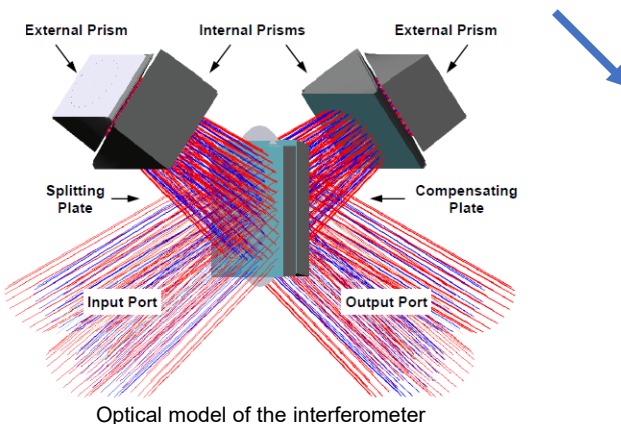
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$\mathbf{Z}_{i=0..4}$  : metrologies opd

$\mathbf{Z}_5 = \mathbf{1}$

$\mathbf{C}_{i=0..5}$  : opd coefficients

They are a combination of :



Measured coefficient :

Are expected to be updated a few times during the instrument in-orbit lifetime

dedicated acquisitions sequences (atmospheric spectra correlation + FPI)

Computed coefficients

- Using a numerical model of the interferometer
- Parameters of the model can be updated in-orbit using WFS



- **Synthesis of the current performances budget (compared to specifications at mission level) for the main requirements**

| Description                      | Specified value / calculated value     | Status                                       |
|----------------------------------|--|--|
| ISRF shift knowledge             | 1E10-6 / 1E10-6                        | compliant                                    |
| ISRF shape error index           | 0.25% / 0.27% (parasitic contribution) | Marginally NC                                |
| Interpixel ISRF stability        | 1E10-4 / 3.3E10-5                      | compliant                                    |
| Radiometric noise                | See previous figure                    | compliant                                    |
| Absolute radiometric calibration | 0.25 K (O) and 0.5 K (T) / 0.24K       | compliant                                    |
| PSF uniformity knowledge         | +/- 10 % / 14%                         | compliant                                    |
| PSF characterization             | -                                      | No budget available                          |
| Spatial resolution               | 11.5 +/- 0.5 km / 11.8km               | compliant                                    |
| Ground sampling                  | 3mrad / 3.2mrad                        | NC in worst case (compliant in typical case) |
| Pointing knowledge               | +/- 3 mrad / 0.9mrad                   | compliant                                    |
| Pointing accuracy                | +/- 3 mrad / 0.91mrad                  | compliant                                    |
| Pointing stability               | 0.1 mrad / 0.062 mrad                  | compliant                                    |
| Sounder geolocation              | 1km / 949 m                            | compliant                                    |
| Geolocation in degraded case     | 5km / 1108 m                           | compliant                                    |