



Impact of dust aerosols on the retrieval of IR land surface emissivity spectrum:

A new simultaneous approach accounting for dust characteristics and surface temperature from IASI

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Rationale

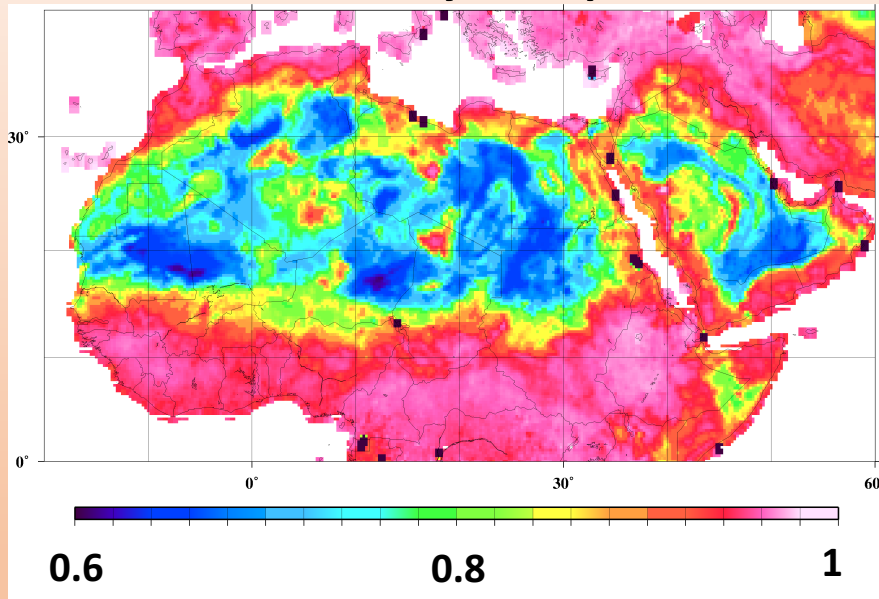
- Create a new IR surface emissivity database from IASI
- Higher spatial resolution than previous ($0.5^\circ \Rightarrow 0.25^\circ$)
- High spectral resolution (0.05 μm from 3.7 to 14 μm)
- Clean from dust contamination
- Day and night separate
- Global
- Monthly

First step: creation of a monthly climatology averaging the 10 years of IASI data

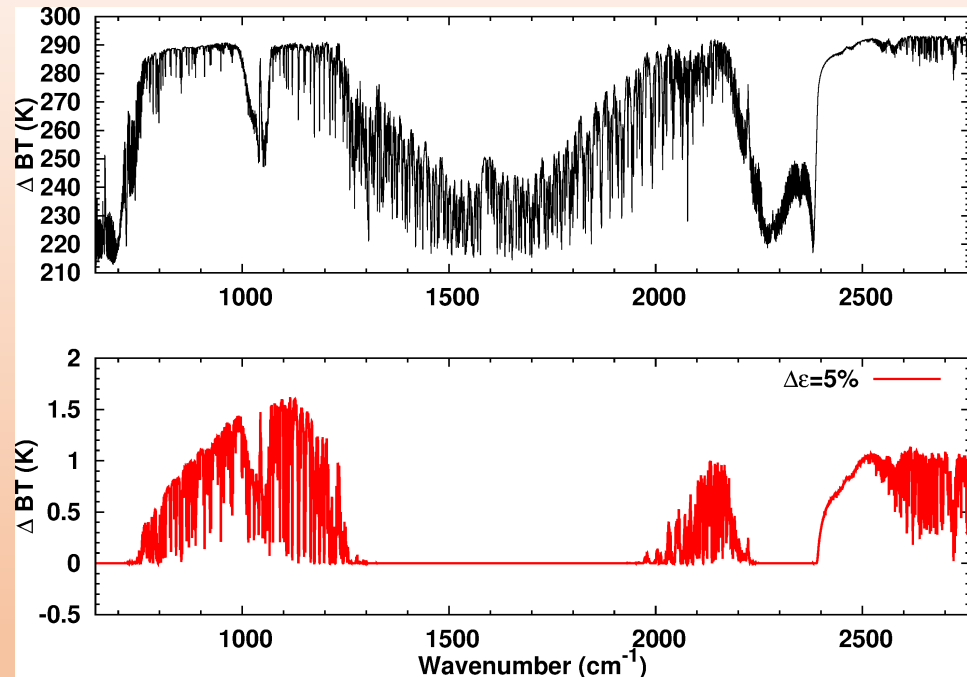
- Create a new IR surface emissivity database from IASI
- **Higher spatial resolution than previous ($0.5^\circ \Rightarrow 0.25^\circ$)**
- **High spectral resolution ($0.05\mu\text{m}$ from 3.7 to $14\mu\text{m}$)**
- **Clean from dust contamination**
- **Day and night separate**
- **Global**
- Monthly

Impact of surface parameters on BT spectrum

January – $8.3\mu\text{m}$

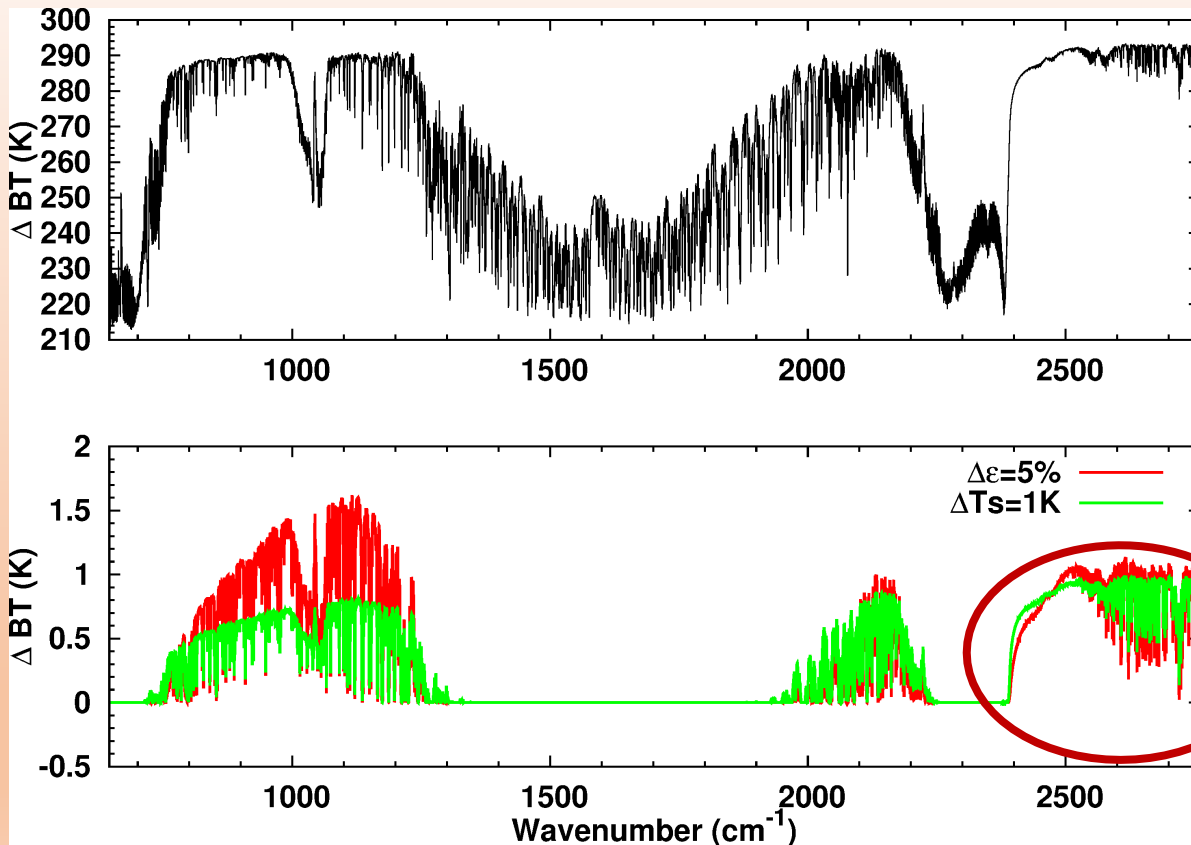


Variation of emissivity
larger than 20% in few km



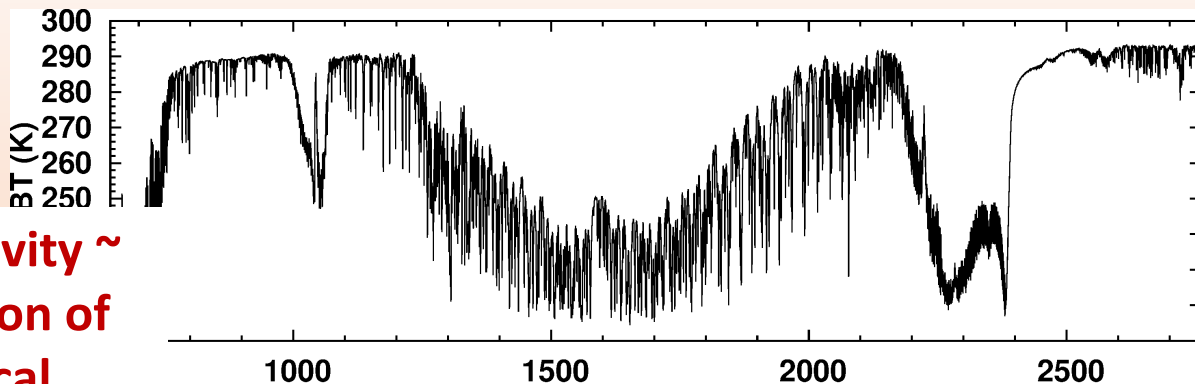
5% of emissivity (~ 0.05) \Rightarrow
1.5K in BT

Impact of surface parameters on BT spectrum

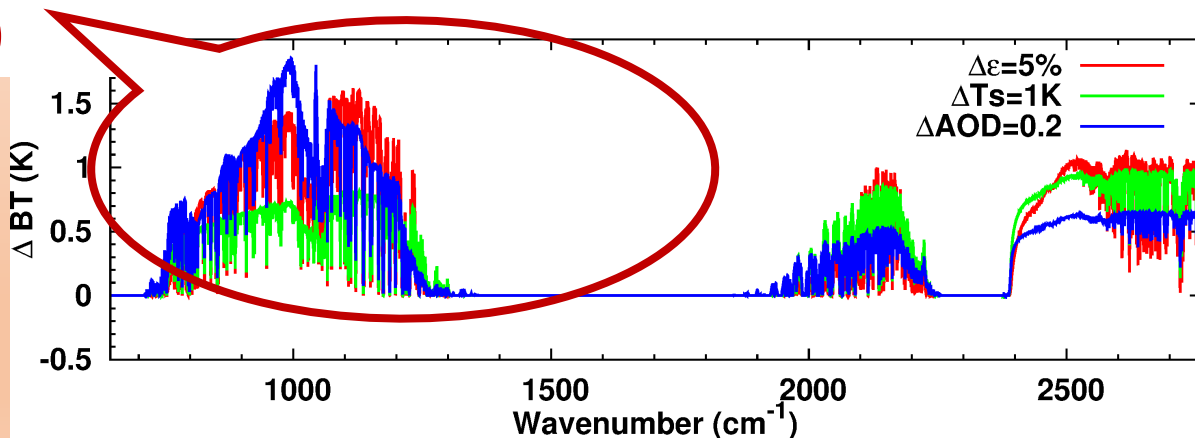


5% of emissivity
~ 1K of T_{surf}

Impact of surface parameters on BT spectrum



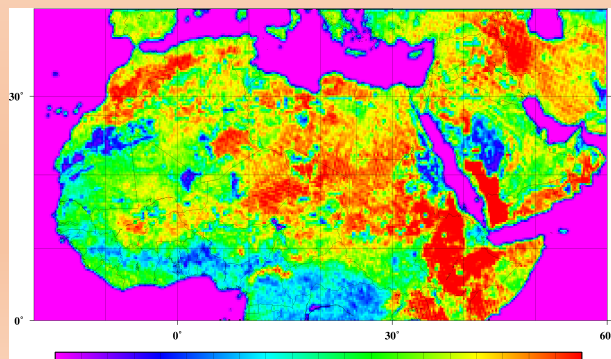
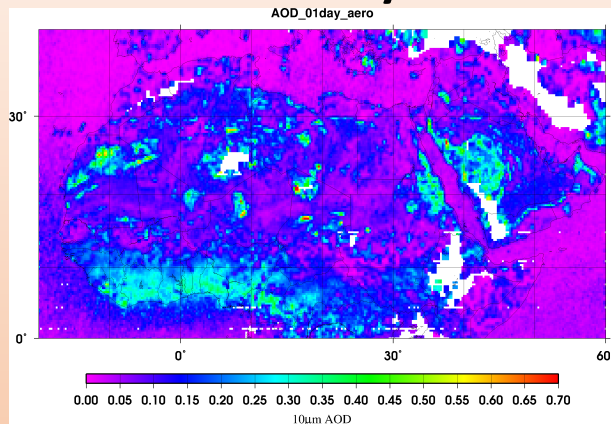
5% of emissivity ~
small variation of
aerosol optical
depth (AOD)



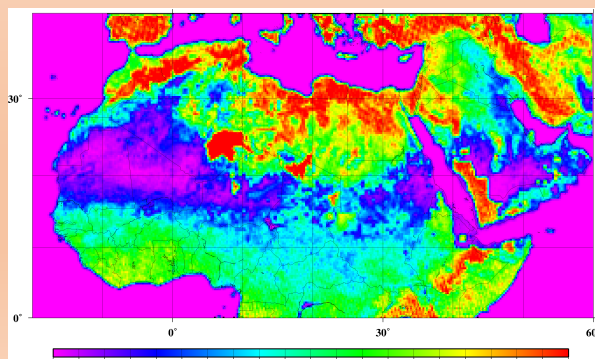
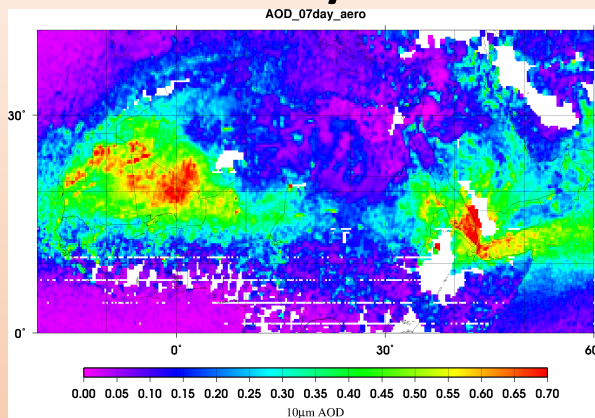
=> Estimating surface parameters requires precise knowledge of aerosol properties and/or a severe filtering.

Possibility of dust filtering? Statistics on the 2007-2017 IASI period

January



July

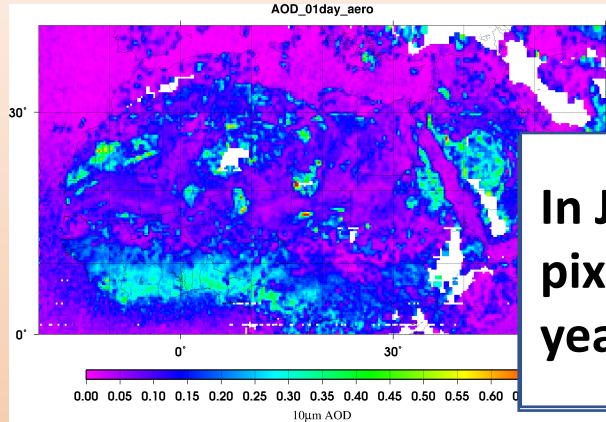


**Mean of 10
µm Dust AOD**

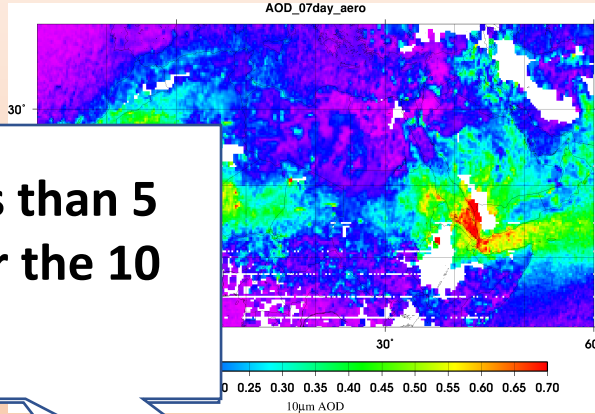
**Number of
observation in
each 0.25°
pixel grid with
AOD<0.15**

Possibility of dust filtering? Statistics on the 2007-2017 IASI period

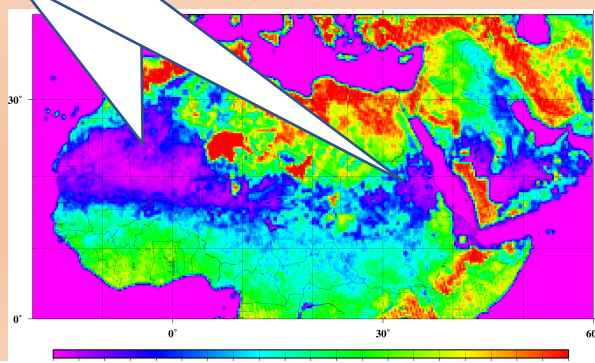
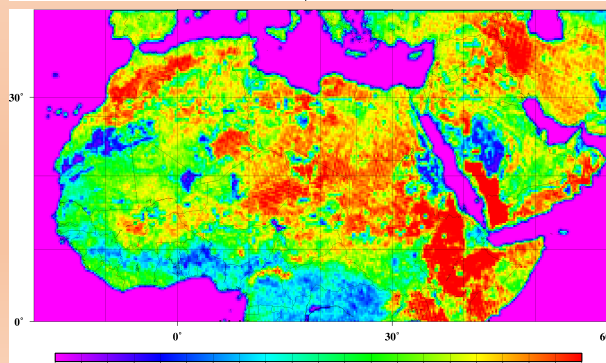
January



July



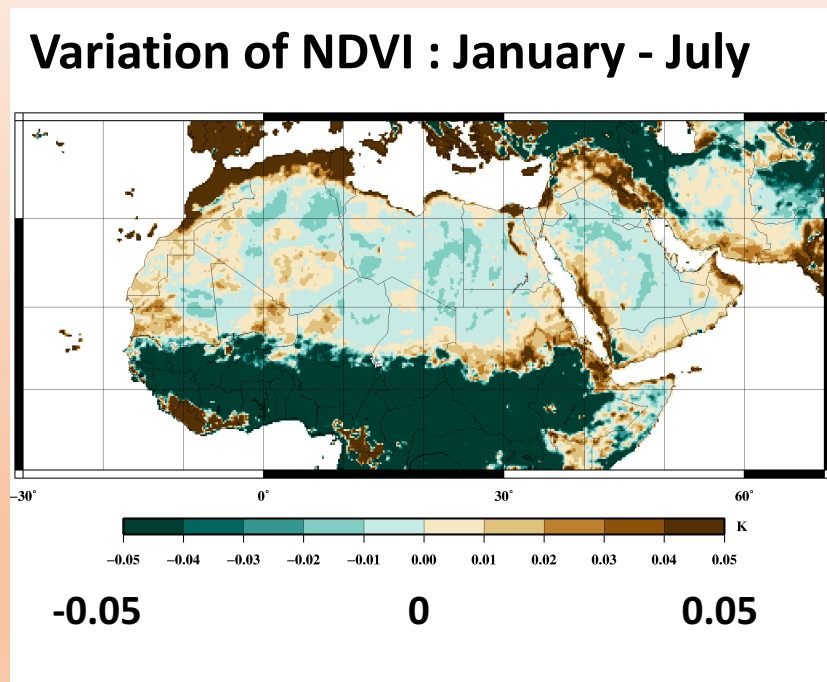
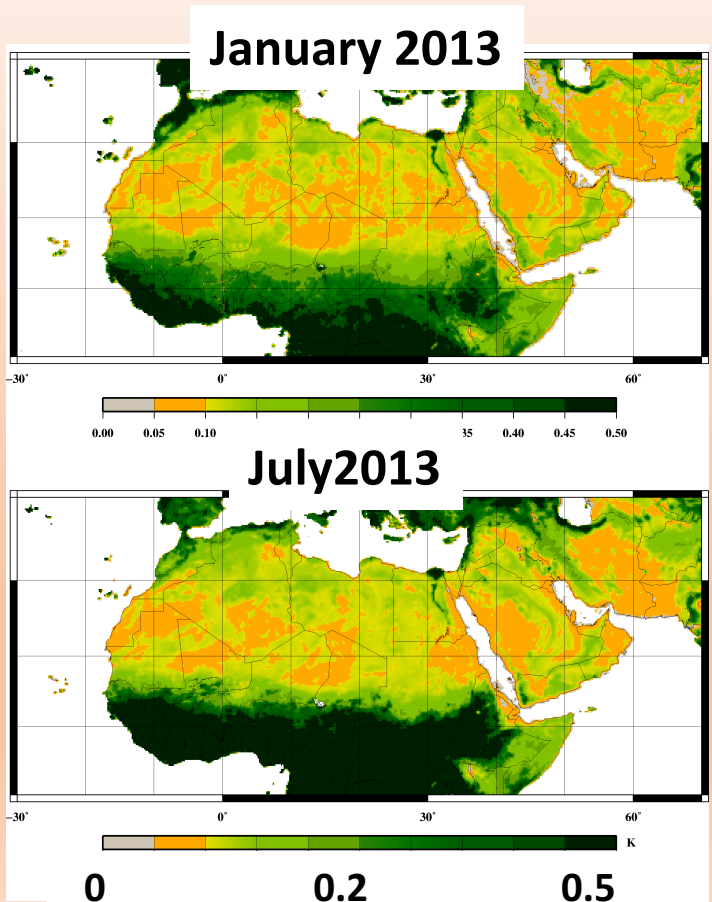
In July less than 5 pixels over the 10 years!



Mean of 10 µm Dust AOD

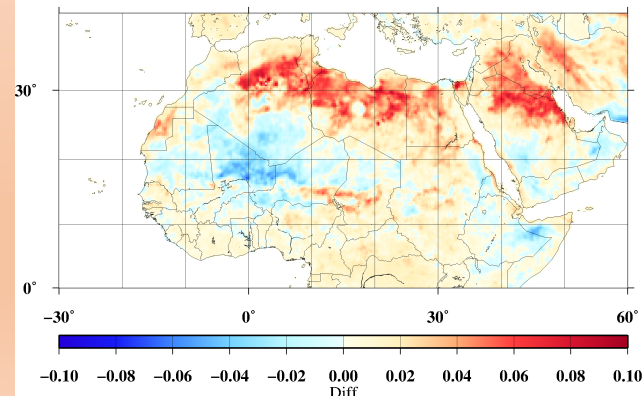
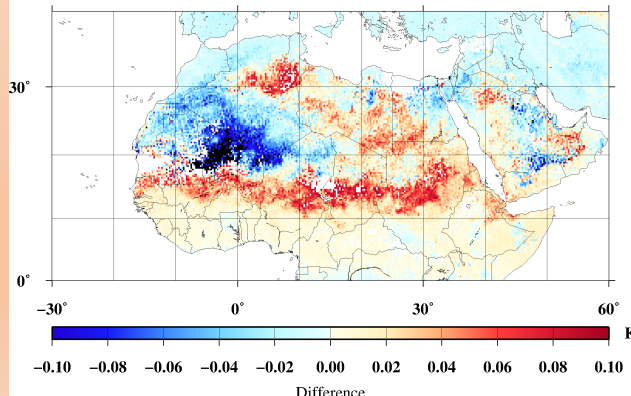
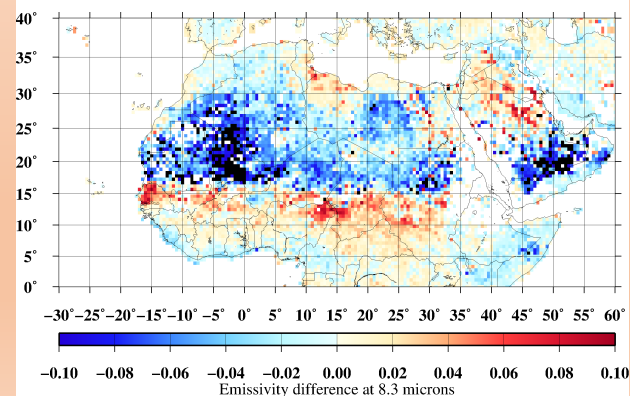
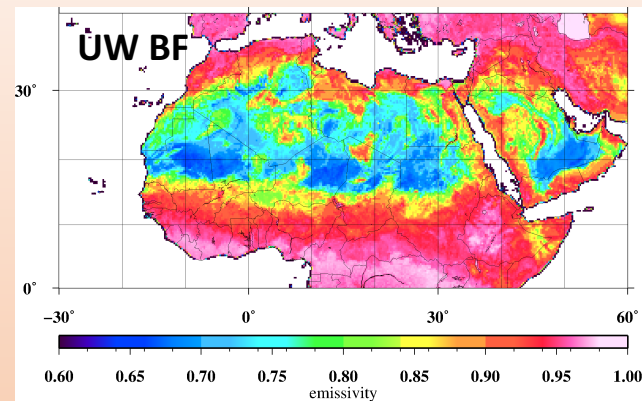
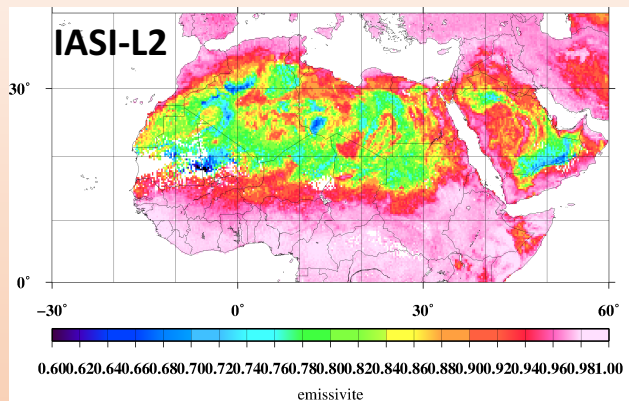
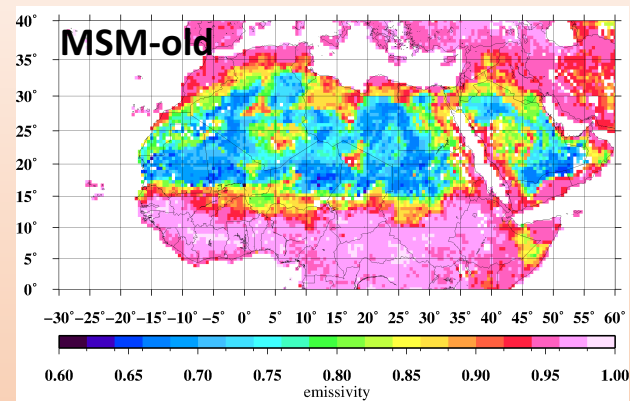
Number of observation in each 0.25° pixel grid with AOD < 0.15

Expected seasonal variation of emissivity: NDVI seasonal variation as proxy



Impact of dust on Emissivity databases

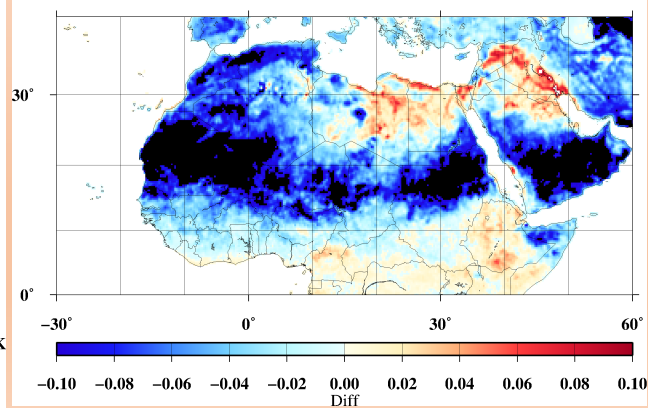
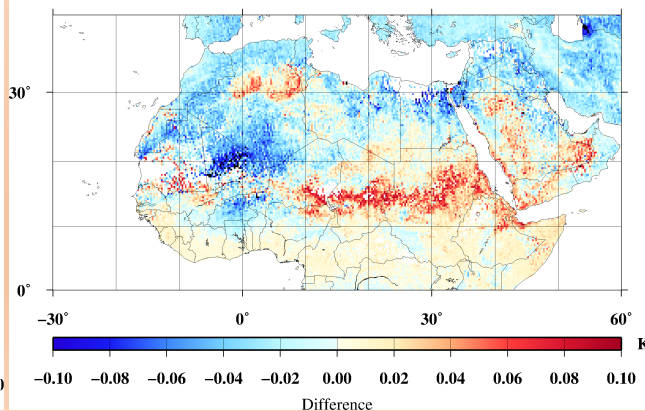
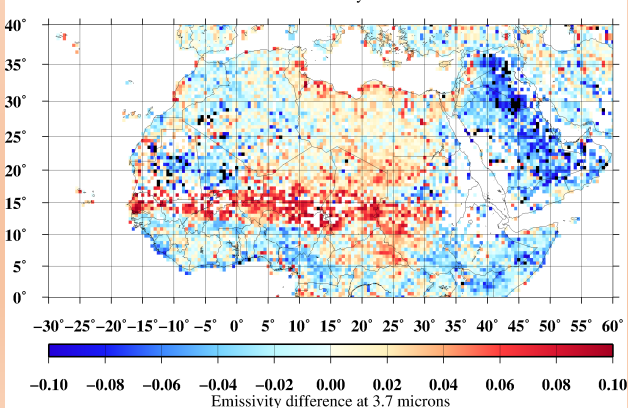
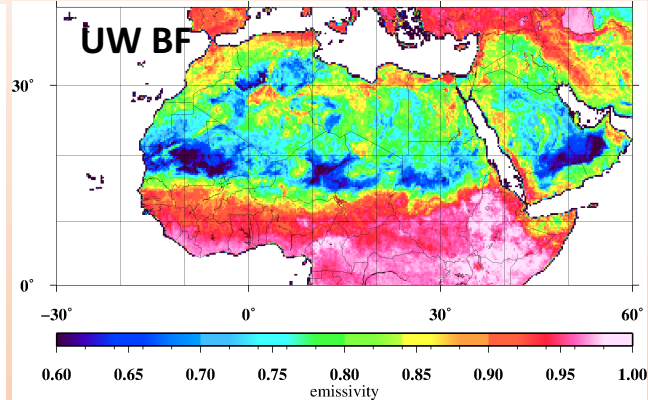
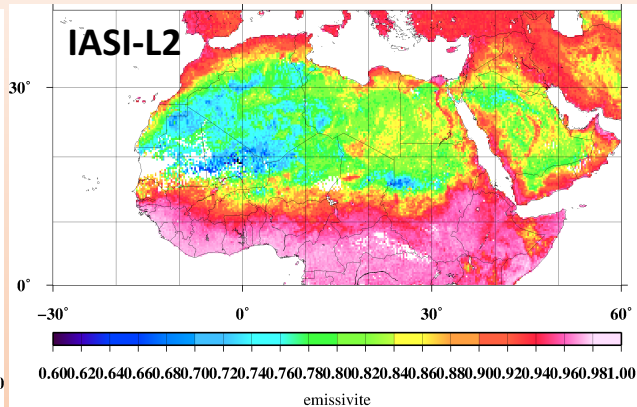
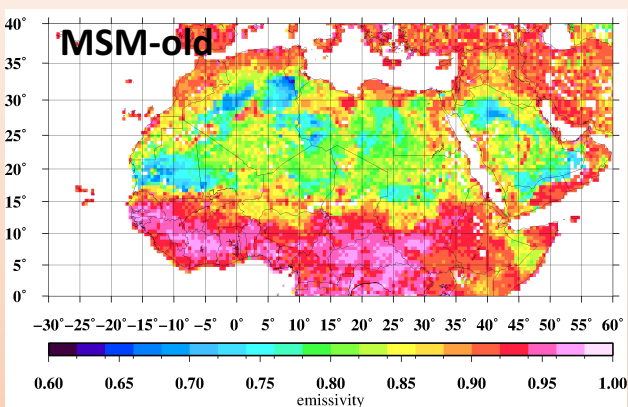
Surface emissivity at 8.3 μm (January 2013)



Surface emissivity difference at 8.3 μm (January –July 2013)

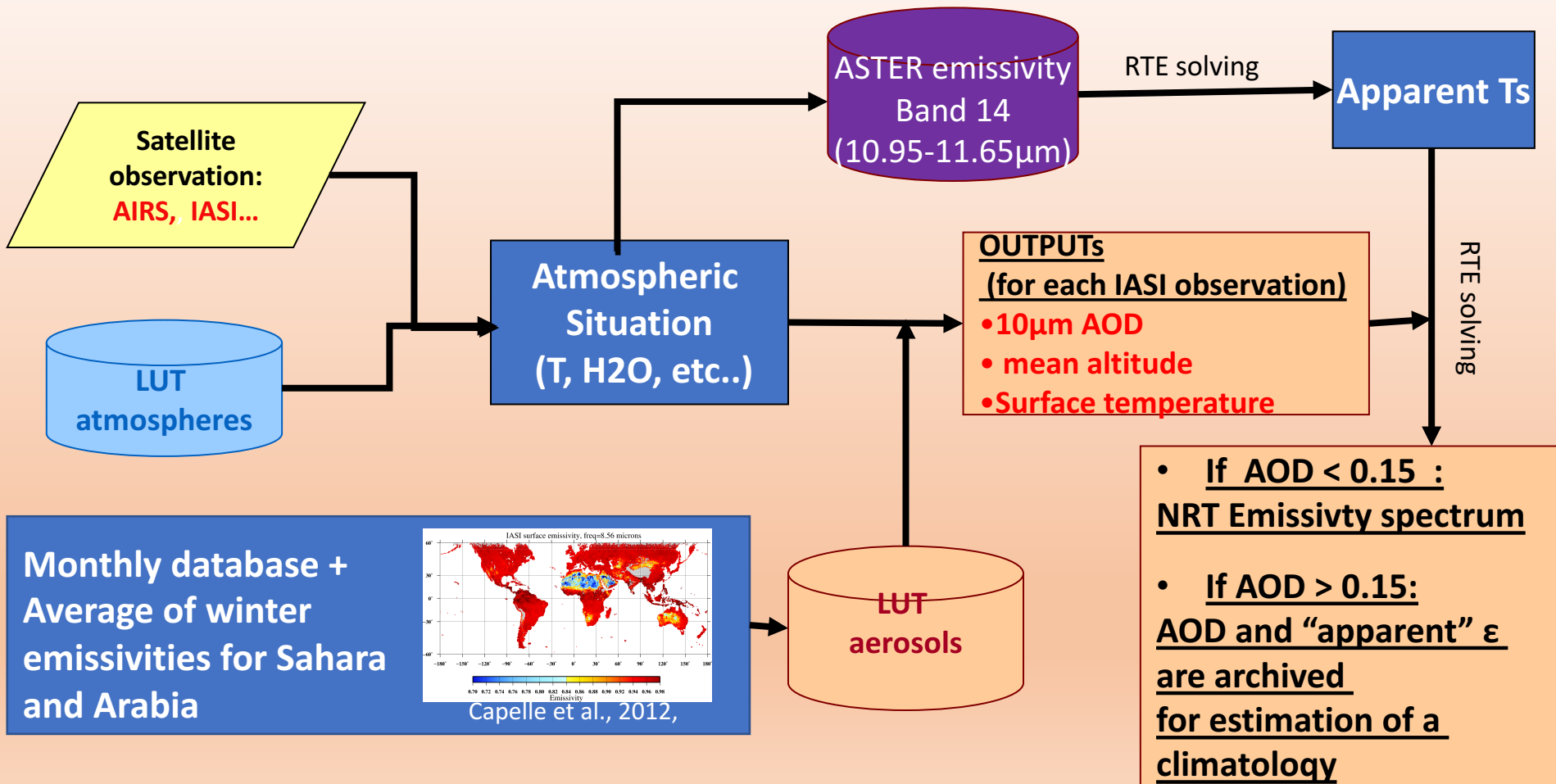
Impact of dust on Emissivity databases

Surface emissivity at 3.7 μm (January 2013)



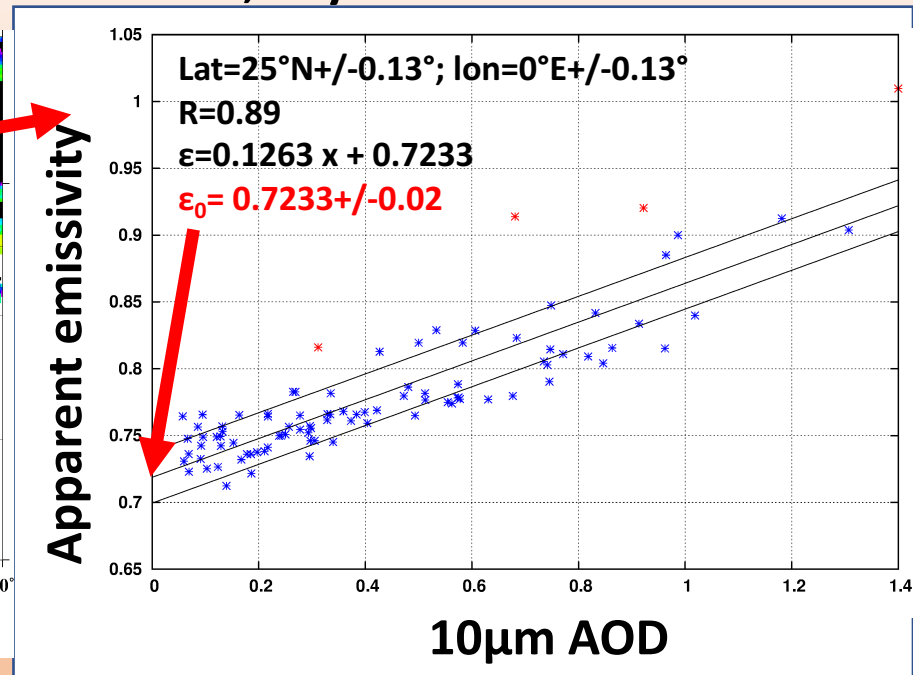
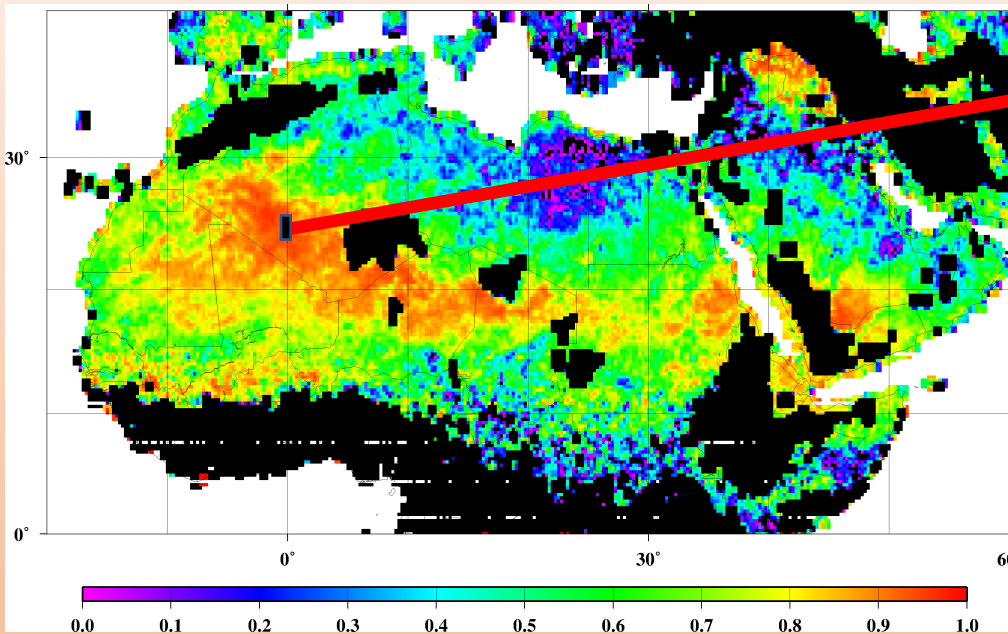
Surface emissivity difference at 3.7 μm (January –July 2013)

Joint surface/aerosol inversion scheme : an iterative process



Dust + emissivity coupling: AOD/emissivity correlation

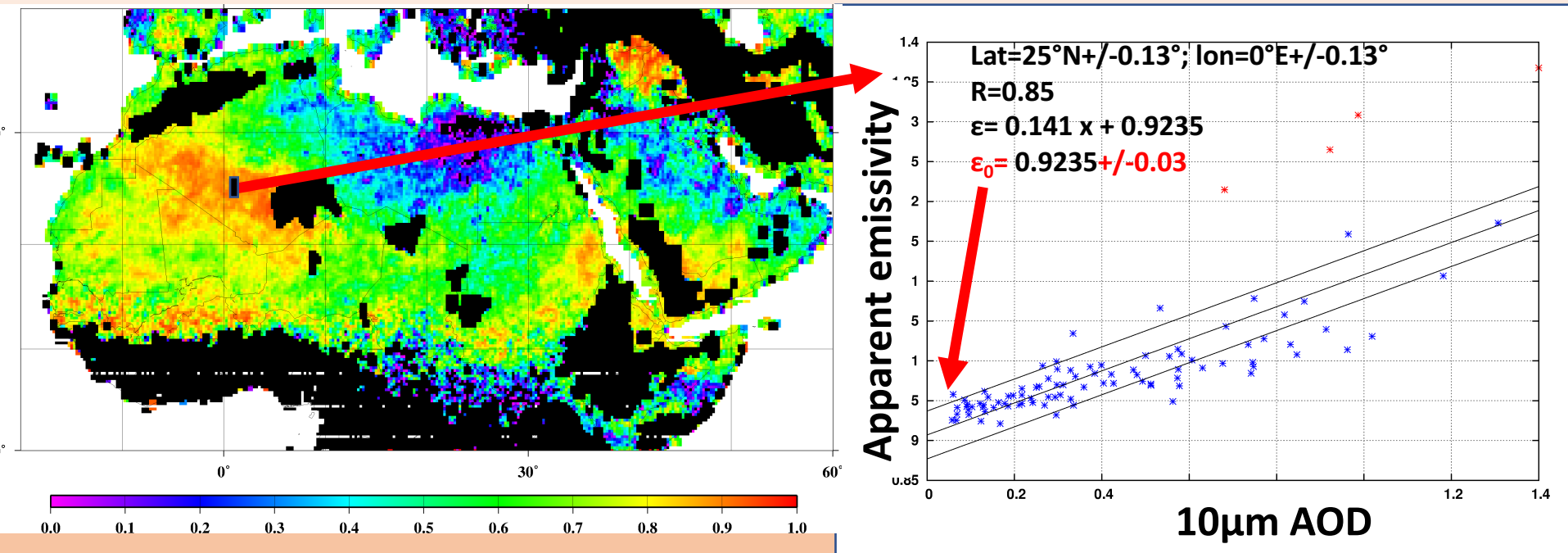
AOD/emissivity at $8.3\mu\text{m}$ correlation: July 2007-2017; daytime



- \Rightarrow AOD/emissivity correlation computed for each pixel of a 0.25° grid and for a selection of window channels
- \Rightarrow ϵ obtained by regression

Dust + emissivity coupling: AOD/emissivity correlation

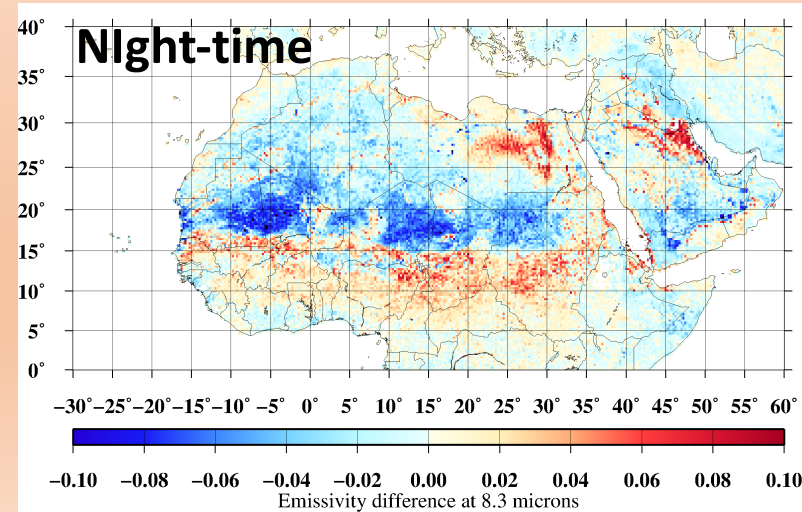
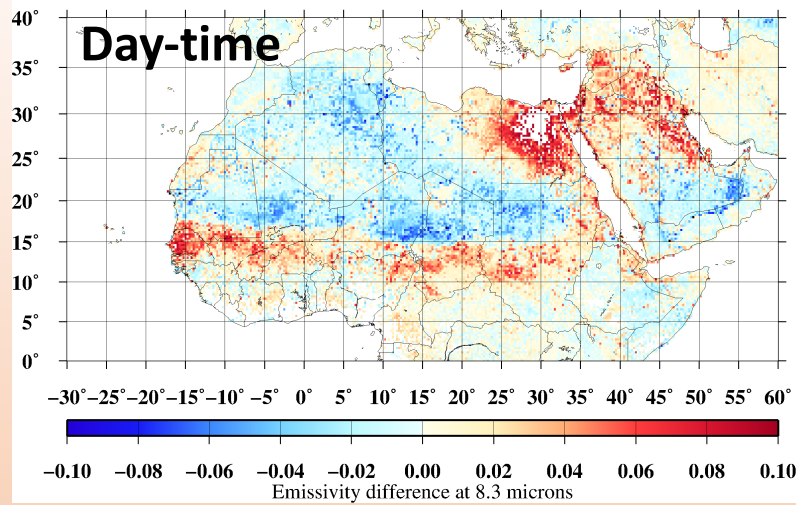
AOD/emissivity at 4.63 μ m correlation: July 2007-2017; daytime



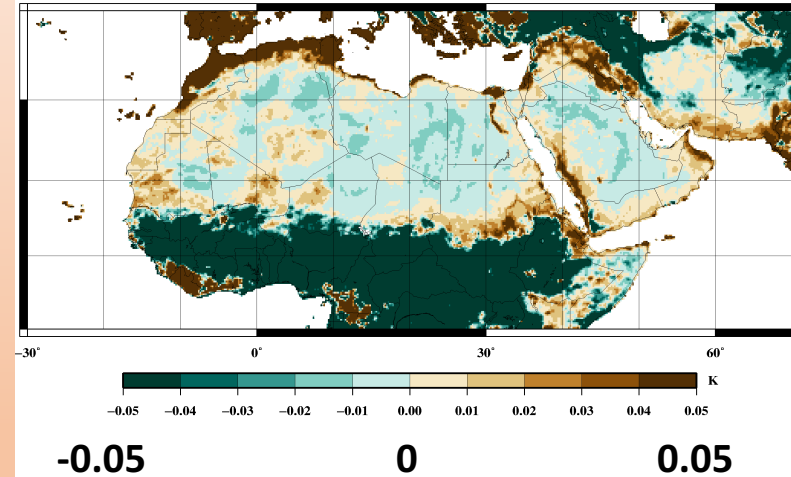
⇒ AOD/emissivity correlation computed for each pixel of a 0.25° grid and for a selection of window channels

⇒ ϵ obtained by regression

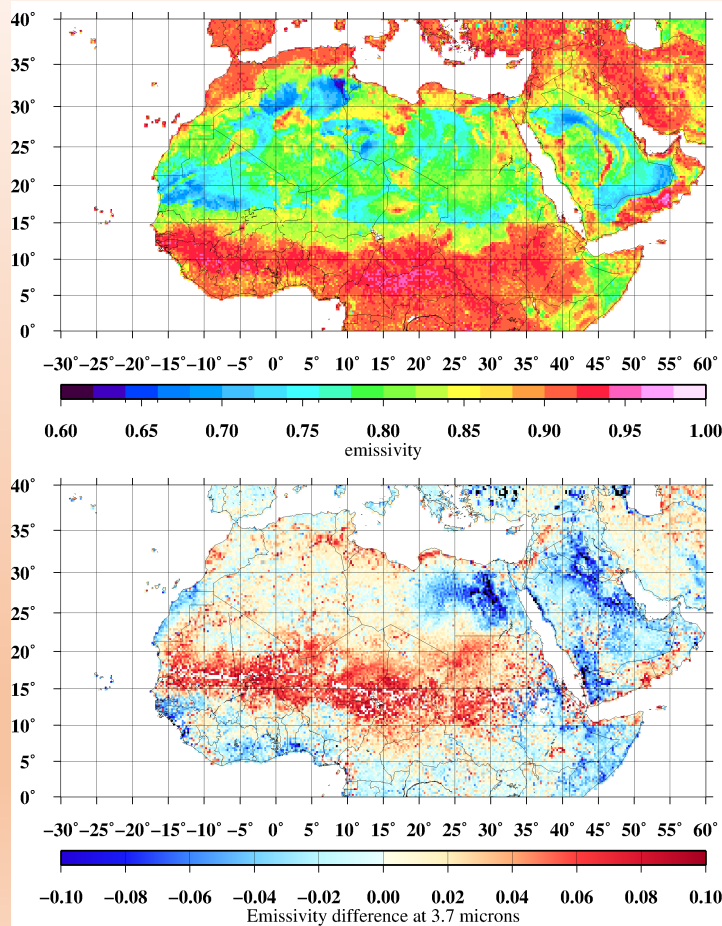
Validation : Seasonal variation at 8.3 μ m



Variation of NDVI : January - July

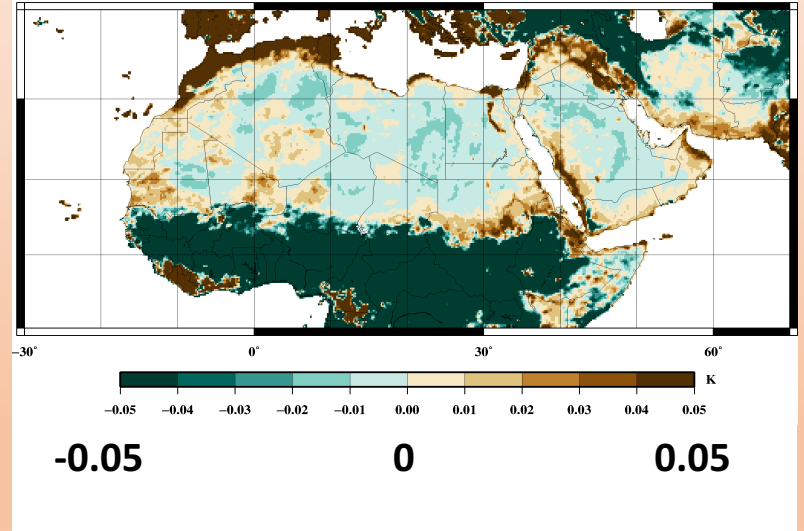


Night-time



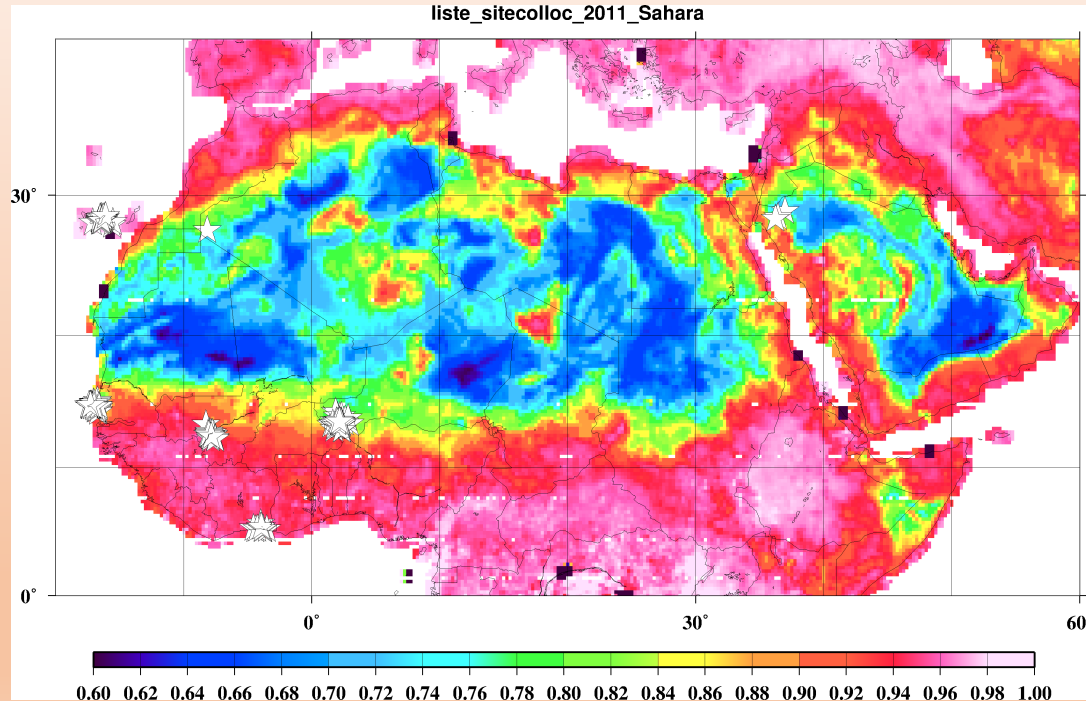
Validation : Seasonal variation at 4.63 μ m

Variation of NDVI : January - July



Validation :

use of 4A/OP + ARSA radiosounding database



Collocation of IASI with the
radio-sounding profile
database **ARSA**

Year: 2011

Night-time conditions

~300 profiles

**Surface temperature: MSM or
L2-IASI**

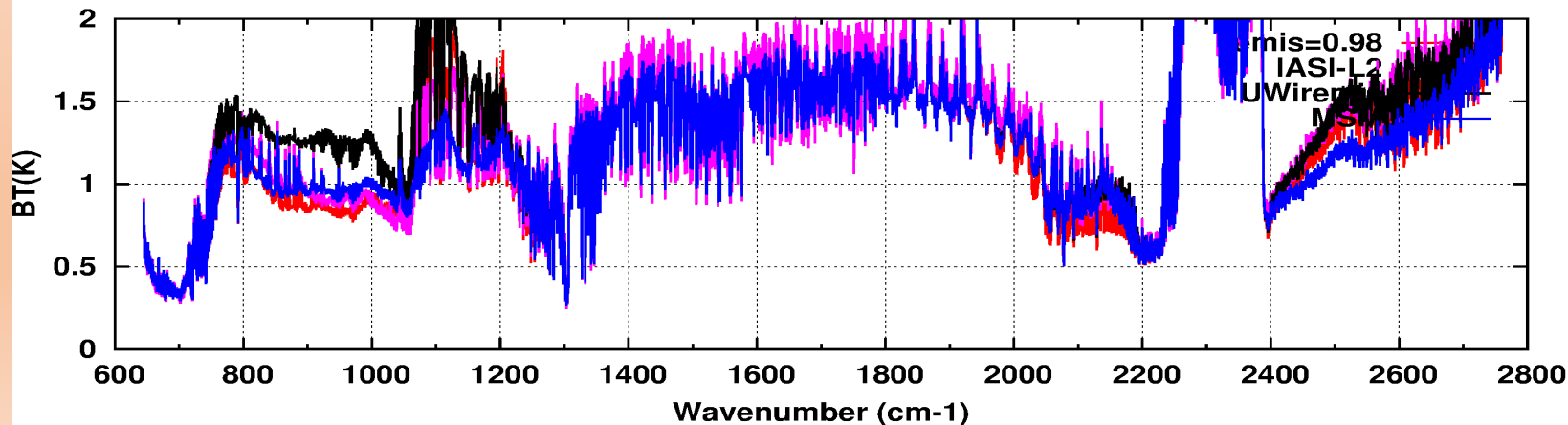
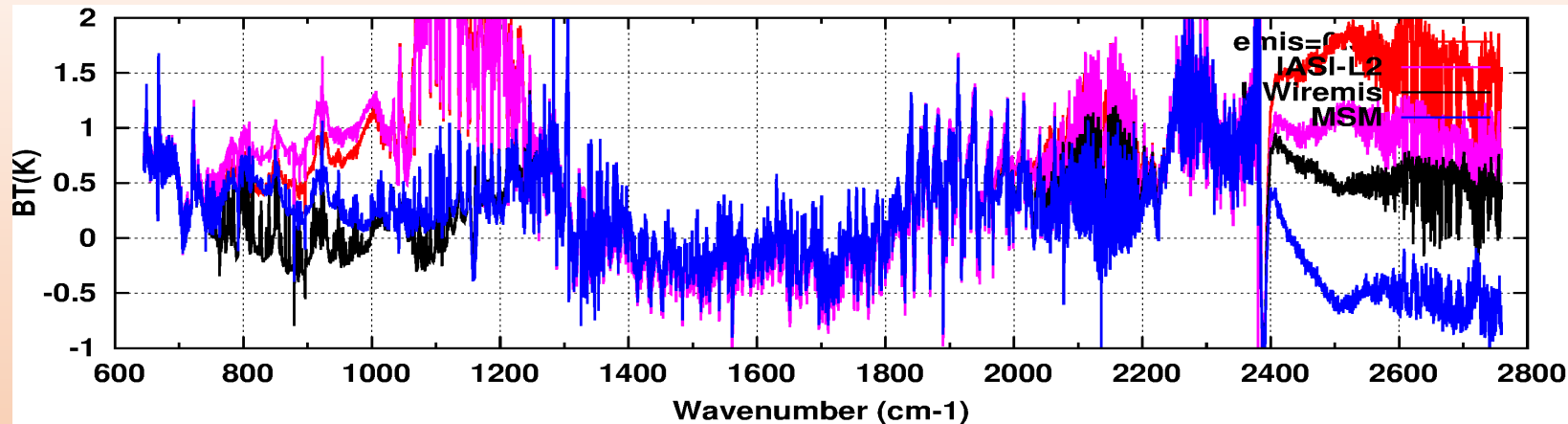
Validation : use of 4A/OP-SPARTE

0.98

IASI-L2

UW HSR

MSM



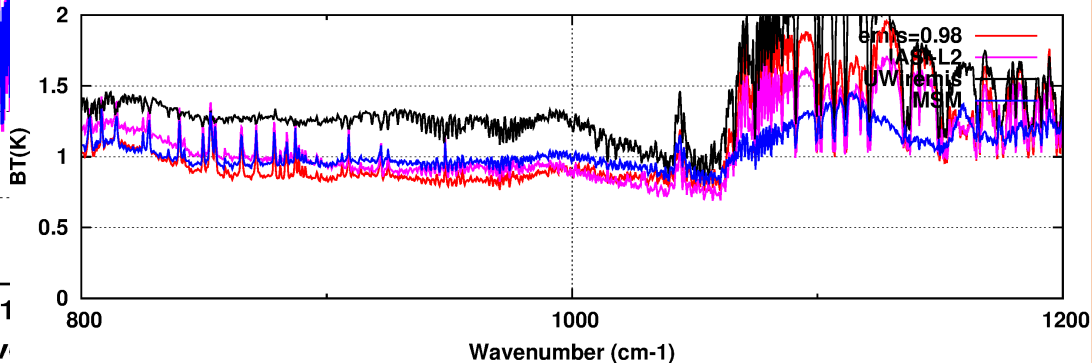
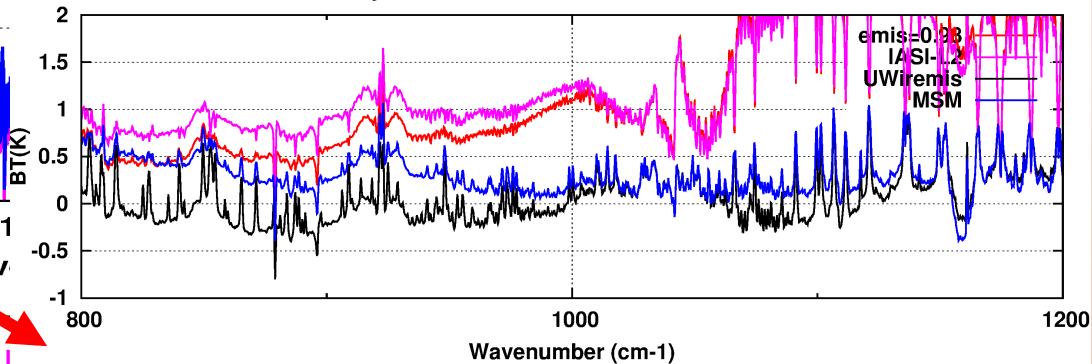
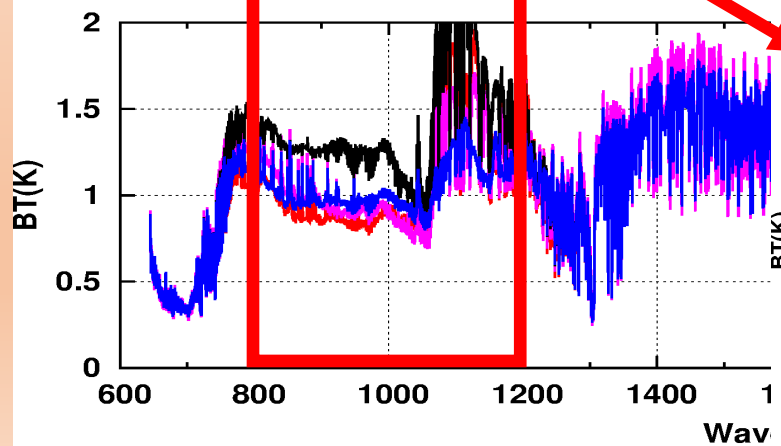
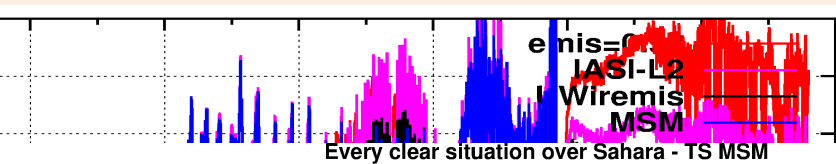
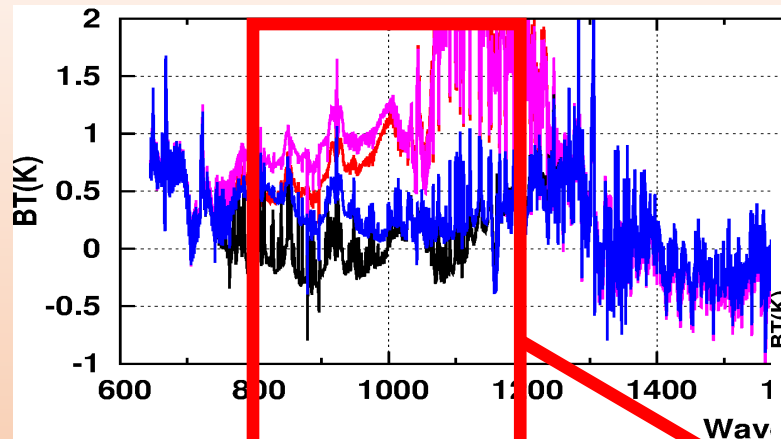
Validation : use of 4A/OP-SPARTE

0.98

IASI-L2

UW HSR

MSM



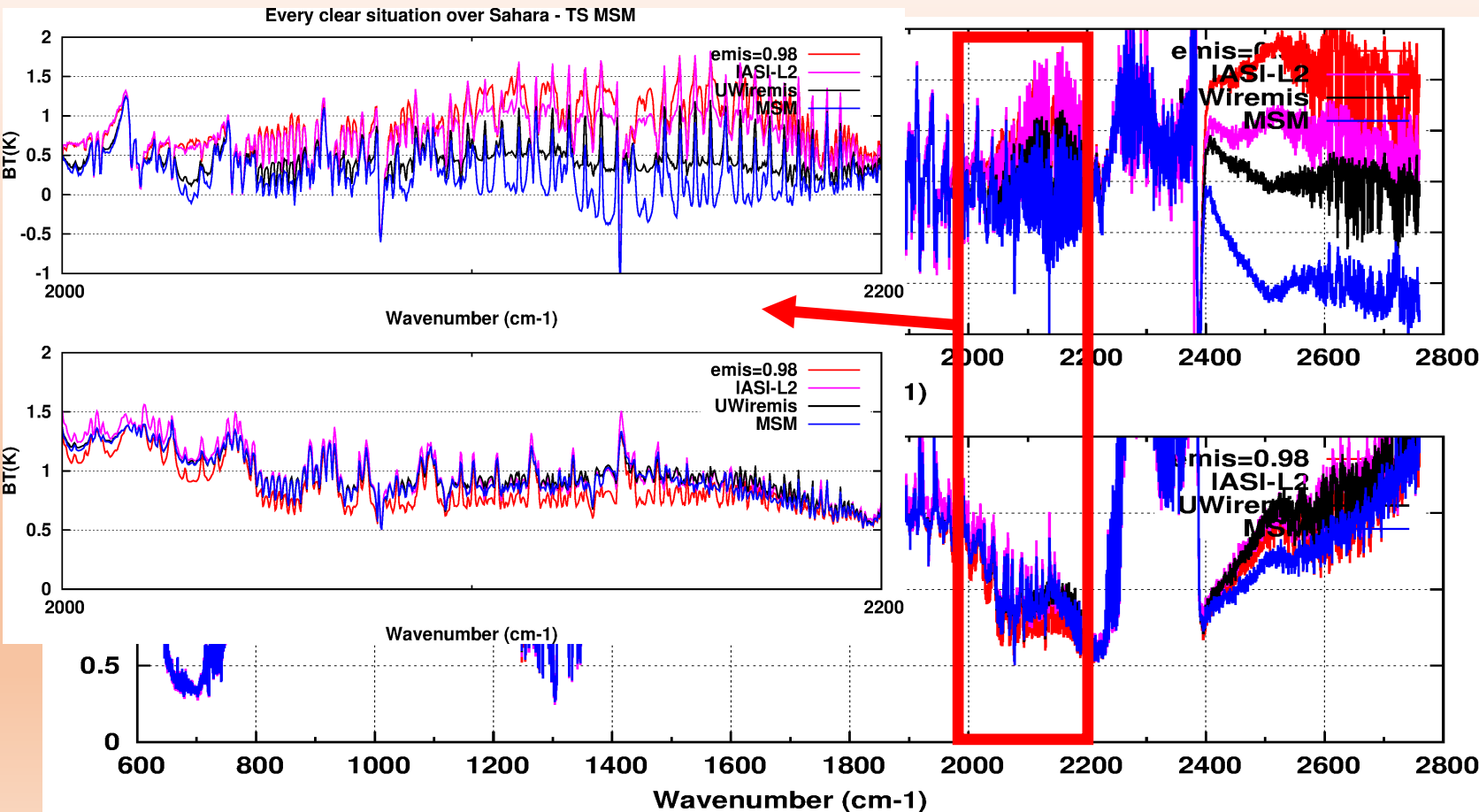
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IASI-L2

UW HSR

MSM



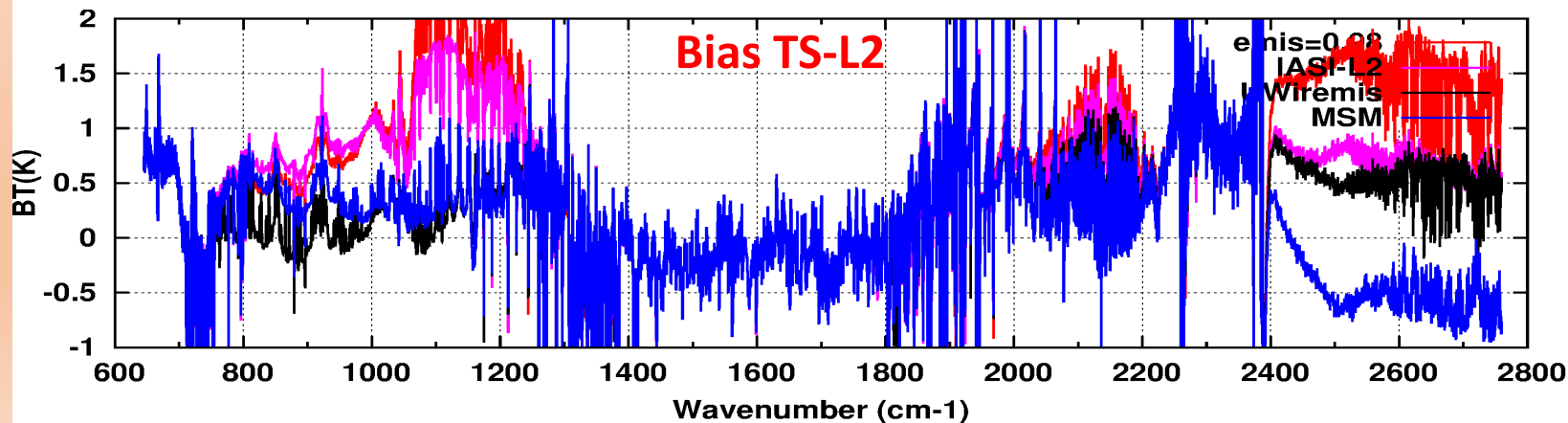
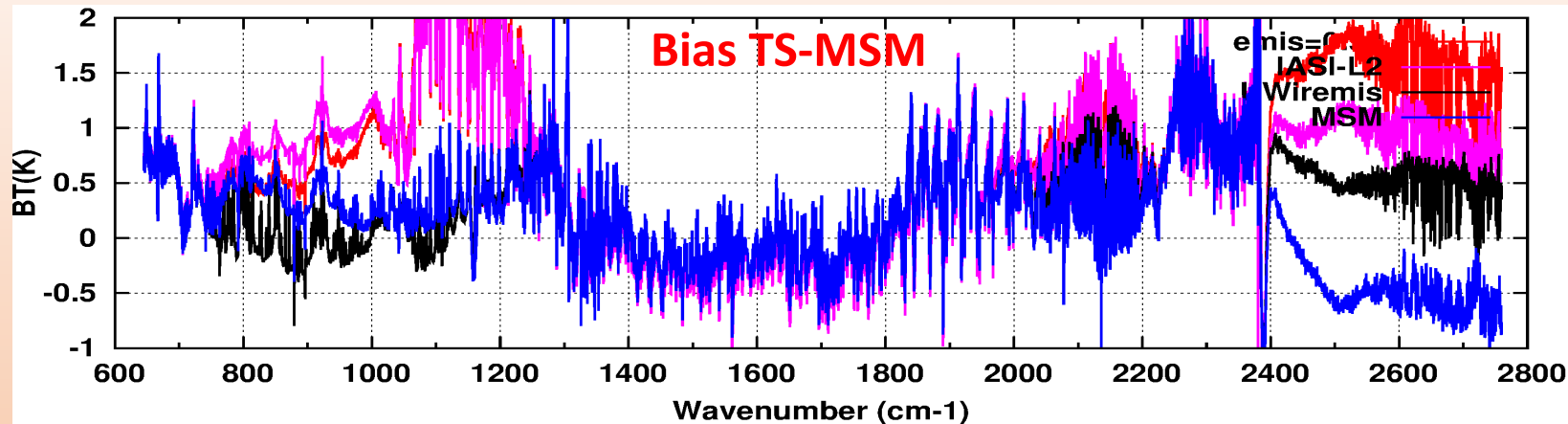
Validation : use of 4A/OP-SPARTE

0.98

IASI-L2

UW HSR

MSM



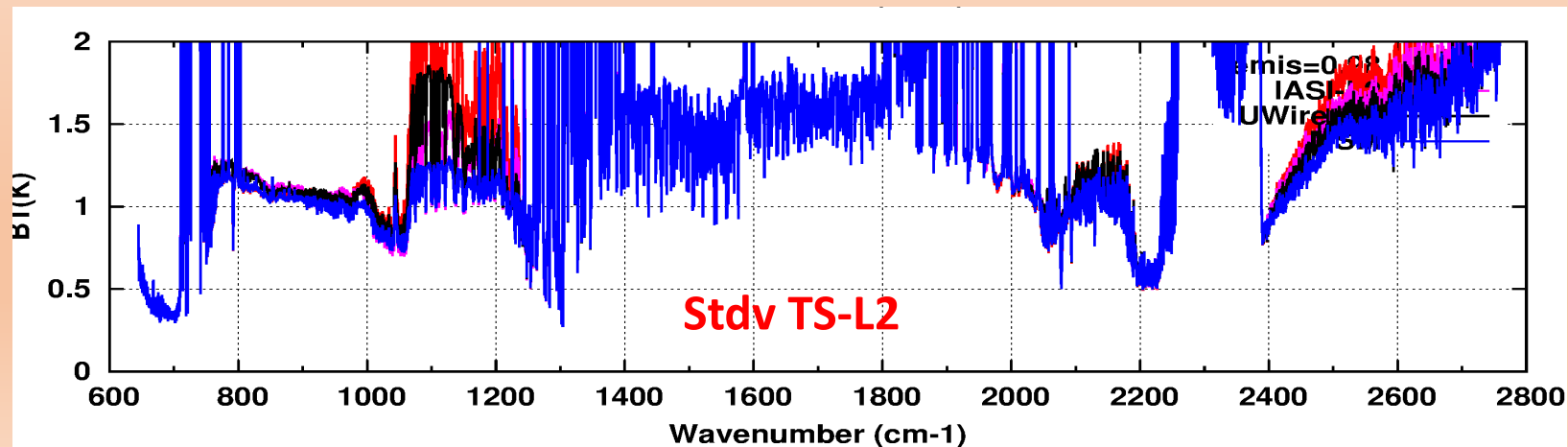
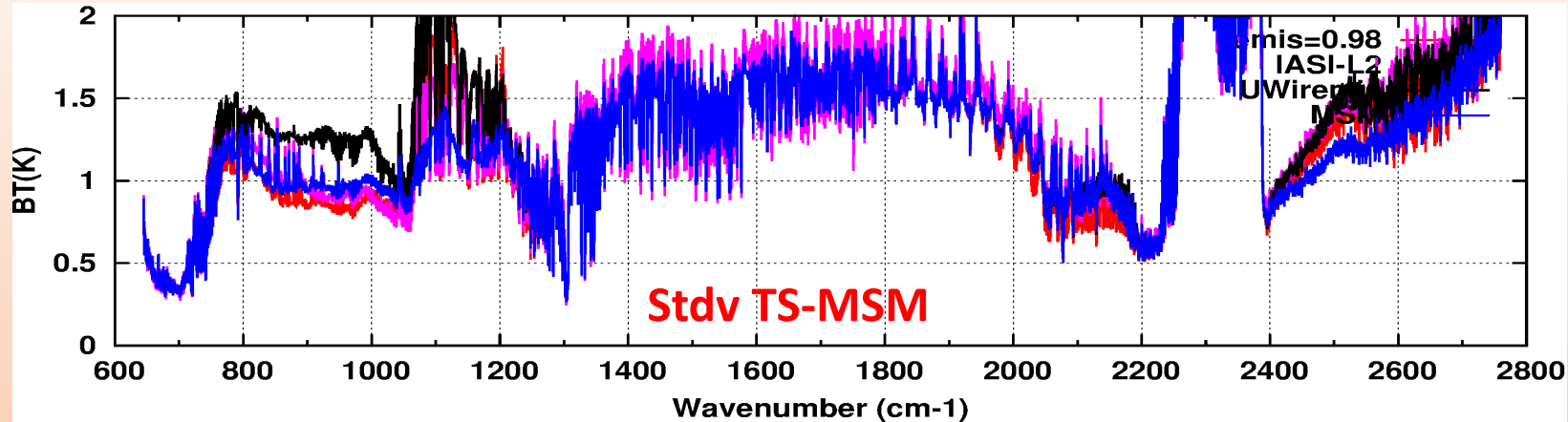
Validation : use of 4A/OP-SPARTE

0.98

IASI-L2

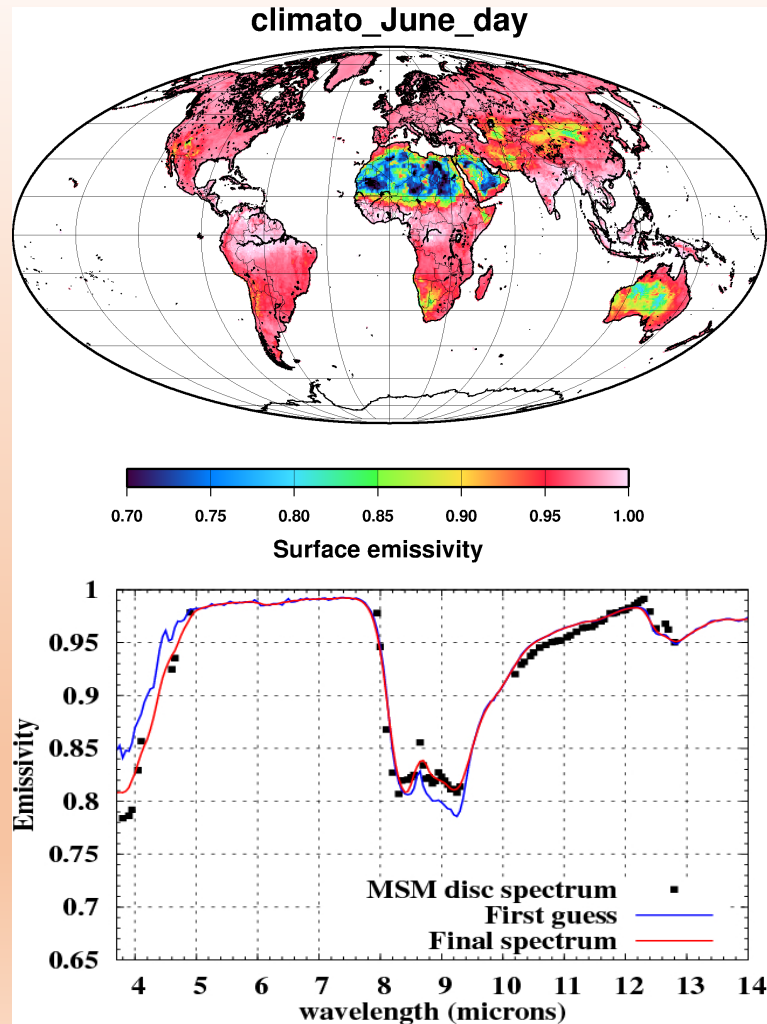
UW HSR

MSM



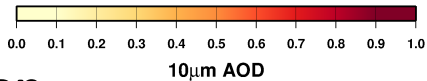
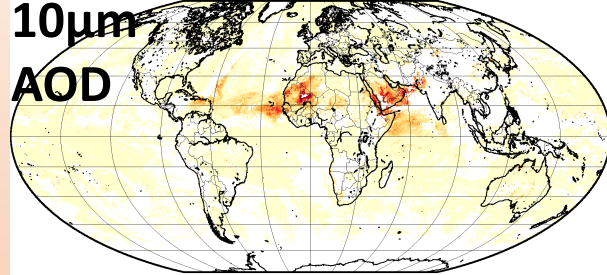
Conclusion

- New IR surface emissivity database from IASI
- Higher spatial resolution than previous ($0.5^\circ \Rightarrow 0.25^\circ$)
- High spectral resolution ($0.05\mu\text{m}$ from 3.7 to $14\mu\text{m}$)
- Clean from dust contamination
- Day and night separate
- Global
- Can be delivered in NRT (D-1) where $\text{AOD} < 0.15$
- Monthly \Rightarrow on-going

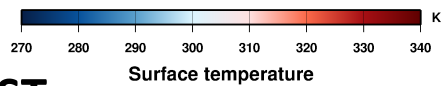
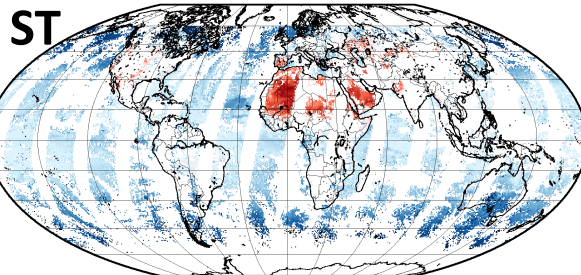


16 June 2013 –Day-time

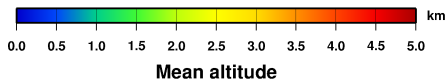
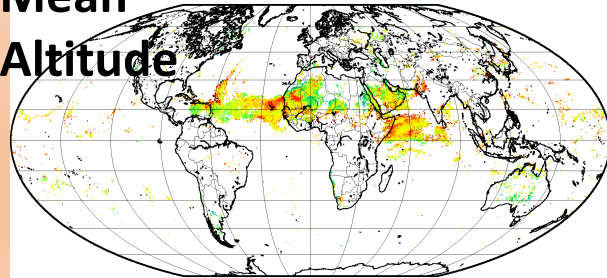
IASIA+IASIB



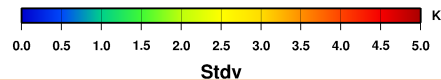
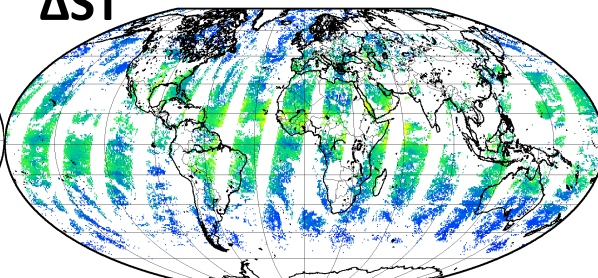
IASIA



Mean Altitude



ΔST

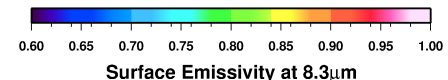
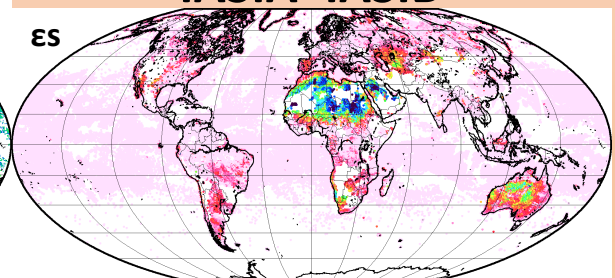


Global and daily restitution
in NRT (D-1) of :

- 10µm dust AOD
- Mean altitude
- Surface temperature
- Emissivity spectrum (if AOD<0.15)

Using **IASI/METOP-A** and **IASI-METOP-B**

IASIA+IASIB

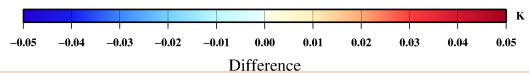
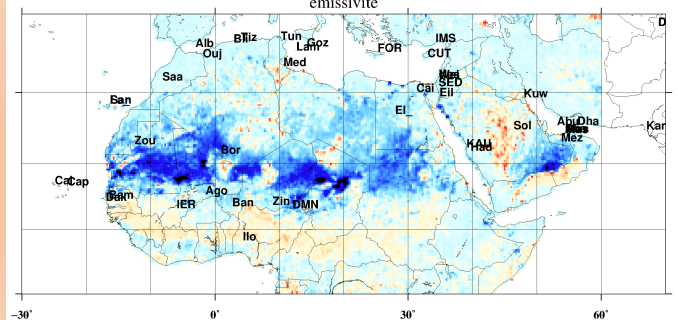
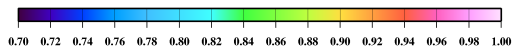
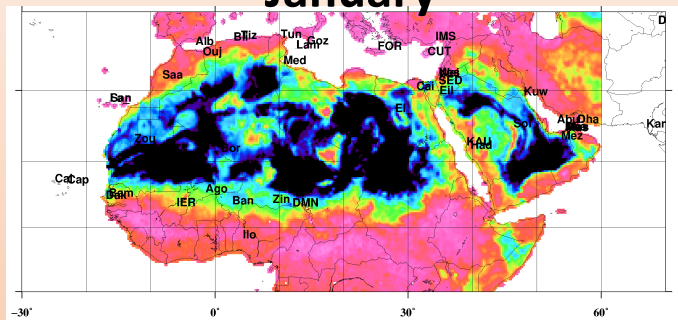


Conclusion

- Day and night variation analysis => on going

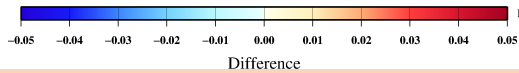
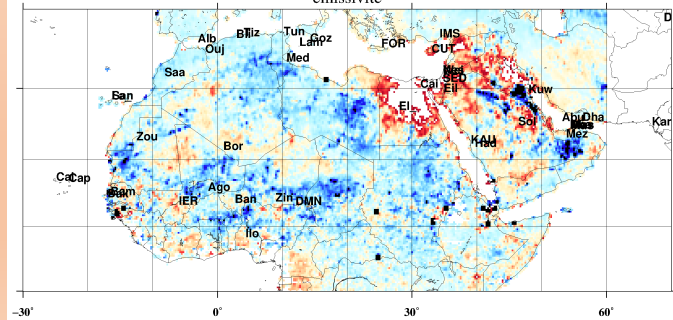
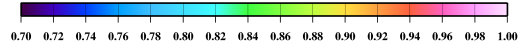
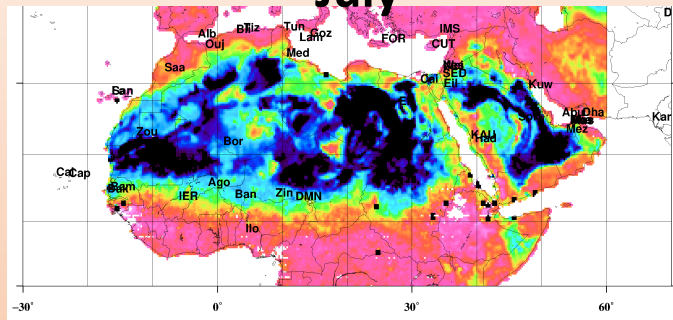
Night – Day at $8.3\mu\text{m}$

January



Night – Day at $8.3\mu\text{m}$

July



Infrared RTE

(lambertian surface, clear sky, night)

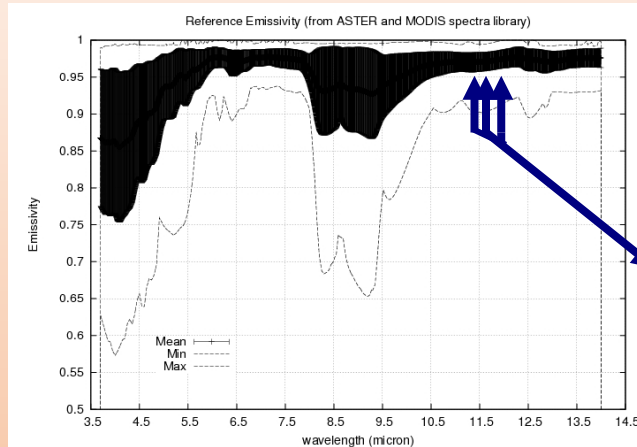
$$I(\lambda, \theta) = \underbrace{\varepsilon_s(\lambda) \tau_s(\lambda, \theta) B(\lambda, T_s)}_{\text{Surface Emission}} + \underbrace{\int_{\tau_s(\lambda, \theta)}^1 B[\lambda, T] d\tau(\lambda, \theta)}_{\text{Upwelling Atmosphere Emission}} + \underbrace{(1 - \varepsilon_s(\lambda)) \tau_s(\lambda, \theta) \int_{\tau_s(\lambda, \theta)}^1 B[\lambda, T] d\tau'(\lambda, \theta)}_{\text{Reflected Downwelling Atmosphere Emission for a Lambertian surface}}$$

$$\tau'(\lambda, \theta) \tau(\lambda, \theta) = \tau_s(\lambda, \theta)$$

$$\varepsilon_s(\lambda) = \frac{I(\lambda, \theta) - \int_{\tau_s(\lambda, \theta)}^1 B[\lambda, T] d\tau(\lambda, \theta) - \tau_s(\lambda, \theta) \int_{\tau_s(\lambda, \theta)}^1 B[\lambda, T] d\tau'(\lambda, \theta)}{\tau_s(\lambda, \theta) \left\{ B(\lambda, T_s) - \int_{\tau_s(\lambda, \theta)}^1 B[\lambda, T] d\tau'(\lambda, \theta) \right\}}$$

Estimate the surface skin temperature

- 3 channels selected for their good transmittance ($\tau \geq 0.6$) and a small variability of the emissivity ($\sigma \sim 0.01$):



IASI channel	Mean ε	ε std dev
754 (12.0 μm)	0.975	0.012
867 (11.6 μm)	0.971	0.013
921 (11.4 μm)	0.970	0.013

- $\varepsilon \sim 0.97$ is no more an unknown for these channels and the skin temperature (T_s) remains the only unknown of the RTE.

$$T_s = B^{-1} \left(\frac{I_{sat}(\lambda_0, \theta) - \int_0^1 B[\lambda_0, T(\tau(\lambda_0, \theta))] d\tau - (1 - \varepsilon_s(\lambda_0)) \tau_s(\lambda_0, \theta) \int_0^1 B[\lambda_0, T(\tau'(\lambda_0, \theta))] d\tau'}{\varepsilon_s(\lambda_0) \tau_s(\lambda_0, \theta)} \right)$$

- Final surface temperature obtained by averaging the results from the 3 channels.

Determination of the continuous emissivity spectrum

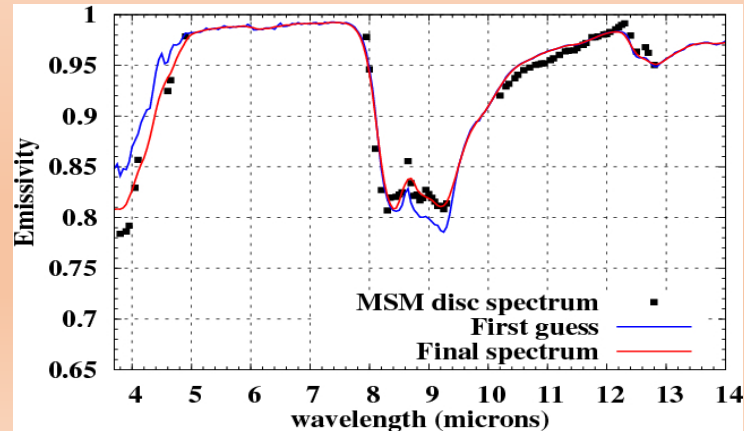
→ ϵ calculated for **101 “hinge points”**
→→→ **Discontinuous spectrum with a demonstrated accuracy accuracy better than 1.5% at 12 μm and $\sim 4.5\%$ at 4 μm .**

**Proximity recognition +
shape fit procedure**

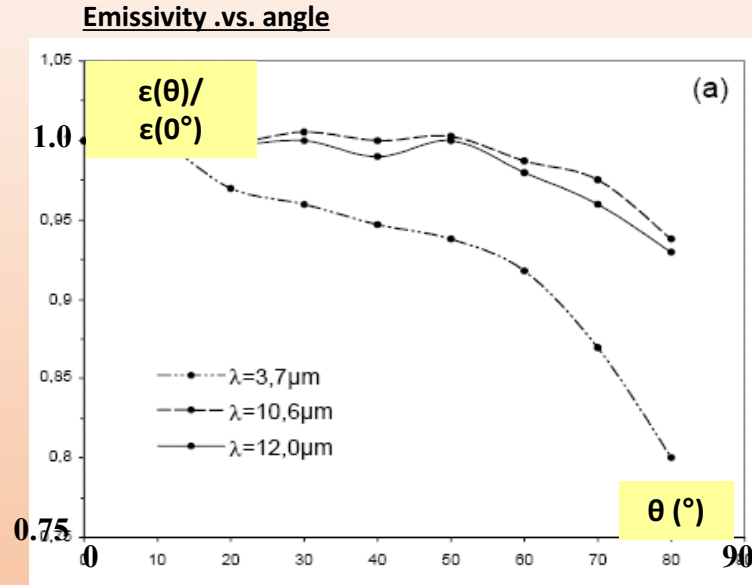
Emissivity continuous spectrum at 0.05 μm resolution between 3.7 and 14.0 μm

MSM emissivity database:

- 165 high spectral resolution emissivity laboratory measurements of different samples of typical Earth surfaces are selected from MODIS/UCSB and ASTER/JPL emissivity libraries.
- They are linearly interpolated at 0.05 μm resolution between 3.7 and 14.0 μm .



Perspective: Introduction of BRDF in transfer modeling



Labeled J. and M.P. Stoll (1991), *International Journal of Remote Sensing*,

Angular variation impact on TIR emissivity:

- negligible impact in longwave ($\lambda > 4\mu\text{m}$)

- Impact $< 5\%$ for $\theta < 40^\circ$

- impact even less given the spatial resolution of the observation

=> Introduction of BRDF may improve results for $\lambda > 4\mu\text{m}$ and/or $\theta > 40^\circ$